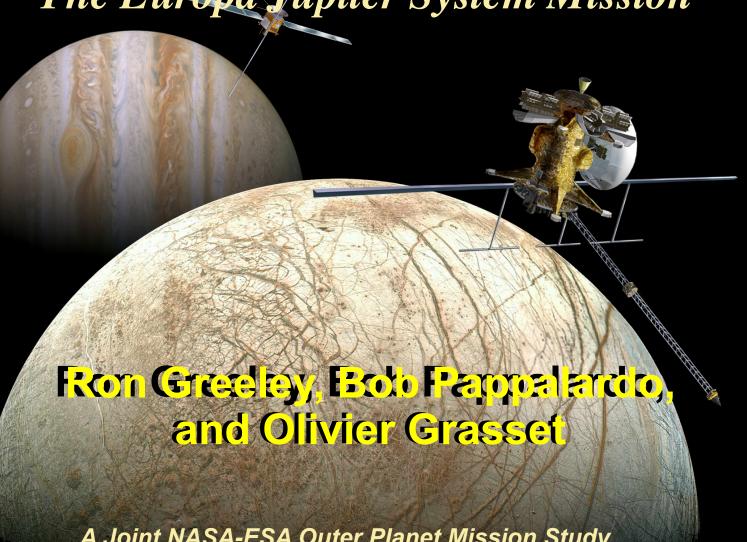




Exploring the Habitability of Icy Worlds: The Europa Jupiter System Mission



A Joint NASA-ESA Outer Planet Mission Study















OPAG Review

- Strategic Science Plan Status
- Europa Jupiter System Mission
 - Overview
 - Jupiter Europa Orbiter
 - Jupiter Ganymede Orbiter
- OPAG review and input















Strategic Science Plan Europa Jupiter System Mission (EJSM)

Objectives

- Continue direct science involvement with JEO, JGO Projects
- "Retire" weaknesses identified in the NASA and ESA reviews
- Build the case for JGO to be an ESA "L" class mission
- Refine joint aspects of EJSM (JEO, JGO)
- Incorporate the best aspects from the Europa and Titan studies
- Engage broad science community and public support















EJSM Study Science Structure

Leads (8)

Working Groups (22 co-chairs)

Full Science Team (57)

Community (OPAG; EU; etc.)

















Science Definition Team Working Groups SDT Co-chairs plus community members

Working Group 1 (Satellites)

Geophysics

Composition

Ice

Geology

Atmospheres/exospheres

Working Group 2 (Jupiter)

Working Group 3 (Magnetospheres)

Working Group 4 (Jupiter System)

Bruce Bills, Hauke Hussman

Federico Tosi, Tom McCord

Don Blankenship, Olivier Grasset

Ralf Jauman, Jeff Moore

Melissa McGrath, Andrew Coates

Pierre Drossert, Leigh Fletcher, Amy Simon-Miller

Krishan Khurana, Norbert Krupp

Tim Van Hoolst, Melissa McGrath

Transverse/cross-cutting Working Groups

Working Group 5 (Origin and Formation)

Working Group 6 (Astrobiology)

Working Group 7 (Cosmic Connections)

Working Group 8 (Radio Science)

Working Group 9 (Education Public Outreach)

Angioletta Coradini, Bill Moore, Hunter Waite

Kevin Hand, Olga Prieto-Ballesteros

Athena Coustenis, Masaki Fujimoto

Paolo Tortora, Essam Marouf

Athena Coustenis, Ron Greeley, Michel Blanc,

Louise Prockter

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Community Engagement: Special Sessions

- Lunar and Planetary Science Conference (Houston, 23-27 Mar. 2009)

 Two sessions
- European Geoscience Union (Vienna, 19-24 Apr. 2009)

 Four sessions
- EJSM Instrument Workshop (Maryland, 15-17 Jul. 2009)
- European Planetary Science Congress (Potsdam, 13-18 Sep. 2009) Four sessions and two workshops
- Division of Planetary Science (Puerto Rico, 4-9 Oct. 2009)

