

Advances in solar forecasting: Computer vision with deep learning

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ABSTRACT

Renewable energy forecasting is crucial for integrating variable energy sources into the grid. It allows power systems to address the intermittency of the energy supply at different spatiotemporal scales. To anticipate the future impact of cloud displacements on the energy generated by solar facilities, conventional modeling methods rely on numerical weather prediction or physical models, which have difficulties in assimilating cloud information and learning systematic biases. Augmenting computer vision with machine learning overcomes some of these limitations by fusing real-time cloud cover observations with surface measurements acquired from multiple sources. This Review summarizes recent progress in solar forecasting from multisensor Earth observations with a focus on deep learning, which provides the necessary theoretical framework to develop architectures capable of extracting relevant information from data generated by ground-level sky cameras, satellites, weather stations, and sensor networks. Overall, machine learning has the potential to significantly improve the accuracy and robustness of solar energy meteorology; however, more research is necessary to realize this potential and address its limitations.

1. Introduction

1.1. Solar forecasting

Solar forecasting is one of the most effective and efficient techniques to mitigate the solar power variability and uncertainty caused by atmospheric changes [1]. Most solar forecasting methods take ground- or remote-sensing information as input to infer current meteorological conditions, and predict the future global solar irradiance (GSI) or power, ranging from seconds to years ahead (Fig. 1).

Solar forecasting has been extensively used in the power and energy industry; it is also known as operational solar forecasting (Section 3.2.2). According to different lead times and horizons, solar forecasting can be roughly categorized into very short-term forecasting, short-term forecasting, medium-term forecasting, and long-term forecasting. Although there is a lack of consistency in these definitions, solar forecasts at all timescales are important to power system operations and planning. As shown in Table 1, long-term solar forecasts are usually used in long-term energy trading, system planning, and resource assessment. Short-term and very short-term solar forecasts, normally

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Abbreviations	
AI	Artificial Intelligence
ARIMA	Autoregressive Integrated Moving Average
ASI	All-Sky Imager
CNN	Convolutional Neural Network
CSP	Concentrated Solar Power
CRPS	Continuous Ranked Probability Score
CSI	Clear-Sky Index
COMS-MI	Communication Ocean and Meteorological Satellite equipped with Meteorological Imager
CVAE	Convolutional Variational Autoencoder
DNI	Direct Normal Irradiance
ECMWF	European Centre for Medium-Range Weather Forecasts
ERA5	ECMWF Reanalysis 5th
FOV	Field of View
GAN	Generative Adversarial Networks
GOES	Geostationary Operational Environmental Satellites
GHI	Global Horizontal Irradiance
GSI	Global Solar Irradiance
HDR	High Dynamic Range
HYTA	Hybrid Thresholding Algorithm
JMA	Japan Meteorological Agency
LIDAR	Laser Imaging, Detection, And Ranging
LSTM	Long Short-Term Memory
MAE	Mean Absolute Error
MBE	Mean Bias Error
MSG	Meteosat Second Generation
MTG	Meteosat Third Generation
MERRA	Modern-Era Retrospective analysis for Research and Applications
MIMO	Multi-Input Multi-Output
MOF	Methods Of Fusion
MSE	Mean Squared Error
MTL	Multitask Learning
NASA	National Aeronautics and Space Administration
NOAA	National Oceanic and Atmospheric Administration
NREL	National Renewable Energy Laboratory
NSRDB	National Solar Radiation Database
NWP	Numerical Weather Prediction
PCA	Principal Component Analysis
PIV	Particle Image Velocimetry
RGB	Red Green Blue
RMSE	Root Mean Square Error
RNN	Recurrent Neural Network
ROC	Receiver Operating Characteristic
ROI	Region Of Interest
SIMO	Single-Input Multi-Output
SEVIRI	Spinning Enhanced Visible and InfraRed Imager
SVR	Support Vector Regression
TSI	Total Sky Imager
XAI	Explainable Artificial Intelligence
WRF	Weather Research and Forecasting
PICP	Prediction Interval Coverage Probability
PINAW	Prediction Interval Normalized Averaged Width

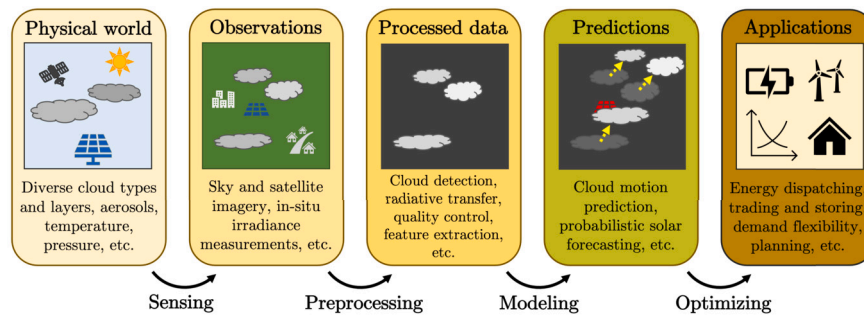


Fig. 1. Solar energy meteorology framework reviewed in the present study.

Table 1
Solar forecasting classification by timescales.

Group	Lead Time	Horizon	Application
Long-term	Months to years	Years	Capacity expansion
Medium-term	Weeks to months	Weeks to months	Capacity market
Short-term	Hours to days	Hours to one week	Day-ahead unit commitment
Very short-term	Minutes to hours	Minutes to hours	Real-time economic dispatch

defined as predictions up to a few days ahead, are used in daily power system operations and energy markets. These definitions and terminologies differ between power systems and markets. It is common practice for system operators and market participants to use medium-/short-term solar forecasts for day-ahead bidding or preliminary generator scheduling and use very short-term solar forecasts for real-time adjustments [2]; therefore, accurate solar forecasts can enhance power system reliability and resilience while reducing economic costs.