 Several relevant sessions
- Geological Society of America (Portland, 18-21 Oct. 2009) Cryovolcanism in the Solar System
- American Geophysical Union (San Francisco, 14-18 Dec. 2009) The Galilean Satellites: 400 years of Discovery
- *Galileo 400th* (Padua, 6-9 Jan. 2010)
- EJSM Instrument Workshop (Noordwijk, 18-20 Jan. 2010)

Similar sessions and activities will be held throughout 2010











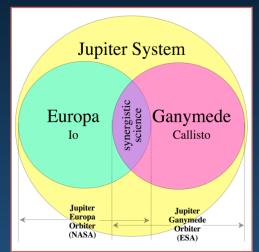




Europa Jupiter Science Mission (EJSM)



- NASA and ESA: Shared mission leadership
- Independently launched and operated orbiters
 - NASA-led Jupiter Europa Orbiter (JEO)
 - ESA-led Jupiter Ganymede Orbiter (JGO)
- Complementary science and payloads
 - JEO concentrates on Europa and lo
 - JGO concentrates on Ganymede and Callisto
 - Synergistic overlap
 - 11-12 instruments each
- Science goals:
 - Icy world habitability
 - Jupiter system processes



Synergistic science: The sum of JEO + JGO is greater than the parts









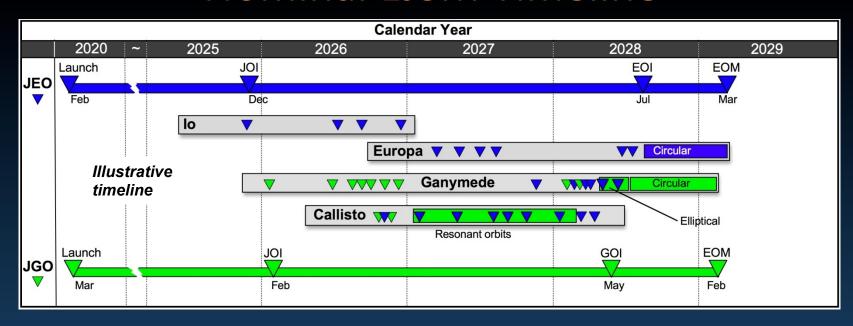








Nominal EJSM Timeline



- Launches: 2020
- Jovian system tour phases: 2–3 years
- Moon orbital phases: 6–12 months
- End of Prime Missions: 2029
- Flexibility if either flight element is delayed or advanced

Coordinated timelines ensure synergistic science













JEO Goal:

Explore Europa to Investigate Its Habitability

Habitabili



Objectives (prioritized):

- Ocean and Interior
- Ice Shell
- Chemistry and Composition
- Geology and Landing Sites
- Jupiter System
 - Satellite surfaces and interiors
 - Satellite atmospheres
 - Plasma and magnetospheres
 - Jupiter atmosphere
 - Rings

Characterizing the archetype of icy world habitability

















JEO Traceability: Europa

Goal		Science Objective	Science Investigation						
	Ocean	Characterize the extent of	A1. Determi	ne the ampl	tude and phase of the gravitational tides.				
		the ocean and its relation	A2. Charact	erize the ma	gnetic environment (including plasma) to determine the induction response				
.		to the deeper interior.			multiple frequencies.				
lit.			A3. Charact	erize surface	motion over the tidal cycle.				
Explore Europa to investigate its habitability.	A.		A4. Determi	ne the satell	ite's dynamical rotation state.				
bit			A5. Investig	ate the core,	rocky mantle, and rock-ocean interface.				
hal		Characterize the ice shell	B1. Charact	erize the dis	ribution of any shallow subsurface water.				
		and any subsurface water,	B2. Search t	for an ice-oc	ean interface.				
e ii	Ice	including their heterogeneity, and							
ate	В.	the nature of surface-ice-ocean			atures and subsurface structure to investigate processes governing material				
tig		exchange.			surface, ice shell, and ocean.				
est					al and global heat flow variations.				
UV	7	Determine global surface			organic and inorganic chemistry, including abundances and distributions of				
) i	Chemistry	compositions and chemistry,			nasis on indicators of habitability and potential biosignatures.				
a te		especially as related to habitability.	C2. Relate c	ompositions	to geological processes, especially material exchange with the interior.				
gdo			C3. Charact	erize the glo	bal radiation environment and the effects of radiation on surface composition,				
ırı);		atmospł	neric compos	sition, albedo, sputtering, sublimation, and redox chemistry.				
E			C4. Charact	erize the nat	ure of exogenic materials.				
ıre	gy	Understand the formation of	D1. Determi	ne the form	ation history and three-dimensional characteristics of magmatic, tectonic, and				
bld		surface features, including sites of	impact l	andforms.					
[X]	eol	recent or current activity, and	D2. Determi	ne sites of n	nost recent geological activity, and evaluate future landing sites.				
	_	identify and characterize candidate	D3. Investig	ate processe	s of erosion and deposition and their effects on the physical properties of the				
	D	sites for future in situ exploration.	surface		of the physical properties of the				
			31112300						
JEO T	JEO Themes: Origins Evolution Processes			Life	Based on 2002 Decadal's "objectives of solar system exploration"				