Solar forecasting techniques can be categorized into physical models [3], data-driven models [4], or hybrid models [5], depending on the involvement of physical laws. Both ground-sensing (i.e., *in situ* or mobile data) and remote-sensing data can be used as input to any type of

model [6]. The most widely used remote-sensing data are satellite data, which can be processed by physical models, such as numerical weather prediction (NWP) models, to provide solar forecasts [7]. Satellite data can also be directly processed by statistical models to generate solar forecasts [8]. The most effective ground-based sensors include pyranometers and pyrhemometers, which provide local surface irradiance measurements. Compared to other target variables in energy forecasting (e.g., load forecasting, price forecasting, and wind forecasting), solar forecasting is most sensitive to cloud movement; hence, it provides unique opportunities for sensors, such as satellites and sky cameras, to play critical roles. Fig. 2 depicts the spatiotemporal resolutions and coverages of different types of sensors, where sky cameras and satellites

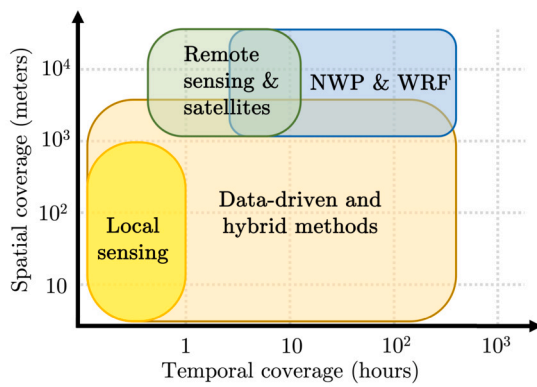


Fig. 2. Temporal and spatial coverage of different solar forecasting techniques. Adapted from [9].

cover a large spectrum of solar forecasting. One important common ground of satellites and sky cameras is that both sensors can provide data in the format of images. These observations of the cloud cover can be processed by advanced computer vision techniques to model the variability of solar power.

1.2. Computer vision in solar forecasting

The cloud cover dynamics accounts for most of the intraday irradiance variability. To better anticipate corresponding solar power fluctuations, computer vision methods based on cloud observations have been developed. Besides an inherent stochastic component in cloud modeling, the partly deterministic nature of cloud motion can be exploited to extrapolate the future cloud cover spatial configuration from past observations. The resulting ability to anticipate cloud displacement offers a significant advantage over forecasting approaches solely based on local meteorological measurements, which fail to offer reliable predictions beyond the short-term autocorrelation period of solar time series. This forecasting strategy is therefore critical to provide truly valuable solar predictions at different spatiotemporal scales [10].

Definition. A solar forecasting method is described as image-based when the covariates include sky features extracted from spatial measurements acquired by light sensing devices (e.g., ground-based imager, satellite imagery, or both). In contrast, physical approaches are based on the position of the sun, whereas statistical methods are based on an autoregressive and moving average covariate structure (i.e., time series analysis).

The principal sources of cloud observations for short-term solar forecasting are sky cameras and satellites. A sky camera equipped with a fish-eye lens is able to capture the local cloud cover spatial configuration over a few square kilometers at a subminute time period, whereas satellites provide continent-wide radiation maps of the atmosphere across a large spectrum at a lower temporal resolution of 5 to 15 minutes [11,12] (Fig. 2). Although the spatiotemporal resolution of remote sensors is expected to increase and close the gap with ground-level imagery, these two imaging technologies are currently complementarily used in solar forecasting due to their different technical characteristics and points of view on the cloud cover [13]. Regarding the forecast horizon, approaches based on sky cameras aim to provide predictions on the local solar variability up to 30 minutes ahead, depending on the local cloud velocity, whereas satellite imagery provides longer-term forecasts, up to several hours ahead, thanks to its wider field of view (FOV) [14]. Unless used in a network setup covering a larger area [15], sky cameras are best suited for locations with a high solar power density, such as a solar plant or a hybrid power plant (Fig. 3). In contrast, satellite-based techniques can provide countrywide solar output predictions benefiting applications such as grid balancing or energy trading.

Traditional computer vision approaches based on sky images or satellite images share similar methods: cloud segmentation, cloud localization, cloud properties estimation, cloud tracking, and cloud motion modeling [17]. The cloud mask in sky images is often derived from thresholding methods based on pixel value statistics [18]. In addition to similar thresholding tests [19], cloud segmentation in satellite images can be further augmented using the effective cloud albedo or cloud index (0: absence of cloud, 1: thick cloud) computed from the ground albedo, the observed albedo, and the maximum empirical albedo [20]. This continuous variable indicates the cloud transmittance, which can also be estimated with a pyrhelimeter and a clear-sky model for sky camera applications [21]. Cloud geolocation is then performed using several sky cameras in stereovision mode [22,23]. Following this, cloud velocity vectors can be approximated, with block matching methods [24–26] or with optical flow [27,28], to predict the future impact of the cloud cover dynamics on solar generation in a deterministic [29,30] or probabilistic manner [31]. From the future position of clouds, it is possible to derive local or regional irradiance forecasts using the estimations of the clouds' physical properties [32,33] (Fig. 4).

Benefiting from recent advances in deep learning and the increasing accessibility of large datasets [34] (see Section 2.5), a range of data-driven approaches have been developed to address this computer vision task. The standard approach consists of training a neural network to extract spatiotemporal features from a sequence of past sky or satellite images to predict the future solar variability. Similar to traditional computer vision methods, both local values (e.g., GSI or photovoltaic power output) [35–37] and regional predictions [38] (e.g., GSI or cloud-index maps) can be obtained with this class of techniques. Despite high performance gains based on averaging metrics, including the root mean square error (RMSE) or the mean absolute error (MAE), deep learning models still face several limitations, such as difficulty in breaking the persistence barrier to predict critical events on time [39,40] or to correlate an image (especially satellite images) with its corresponding solar value (GSI or photovoltaic power output). As a result, some models rely on auxiliary data, including GSI measurements or the solar elevation.