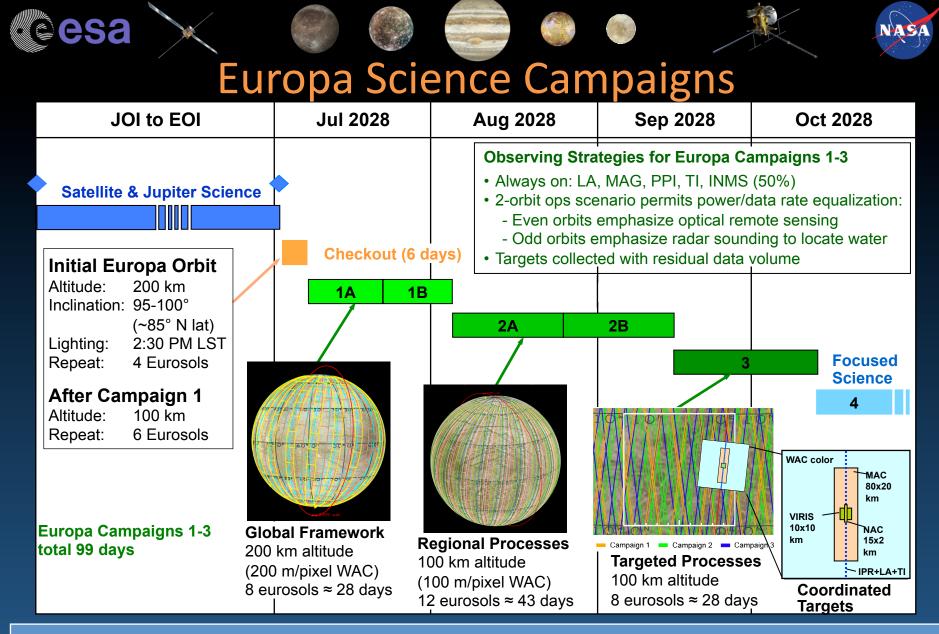






JEO Traceability: Jupiter System Science

Goal	Science Objective		Science Investigation				
	Understand Europa in the context of the Jupiter system.	Satellite surfaces and interiors	 E1. Investigate the nature and magnitude of tidal dissipation and heat loss on the Galilean satellites, particularly Io E2. Investigate Io's active volcanism for insight into its geological history and evolution (particularly of its silicate crust) E3. Investigate the presence and location of water within Ganymede and Callisto. E4. Determine the composition, physical characteristics, distribution and evolution of surface materials on Ganymede. E5. Determine the composition, physical characteristics, distribution and evolution of surface materials on Callisto. 				
vestigate its	system science	itellite Atms.	 E6. Identify the dynamical processes that cause internal evolution and near-surface tectonics of Ganymede and Callisto. E7. Characterize the composition, variability and dynamics of Europa's atmosphere and ionosphere E8. Understand the sources and sinks of Io's crustal volatiles and atmosphere. E9. Determine the sources and sinks of the Ganymede and Callisto atmospheres. 				
Explore Europa to investigate its habitability.	E. Jupiter	Plasma and magnetospheres	 E10. Characterize the neutral atoms and molecules escaping Europa's gravity. E11. Characterize the composition of and transport in Io's plasma torus. E12. Study the pickup and charge exchange processes in the Jupiter system plasma and neutral tori. E13. Study the interactions between Jupiter's magnetosphere and Io, Ganymede and Callisto (incl. characterize Ganymede's magnetic field) E14. Understand the structure, composition and stress balance of Jupiter's magnetosphere. E15. Determine how plasma and magnetic flux are transported in Jupiter's magnetosphere. E16. Characterize the abundance of minor species (especially water and ammonia) in Jupiter's atmosphere to understand the evolution of the Jovian system, including Europa. E17. Characterize Jovian atmospheric dynamics and structure. 				
	Ö	36	E18. Characterize the properites of the small moons, ring source bodies and dust E19. Identify the dynamical processes that define the origin and dynamics of ring dust.				

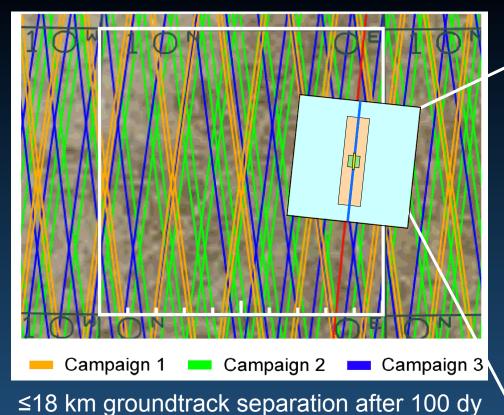


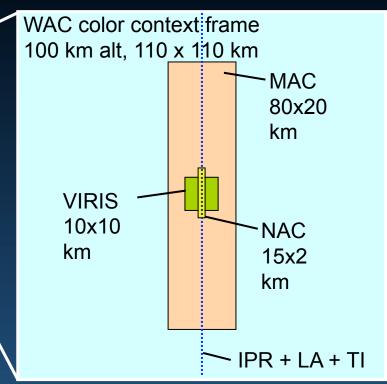
Europa science objectives addressed in first 100 days in orbit





Europa Science Campaigns: Profiling and Targeted Observations





290 Mb coordinated targets

~1700 coordinated targeted observations obtained after 9 mo.







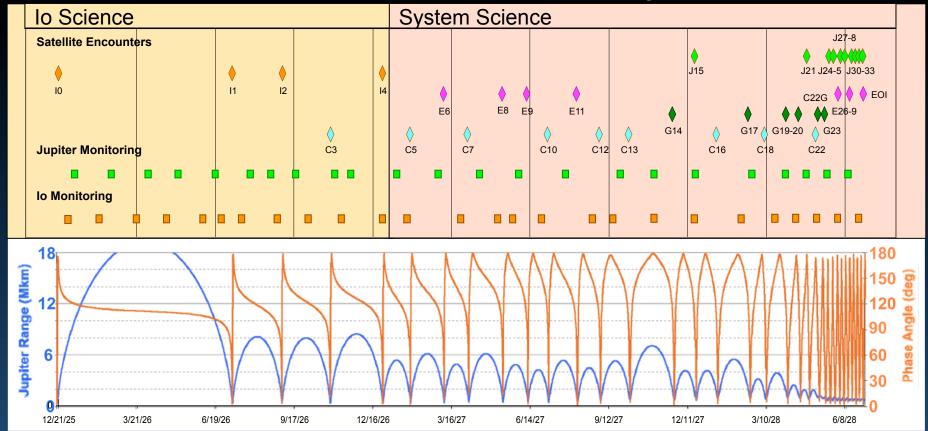








JEO Jovian Tour Example



- 33 perijoves during Jovian Tour
 - 23 with satellite flybys
 22 permit JEO-Earth radio occultations

Rich opportunities to acquire Jupiter System Science















JGO Goals and Objectives

- Key JGO science phases
 - Jupiter system: In-depth exploration
 - · From Jupiter orbit, synergistically with JEO
 - Callisto: In-depth study and mapping
 - · Multiple flybys using a resonant orbit
 - Ganymede: Detailed orbital study
 - · Elliptical orbit first, then circular orbit
- Science Objectives:
 - Ganymede: Characterize Ganymede as a planetary object, including its potential habitability
 - Satellite System: Study the Jovian satellite system
 - Jupiter: Study the Jovian atmosphere
 - Magnetosphere: Study the Jovian magnetodisk / magnetosphere
 - Jupiter system: Study the interactions occurring in the Jovian system



Characterizing the Jupiter system and its outer Galilean moons



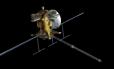














JGO Traceability: Ganymede

Goal	Science objective	Science investigation						
	A. Ice shell and	A1.	Time dependent altimetry and gravity to determine Love numbers h2 and k2.					
	ocean							
Ct			Study the magnetic field at multiple frequencies					
object /		A3.	Subsurface characterization - Determine the properties of the icy shell and the presence and location					
			of shallow liquid water.					
<u>`</u>		A4.	Constrain the amplitude of forced libration and obliquity and non-synchronous rotation					
planetary	B. Induced and	B1.	Globally characterize Ganymede's intrinsic magnetic field (to accuracy of 0.1nT).					
an	intrinsic magnetic	B2.	Characterize particle population within Ganymede's magnetosphere and its interaction with Jupiter's					
pli hat	fields	$\overline{}$	magnetosphere					
_ a _			Investigate the generation of Ganymede's aurora					
Ē. a			Study of the ionosphere and exosphere of Ganymede					
nede as potenti			Investigate surface composition and structure on open vs. closed field line regions					
ne	C. Geology and	C1.	Improve global and regional mapping					
Ganymede ng its poter	search for past and	$\overline{}$	Topographic mapping of large fractions of the surface.					
ian y it	present activity	C3.	Subsurface characterization					
_			Constrain global and regional surface ages					
iz	D. Surface comp.	D1.	Nature and location of non-ice and organic compounds					
جو ال	and physical							
ac i	properties of	D2.	Constrain the existence and rate of mass transfer processes					
Characterize includi	subsurface layers							
ည်	E. Deep interior	E1. Precise determination of low-degree static gravity field and shape						
		E2.						
		E3.	Search for deviations from hydrostatic equilibrium and for mass anomalies					

JEO Themes:

Origins

Evolution Processes Habitability

Life

To be modified wrt Cosmic Vision themes before June 2010















JGO Traceability: Satellites

Goa		ence obj										
		allisto: S	tudy		•	varying po	tential and sh	hape -	Time dependent altimetry and gravity to determin			
		urface		Love n	umbers							
	com	position,		F2. Study t	ne induced m	nagnetic fiel	d at multiple	frequ	iencies			
	phys	sical prop	perties,	F3. Subsur	ace characte	rization						
	puta	tive ocea					and organic c					
	inte	rnal stru	-						liquity and non-synchronous rotation			
Ε			-						field and shape			
ste e							fractions of th					
8	•		EJSM	F8. Charac	erization of	Callisto ion	osphere and e	exospl	here.			
9				F9. Constra	in the existe	nce and rate	of mass tran	nsfer p	processes between a) leading vs trailing			
_				hemisp	heres (role o	f impactors	and dust); b)	north	vs south hemispheres.			
Ite				F10. Constra	in global and	d regional si	urface ages					
Sa				F11. Improv	e imaging co	verage of C	'allisto's surfa	ace .				
an	G. I	o and Eu	_						plasma/neutral tori			
Š								itudes	and local times			
J.					erization of							
Study the Jovian satellite system	H. S	tudy the		H1. Charac	eristics and	chemical co	mposition of	the su	urfaces of outer irregular satellites			
<u>></u>	irre	irregular satellites (if close flybys are		H2. Astrom	etric observa	tions and m	ass determin	ation	of irregular satellites			
pn	(if cl											
S	feasi	feasible) EISM		H3. Search	for new oute	r irregular	satellites					
	I. In	vestigate	the	II. Physica	l characteriz	ation and cl	nemical comp	positio	on of the ring system in 3D and over different			
	inne	inner region of the Jupiter system including the ring		timesca	les and searc	ch for new a	ssociated sate	ellites	3			
	Jupi			I2. Charac	eristics and	chemical co	mposition of	f the su	urfaces of Thebe, Amalthea and other small inner			
	_			satellite	S							
	syste		EJSM	I3. Provide	improved e	phemerides	and mass est	timates	es for small inner satellites			
JEO The	JEO Themes: Origins Evolution			Processes	Habitability	Life	To be modif	fied v	wrt Cosmic Vision themes before June 2010			

