1.3. Bibliometric analysis

Computer vision-based solar forecasting has drawn increasing attention in the solar community, as reflected in the number of publications on this topic in the recent 12 years. Fig. 5 shows the total number of journal articles, conference articles, and book chapters published each year. It is visible that the computer vision-based solar forecasting research has exponentially increased in recent years. In 2010 and 2011, satellite-based solar forecasting dominated the field. From 2012 to 2017, however, sky image-based solar forecasting played a leading role. In the past four years, satellite- and sky image-based solar forecasting have been developed in similar proportions. Fig. 6 lists the top 12 journals in terms of number of related publications (note that publication quantity does not imply quality). *Solar Energy*, the official journal of the International Solar Energy Society, is the leading journal in computer vision-based solar forecasting. Most of the well-known journals in this field are energy journals, except for *Remote Sensing* and *Atmospheric Measurement Techniques*. The former covers various aspects of remote-sensing science, where image processing is a major publishing scope. The latter is dedicated to publishing advances in remote-sensing, *in situ*, and laboratory measurement techniques for the constituents and properties of the Earth's atmosphere, including irradiance sensing and measurement research.

Although all these journals publish both ground-sensing and remote-sensing solar forecasting articles, most of them have imbalanced amounts of publications in the two areas. For example, *Solar Energy*, *Atmospheric Measurement Techniques*, *Journal of Renewable and Sustainable Energy*, and *Energy Conversion and Management* predominantly focus on ground sensing forecasting articles, whereas *Remote Sensing* and *Renewable and Sustainable Energy Reviews* specialize in remote-sensing solar

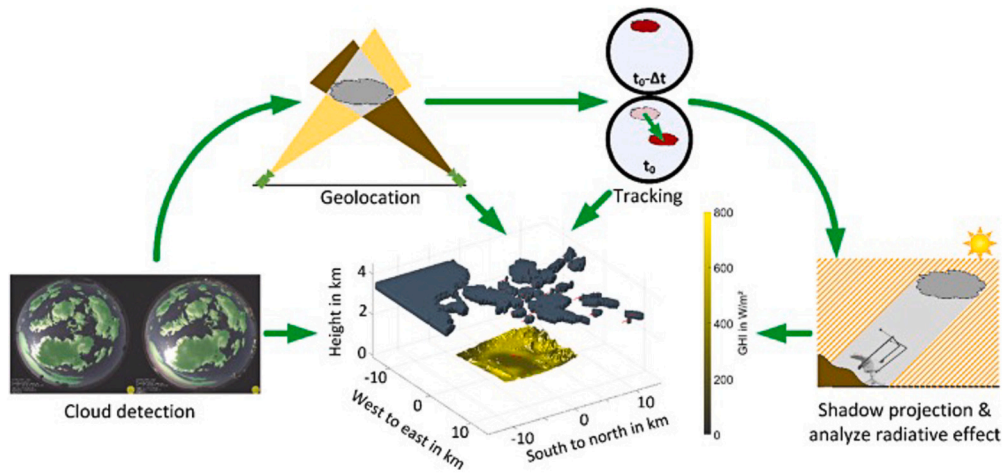


Fig. 3. Illustration of a standard physics-based solar forecasting approach using sky cameras in stereovision mode: cloud detection, geolocation, tracking, shadow projection, and analysis of radiative effect. Taken from [16].

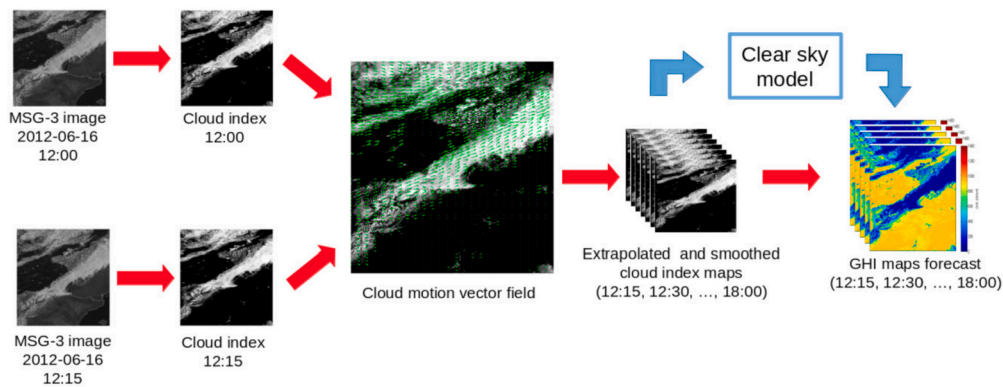


Fig. 4. Example of a physics-based solar forecasting method using satellite observations: cloud detection via cloud index mapping, cloud motion estimation, extrapolation to future time steps, and radiative transfer modeling. Taken from [33].

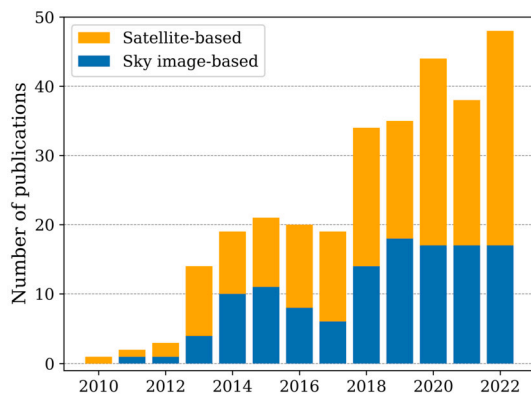


Fig. 5. Number of publications in satellite-based and sky image-based solar forecasting.

forecasting articles. This bibliometric analysis suggests the increasing research trend and top journals, which can help researchers track the state of the art in this field.