JGO Traceability: Jupiter

Goal	Science objective	Science investigation							
	J. The upper	J1.	Determination of general circulation & composition in the upper atmosphere						
	atmosphere	J2.	Characterization of the vertical coupling in the atmosphere & of its drivers, ion drag or wave						
			activity)						
		J3.	Temperature structure retrieval from upper atmosphere to the troposphere						
	EJSM	J4.	Characterization of ionospheric total electron densities & variations						
		J5.	Characterization of the wave activity at low- to mid-latitudes and eddy activity and eddy meridional						
ā			transport						
atmosphere	K. The stratosphere	K1.	Determination of the composition: H2O (characterisation of latitudinal variations, dynamics, role in						
lds			atmospheric chemistry); HCN (dispersion following the SL9 impact), hydrocarbons (stratospheric						
ě			chemistry) and haze; characterization of the strength of vertical mixing						
at .		K2.	Determination of temperature structure from stellar and solar occultations over a wide range of						
			latitudes in the upper stratosphere (1-km at 20 K per measurement).						
/ia			Determination of the general circulation in the stratosphere						
jo	L. The troposphere	L1.	Determination of chemical composition: condensable species (NH3, H2O) and disequilibrium						
<u>a</u>			species (PH3, CO)						
Study the jovian		L2.	Characterization of the strength of the vertical coupling in the atmosphere down to the troposphere						
ρn		L3.	Determination of the composition & vertical structure of clouds and cloud size distribution						
St			Study of the relation between the upper troposphere circulation & the deep circulation below the						
			clouds & processes driving the jets circulation.						
			clouds to processes driving the jets enculation.						
			Potential vorticity retrieval from combined dynamics and thermal measurements						
	M. Internal	M1.	Constrain the existence and size of a core, and the nature of the H-H2 phase transition -						
	structure of		F man a constant of the consta						
	Jupiter								
150 - 1		la di a	Process Helifability I I'm T. I. I'm I C.						
JEO Themes: Origins Evo		lution	Processes Habitability Life To be modified wrt Cosmic Vision themes before June 2010						