1.4. Existing review articles and motivation for the new review

Since 2010, solar forecasting has witnessed exponential growth, largely due to increasing data availability and emerging techniques, together with the rising value of solar predictions. The first wave of solar forecasting development resulted in several classic and highly cited re-

view articles (Table 2). A comprehensive and detailed review covered clear-sky models, regressive methods, remote-sensing models, NWP models, and limited types of artificial intelligence (AI) techniques [8] for solar forecasting. The same year, a similar but briefer review focused on three types of solar forecasting techniques: statistical methods, cloud imagery and satellite-based models, and NWP models [41]. A more recent review article summarized and compared techniques for direct normal irradiance (DNI) forecasting and applications to concentrated solar power (CSP) output forecasting. Antonanzas et al. [42] reviewed specific techniques for photovoltaic power output forecasting grouped by spatial and temporal characteristics. Probabilistic solar forecasting was emphasized in the same work.

Machine learning techniques started to be applied to solar forecasting in the early 2010s and were first systematically reviewed in [4]. This work grouped machine learning methods into classification, supervised learning, unsupervised learning, and ensemble learning. The recent advances in computational power enabled deep learning methods to be widely implemented in solar forecasting. The first review article for deep learning solar forecasting methods was published in 2021, which also covered wind forecasting methods [43]. In the same year, two similar review papers were published in slightly different scopes: GSI forecasting and satellite image prediction [44,45]. The development of machine learning and deep learning has also led the advances in intra-hour solar forecasting, which could also be reflected by some recent review articles [9,46,47].

The rapid advances in the emerging field of computer vision-based solar forecasting with deep learning motivate this review. This work

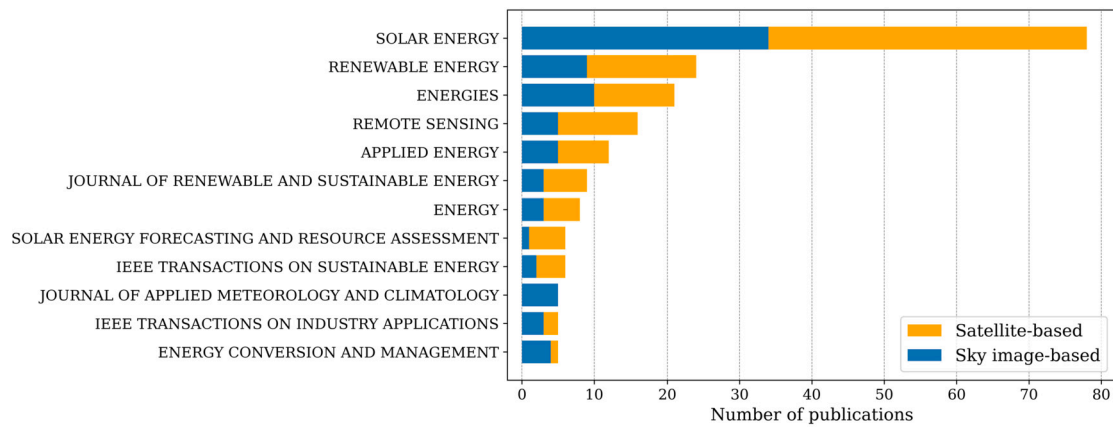


Fig. 6. Top 12 journals with the most computer vision-based solar forecasting publications.

Table 2

Solar forecasting review articles reference, publication year, and scope.

Article	Year	Scope
Inman et. al [8]	2013	Theory and application of utility-scale solar forecasting
Diagne et. al [41]	2013	GSI forecasting and proposition for small-scale insular grids
Law et. al [48]	2014	DNI forecasting
Antonanzas et. al [42]	2016	Latest advancements and future trends in solar power forecasting
Voyant et. al [4]	2017	Machine learning methods of GSI forecasting
Barbieri et. al [49]	2017	Very short-term photovoltaic power forecasting with cloud modeling
Sobri et. al [50]	2018	Photovoltaic power forecasting techniques
Kurzrock et. al [51]	2018	Geostationary satellite-based short-term cloud forecasting
Yang et. al [52]	2018	Technological infrastructure and key innovations in solar forecasting
Sweeney et. al [2]	2019	State-of-the-art and future solar and wind forecasting
Kumar et. al [53]	2020	Solar forecasting methods and sensor networks
Hong et. al [54]	2020	Energy (load, solar, wind, price) forecasting
Li et. al [55]	2020	Probabilistic solar forecasting applications in power systems
Guermoui et. al [56]	2020	Hybrid models for solar radiation forecasting
Ahmed et. al [57]	2020	Solar forecasting methods and optimization
Wang et. al [58]	2020	Taxonomy of AI for solar power forecasting
Juncklaus Martins et. al [59]	2021	Forecasting/nowcasting based on ground-based cloud imaging
Alkhatay et. al [43]	2021	deep learning-based solar and wind forecasting methods
Sharma et. al [46]	2021	Sky image-based short-term intra-hour solar prediction
Kumari et. al [44]	2021	deep learning GSI forecasting models
Moskolaï et. al [45]	2021	deep learning-based satellite image prediction
Sawant et. al [60]	2021	Cloud detection, identification, and forecasting
Chu et. al [9]	2021	Intra-hour GSI forecasting
Fan et. al [47]	2022	Ground-based sky image-based intra-hour solar forecasting
Erdener et. al [61]	2022	Aggregate regional behind-the-meter solar generation forecasting
Yang et. al [62]	2022	Solar forecasting from atmospheric science and power engineering perspectives
Lin et. al [12]	2023	Intra-hour solar forecasting with ground-based sky image methods
Krishnan et. al [63]	2023	How solar radiation forecasting impacts the utilization of solar energy

considers the diverse aspects of this field, from the datasets [34,64,65] to end-user applications [66] and cloud modeling techniques [67]. The review primarily focuses on computer vision methods using deep learning to analyze spatial data (images or videos) for solar power modeling [36,35,68–71].