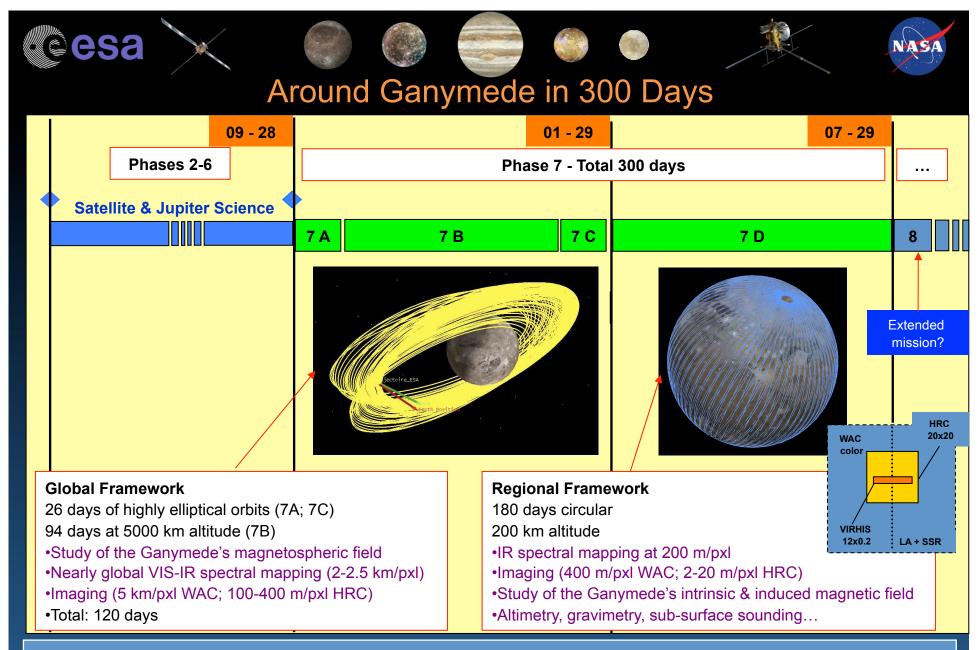






JGO Traceability: Magnetosphere & Jupiter System

Goal	Science objective	Science investigation
	N. The	N1. Characterize the properties of the magnetodisk with nearly 3D coverage in order to obtain good and
Ø	magnetosphere as a	reliable plasma moments (density, pressure, bulk flow velocity)
erc	fast magnetic	N2. Improve our understanding of the plasma processes acting in the magnetodisk
d d	rotator	N3. Investigate the plasma sources, mass loading variability, composition, transport modes, and loss
n os		processes in the magnetosphere
vian netosphere		N4. Study of the dust - plasma interactions
Jovian		N5. Characterize the large-scale coupling processes between the magnetosphere, ionosphere and
le .		thermosphere
ţ,		N6. Magnetospheric response to solar wind variability
Study the Jo magnetodisk/mag	EJSM	N7. Look for direct evidence of the effects of the solar wind and planetary rotation on driving
od		magnetospheric dynamics
S	O. The	O1. Characterize the time evolving Jovian radiation environment
gr	magnetosphere as a	O2. Improve our understanding of the particle bombardment of the surfaces of the moons
ma	giant accelerator	O3. Detail the particle acceleration processes
_	EJSM	O4. Study the loss processes of charged energetic particles
		O6. Observations of the moon auroral magnetic footprints
ns	P. Satellite / mag.	P1. Study of pick-up & charge-exchange processes in plasma/neutral tori
octions jovian	interactions: the	
10	magnetosphere as a	P2. Search for plasma effects on satellites (including irregular)
in the	magnetized binary	P3. Analysis of absorption signatures by moons, rings and dust
	system	13. Pinarysis of absorption signatures by moons, rings and dust
	Q. Tidal coupling	Q1. Determine short-term and long-term changes of the orbits of the Galilean satellites and the inner
	among Jupiter and	satellites
Study occur	the setallites	Q2. Study the coupled evolution of Io Europa and Ganymede by determining internal structures, heat
Sti	the satemites EISM	flows, and tidal responses (including tidal phase lags) of the moons.
EO Ther	mes: Origins Evoluti	on Processes Habitability Life To be modified wrt Cosmic Vision themes before June 2010



Ganymede science objectives addressed in 300 days in orbit









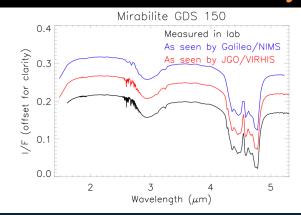


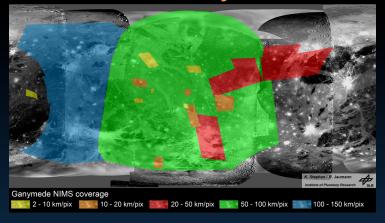






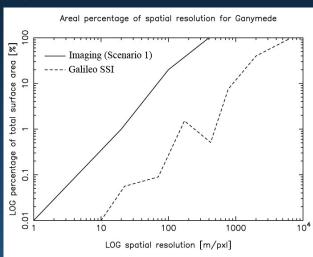
Around Ganymede in 300 Days: Some Examples

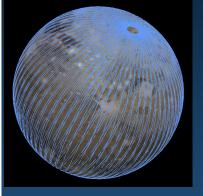




Remote sensing

Significant improvement in spectral (x5) and spatial (80% at 2-2.5 km/pxl) resolution.