The remainder of this article is organized as follows. Section 2 presents data acquisition and preprocessing techniques used to model solar power via cloud cover observations. Section 3 describes the fundamentals of computer vision-based solar forecasting, including computer vision tasks, evaluation methods, and end-user applications. A list of accessible resources, datasets, and codes is described in Sections 2.5 and 3.4. Section 4 focuses on the deep learning methods applied to solar power modeling with computer vision such as data fusion, transfer learning, multitask learning, data-centric techniques and interpretable AI. Future research challenges as well as the state of the adoption of vision-based forecasting technologies are discussed in the penultimate Section 5. Section 6 concludes the study.

2. Data collection and preprocessing

From the perspective of traditional computer vision methods, the information in a scene has both global and local features [72]. In solar forecasting, sky images contain local information about atmospheric conditions, whereas ground-level GSI measurements are global features represented by the sky images. The section about data is divided into atmospheric imagery and *in situ* measurements (Sections 2.1 and 2.2). In addition, effective image preprocessing methods for cloud cover observations are described in Section 2.3, and the derivation of the physical features to incorporate into physics-informed deep learning models is mentioned in Section 2.4.

2.1. Cloud cover observations

This section summarizes different data acquisition systems and their technical specifications, in addition to recent technological advances.

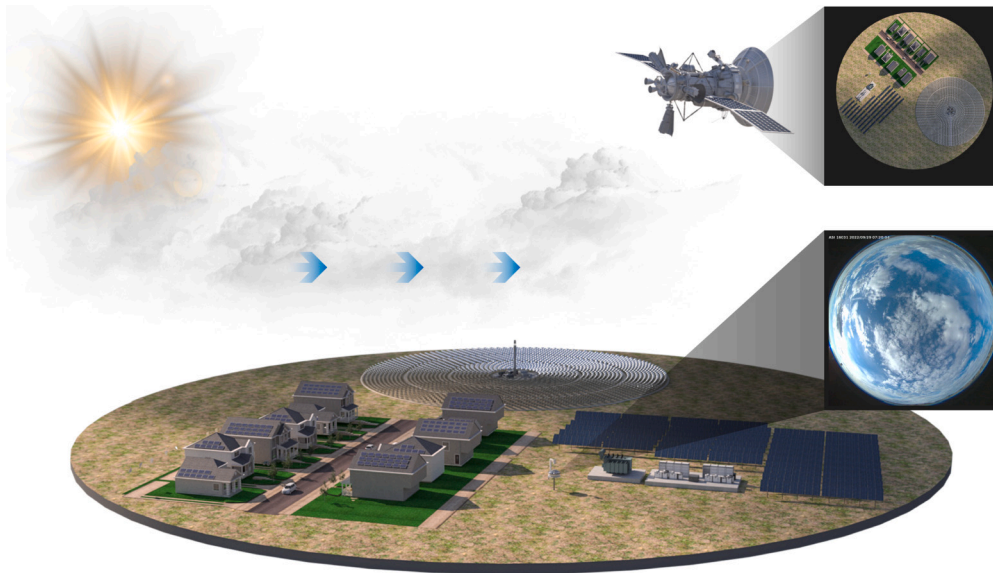


Fig. 7. Meteorological influences such as moving clouds can cause a significant variability in the production of energy from photovoltaic or concentrated solar power. Observing the cloud cover in the vicinity of solar sites via satellites or ground-level sky cameras helps to anticipate these fluctuations.

Information about publicly accessible datasets for direct download is presented in Section 2.5.

2.1.1. Satellite imagery

Geostationary satellites are the most common data acquisition system for solar forecasting, as they provide continuous imagery over large areas of Earth, albeit at lower spatial resolutions (around 1–3 km per pixel) than lower orbiting satellites such as Landsat. This enables them to quickly see changing cloud and atmospheric conditions over nearly any solar site, and be available as input for any forecast time. The Geostationary Operational Environmental Satellites (GOES-16) operated by the National Oceanic and Atmospheric Administration (NOAA) provides measurements covering 16 spectral bands every 5 to 10 minutes [73]. The satellite Fengyun-4A, managed by the China Meteorological Administration, provides measurements from 14 spectral bands [74]. The Communication Ocean and Meteorological Satellite equipped with Meteorological Imager (COMS-MI), operated by the Korean Meteorological Association, observes 8 spectral bands [75]. The Meteosat Second Generation (MSG), operated by EUMETSAT, is equipped with the Spinning Enhanced Visible and InfraRed Imager (SE-VIRI) capable of sensing the Earth using 12 different spectral bands every 5 to 15 minutes [37,38,76,13]. The Himawari-8/9 satellite, operated by the Japan Meteorological Agency (JMA), provides 16 different spectral bands every 10 minutes [77]. The future generations of geostationary satellites like GOES-18 and Meteosat Third Generation (MTG) will provide improved spatiotemporal resolution and richer spectral information (e.g., from 12 channels for MSG to 16 channels for MTG, and from 5 to 15 minutes temporal resolution down to 2.5 to 10 minutes). Non-geostationary satellites, such as the Landsat series of imagers or MODIS satellite can provide higher resolution imagery, on the order of 0.5 to 30 meters per pixel, but with revisit times on the order of hours or days. Imagery from these satellites can be used for solar potential estimation [78–81] but are generally not suitable for solar nowcasting or short-term forecasting, because of their limited temporal resolution.

2.1.2. Ground-based sky imagers

The most commonly used sky images in solar forecasting applications are in the Red Green Blue (RGB) color model. The total sky imager (TSI) was first introduced in the atmospheric science community to quantify the fraction of the sky with clouds [82]. The TSI is a reflective sky imager that acquires sky images with a large FOV. A TSI records sky images by reflecting light beams using a convex mirror to converge

on the focal point of a visible light camera. A decade later, the all-sky imager (ASI), i.e., skycam or whole-sky imager, was proposed as a low-cost alternative to the TSI [83]. The ASI uses an inexpensive camera with a fish-eye lens attached to enlarge the FOV. The TSI and ASI suffer from saturation of the pixels in the circumsolar area, but different alternatives to reduce the saturation of circumsolar pixels have been proposed since their introduction. Some of these alternatives use the Y16 brightness model (i.e., high dynamic range -HDR- and infrared images) instead of RGB.

2.1.3. Sun tracking

The sky conditions affect in situ GSI measurements mainly by passing clouds (i.e., stochastic effects) or other aerosols suspended in the atmosphere. These effects are commonly known as direct and diffuse components. In particular, assessing the effects of the direct component requires estimating when a cloud will intersect with the sun (i.e., inner circumsolar region), while the diffuse component effects require estimating the cloud density and coverage (i.e., outer circumsolar region). In any case, it is necessary to predict the future position of the sun in the sky images to identify which air parcel will cover the sun. The sun follows a trajectory in the sky varying daily, but it is deterministic and can be easily estimated.

A possible solution is mounting a sky imager on a solar tracker reducing the complexity of the problem to moving clouds [84,85]. Fast and accurate solar position algorithms exist [86–89]. However, there are precision errors in the estimation of the solar position that affect the accuracy of the clear sky index documented earlier [90]. These errors have been thoroughly assessed [91], and more recent research concluded that a consequence of the solar zenith angle estimation [92], which is particularly important for ground-based sky imagers mounted on a solar tracker. A proposed solution to this problem is to synchronize the solar tracker with satellite positioning [93] or use ground-based solar sensors [94]. Another alternative is to apply image processing methods and directly remove the sun from the images [95]. However, the uncertainty in the physical size of the pixels poses a limitation to this approach. An alternative to solar position algorithms is to detect the position of the sun in sky images to drive the solar tracker and center the images around the sun [96]. An approach following this trend proposed a model to predict the solar position throughout the day in sky images from past observations [97] (Fig. 8).

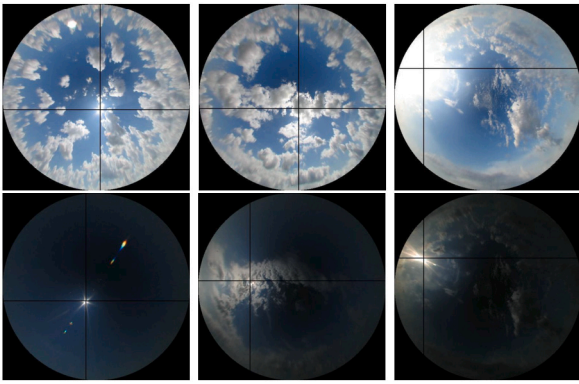


Fig. 8. Position of the sun in long- and short-exposure sky images predicted by an image-based sun tracker from past observations. Knowing the position of the sun when it is hidden by a cloud (middle column) or surrounded by a large saturated area (top right image) provides valuable information for short-term solar forecasting. Taken from [97].

2.1.4. Technological challenges

As shown in the top right image of Fig. 8, the brightness temperature of solar radiation saturates pixels in visible light sky imagers (i.e., RGB). This effect removes otherwise useful information for computer vision-based solar forecasting. For that reason, a sun-blocking mechanism is generally attached to a sky imager [98]. This mechanism is mounted on a 1-axis solar tracker and is similar to the shadow bands used to measure diffuse radiation with pyranometers [99]. Alternative mechanisms that remove less information from sky images exist but are not commercially available [100].

Visible light cameras record multiple frames per second (i.e., > 15 fps). It is possible to exploit this to improve the quality of the sky images and reduce the noise-to-signal ratio [85], and to adjust the exposure time (i.e., control the shutter) for recording frames with different exposure times [101]. This allows for a saturation reduction of circumsolar pixels via software. Other approaches following this trend propose nonparametric methods for fusing sun-centered sky images at multiple exposure times with reduced noise-to-signal ratios to generate HDR 16-bit grayscale sky images [85]. This method was developed for an ASI mounted on a 2-axis solar tracker (Fig. 9).

Optical filters are another alternative to reduce pixel saturation in the circumsolar area. Neutral density filters uniformly attenuate the solar radiation in the entire light spectrum [85], whereas near-infrared filters are capable of reducing the saturation of the pixels in the circumsolar area [98,102]. However, the most effective method to reduce the saturation of the circumsolar pixels is to use far-infrared sky images [103]. These types of cameras are made of low-cost uncooled microbolometers [104]. As visible light sky imagers, far-infrared imagers can have a large FOV when reflecting the entire sky over an aluminum sphere [105]. These approaches are similar to TSIs and are commercially available.¹ Additionally, it is possible to reproduce the capabilities of an ASI by fusing infrared images acquired from multiple sensors, producing static far-infrared sky images with a large FOV [106]. Similarly, when a small FOV far-infrared camera is mounted on a 2-axis solar tracker, it is possible to acquire sun-centered sky images with low saturation of the pixels in the circumsolar region [85].

2.2. In situ measurements

A computer vision-based solar forecasting model intrinsically aims to forecast GSI measured on the ground, or photovoltaic power output, by analyzing the movement of passing clouds using sky or satellite

images. Consequently, a dataset for these applications must include consecutive sequences of sky or satellite images (i.e., covariates) paired with ground GSI measurements or photovoltaic power output (i.e., predictors).