In situ continuous acquisition

Full investigation of intrinsic and induced magnetic fields

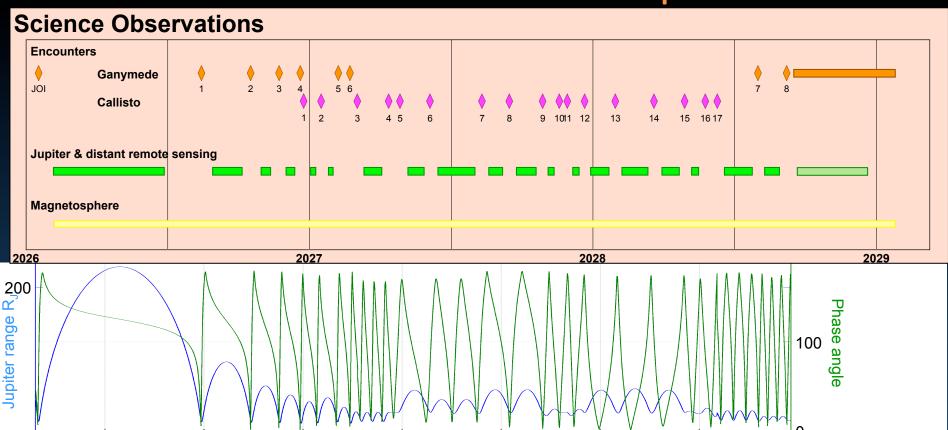


> 400 Gb of compressed data





JGO Jovian Tour Example



- 17 Callisto flybys
- 8 Ganymede flybys

33 perijove at ~15 R_J

Rich opportunities to acquire Jupiter System Science



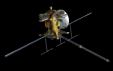






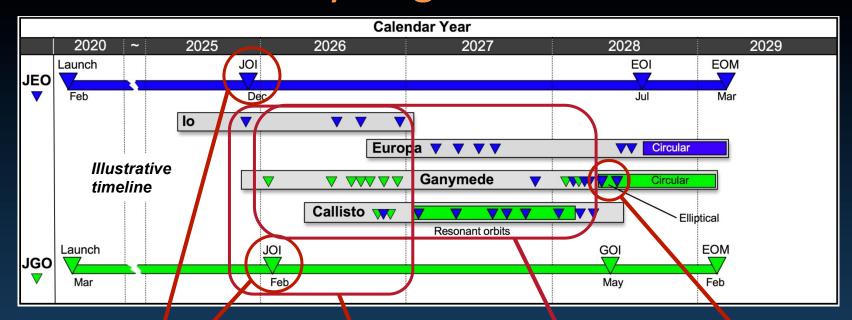


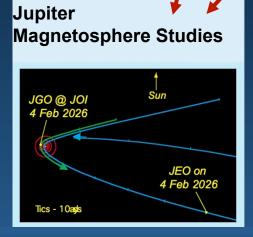






EJSM Synergistic Science

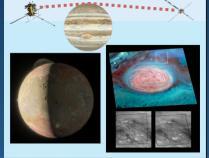








Satellite & Jupiter Monitoring; Radio Occultation Science?



Ganymede Magnetosphere Studies

















OPAG Feb 2010: What's Needed?

- Discussion sessions (Tues. morning)
 - Satellites (Dave Senske, Olivier Grasset)
 - Jupiter (Amy Simon-Miller, Leigh Fletcher, Bob Pappalardo)
 - Review the science
 - "Walk through" the example Jovian Tour
 - Make recommendations
- Jupiter System Science (Melissa McGrath)
 - Synergistic Magnetospheric Science (Norbert Krupp)
 - General Discussion (all)

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