2.2.1. Global solar irradiance

A pyranometer, located on a horizontal surface normal to the ground, measures GSI. The GSI, \mathcal{R}_{GSI} , can be decomposed in DNI, \mathcal{R}_{Direct} , and diffuse, $\mathcal{R}_{Diffuse}$, and reflected, $\mathcal{R}_{Reflected}$, irradiance, such as $\mathcal{R}_{GSI} = \mathcal{R}_{Direct} + \mathcal{R}_{Diffuse} + \mathcal{R}_{Reflected}$ [107]. The DNI is measured using a pyrliometer, but its approximation is possible from GSI measurements [108]. DNI is necessary to estimate CSP output [109]. In contrast, GSI is a requirement to approximate the power output of photovoltaic power plants [110]. The GSI includes diffuse irradiance, which quantifies the scattering effect produced by particles or molecules floating in the atmosphere (i.e., clouds) [111]. When estimating the power output from a GSI forecast, it is necessary to know photovoltaic system design parameters; the number of photovoltaic arrays and active area (i.e., shadows and dust) [112,113]; the manufacturer's specific parameters, such as module and inverter efficiency; plus the local degradation pattern of the photovoltaic system, which requires an empirical model based on data for accurate estimation [114]. But there are simplified models [115] that are commonly used in operational planning applications [116].

2.2.2. Clear-sky index

The GSI signal has a deterministic and a stochastic (i.e., random) component. The deterministic component is considered to be the cyclostationary effects produced by Earth's rotation and revolution [95]. These movements cause the GSI to be a periodic function with local maxima at noon. The amplitude of the daily local maxima varies yearly, with global maxima in the summer solstice and a global minimum in the winter solstice. The effects produced by clouds and aerosols are the stochastic component. The random effects produced by aerosols and clouds depend on the distance that a light beam travels through the atmosphere [117]. These effects are greater during the sunrise and sunset hours. The clear-sky index (CSI) is a stochastic component, defined as $i_k = r_k / \hat{r}_k$, where r_k is the pyranometer signal, and \hat{r}_k is the deterministic GSI computed from a physical model [118]. This practice is known as *detrending* in time series analysis (Fig. 10).

Previous work validated the performance of deep learning solar forecasting using CSI or GSI and found that the CSI reduces the performance [119]; however, more recent investigations used CSI as the predictor instead of GSI, which is becoming a common practice in the field to increase the performance of a solar forecast [117,120]. Additionally, the pyranometer signal might have biases with respect to the deterministic component that are only noticeable in high-resolution measurements [95]. These biases are produced by slight inclinations in the pyranometer's support or deviations with respect to the deterministic component. These biases are stationary and deterministic; thus, their removal increases the performance of a solar forecast [121,122].

2.3. Image preprocessing

This section describes existing image processing methods known for increasing the performance of solar forecasts. The key idea behind applying preprocessing to the sky images is to remove complexity from the input space to facilitate training a deep learning model, thus speeding up or increasing the performance.

2.3.1. Irradiance maps

There are other ground-based approaches available for indirectly visualizing cloud motion (i.e., projected shadows) using information recorded from the surface. These methods are particularly useful to derive photovoltaic power output from utility-scale photovoltaic power plants. For instance, it is possible to derive the shadows projected (i.e.,

¹ <https://reuniwatt.com/en/247-all-sky-observation-sky-insight/>.

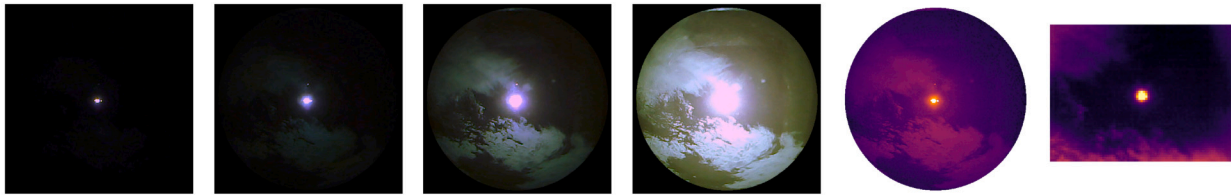


Fig. 9. From left to right, visible light images acquired at different exposure times (1, 4, 12, and 28 ms), HDR image resulting from fusing the visible light images, and infrared image of the circum-solar region. The resolution of the visible light and the HDR images is 450×450 , and the resolution of the infrared image is 80×60 .

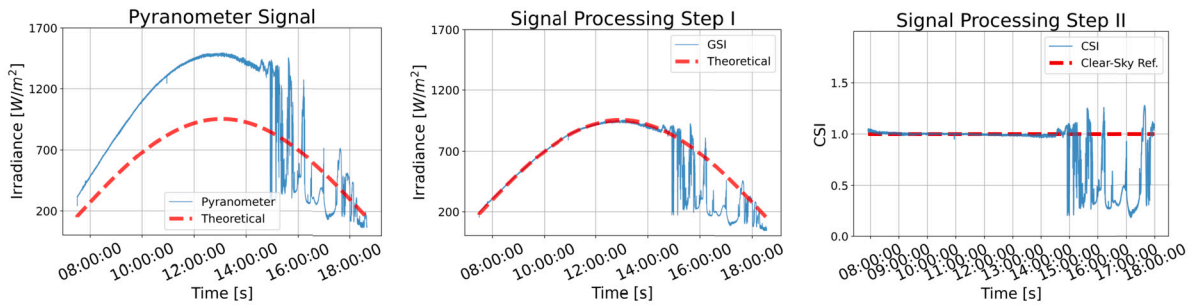


Fig. 10. Signal processing applied to the pyranometer measurements to obtain the CSI time series. From left to right, raw pyranometer measurements, after correcting amplitude and shifting bias, and CSI time series after detrending.

DNI map) by a cloud on the ground using multiple sky imagers [123]. The shadow camera is a visible light imager that records projected shadows instead of directly recording the clouds using sky images. The shadows are projected by passing clouds on the ground [124]. An alternative approach proposes visualizing passing clouds (in utility-scale power plants) by measuring the energy generated by photovoltaic modules and treating them as pixels. In this way, when a cloud crosses the photovoltaic modules, energy generation decreases, and an image of the shadow projected by a cloud is formed to track its motion [125]. Similarly, it is possible to scale this method in locations with multiple utility-scale photovoltaic plants [126]. To track the motion of the clouds across large space regions and use this information in solar forecasting, the energy generated at each photovoltaic plant can be defined as the pixels in an image [127].

2.3.2. Image reduction

The images acquired using visible light imagers (i.e., TSIs or ASIs) have a relatively high resolution for solar forecasting applications, generally ranging from 720×720 to 1280×1280 . The resolution in the input space (Height \times Width \times Number of channels) increases the number of parameters in Convolutional Neural Networks (CNNs) (i.e., number of convolutional filters and dimensions) and adds complexity to the model without showing any improvement in the performance. To accelerate model training, the original sky image is downsized using common image processing libraries. Consistent results show that the optimal resolution is $\geq 128 \times 128$ [36] and $< 256 \times 256$ [120] (Fig. 11). To have a competitive performance in solar forecasting applications, the smallest resolution possible is ≥ 64 [119]. When downsizing sky or satellite images, a low-pass filter should be applied first to avoid aliasing effects [13], which have been shown to alter image content if not handled properly [128].

2.3.3. Image transformations

The sky images acquired using a fish-eye lens or a reflective surface have a distortion produced by the concentration of light beams on the imager's sensor plane [130] (Fig. 12). The images can be undistorted by detecting the corners of a checkerboard to numerically approximate the affine transformation [131]. Additionally, further work in solar forecasting developed a method to transform the pixels' coordinate system (i.e., Euclidean geometry) to a longitude and latitude coordinate system (i.e., spherical geometry) [132].



Fig. 11. Sky image at different pixel resolutions: 512×512 , 128×128 , 32×32 . Source: SIRTA [129].

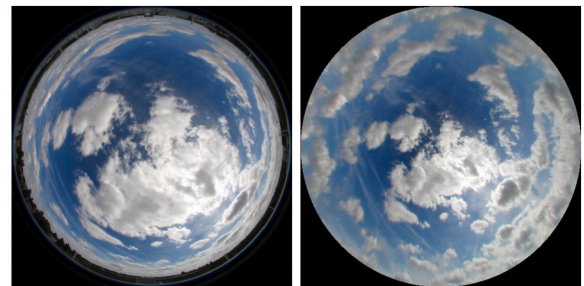


Fig. 12. Distorted image of the sky taken by an ASI and the corresponding undistorted image (azimuthal equidistant projection). Source: SIRTA [129].

Knowing the latitude and longitude of the air parcel corresponding to a pixel (i.e., voxel), the derivation of its area and volume is possible by applying a geospatial perspective reprojection [133]. A reprojection is particularly useful to approximate the cloud velocity vectors in sky images relative to the differences between consecutive sky images [134], which is a data processing (i.e., feature extraction) technique commonly applied in deep learning-based solar forecasting [103]; however, computing the absolute (i.e., actual) cloud velocity vectors is necessary to approximate the cloud height [135].

2.3.4. Ground-based image understanding

The circum-solar region has the information necessary to perform an accurate forecast when the forecast ranges from 30 seconds to 150 seconds [103]. When developing a multitask solar forecasting model ranging from 10 minutes to 60 minutes, it is necessary to use sky images with a large FOV [120]. Introducing knowledge based on image

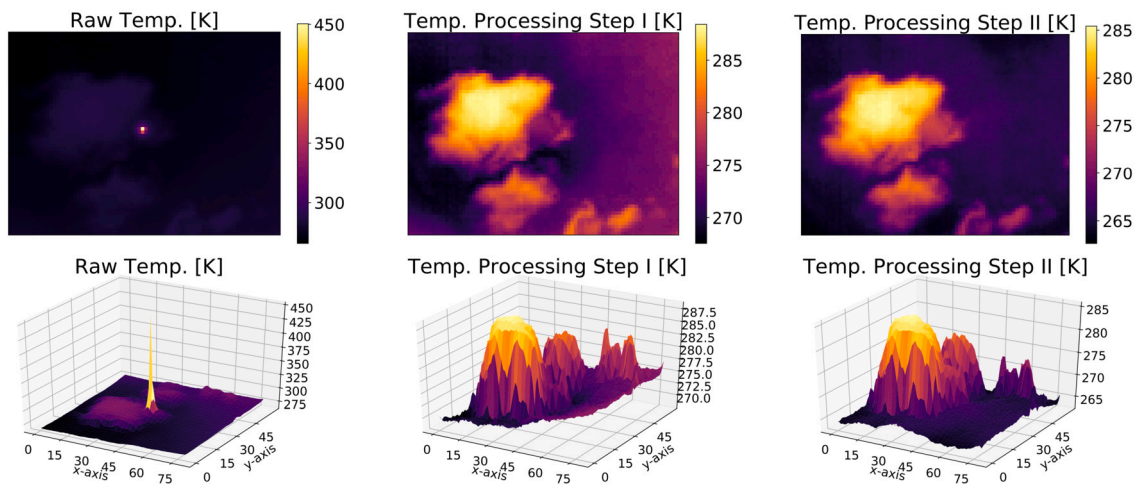


Fig. 13. Preprocessing applied to ground-based infrared images to remove radiation emitted by the sun (left), scattered radiation by particles in the atmosphere (middle), and radiation emitted by dust on the outside window of the sky imager (right). The bottom row shows the 3D graphs of the images in the top row.

understanding to the model allows for identifying the region of interest (ROI) in sky images.

Image understanding techniques to select potential ROI at different forecasting horizons are proposed in ground-based solar forecasting. Image transformation techniques discriminate between the circumsolar area (i.e., closeup of the sun) and the position of the sun in the sky (i.e., sun-centered) [136]. In particular, the ROI depends on the trajectory of the clouds and the sun (i.e., intersection). The *ladder* method was proposed to select an ROI in sky images by identifying which area in the ladder will cover the sun at a given forecasting horizon [137]. For that purpose, the *sun intercepting air parcel* method was introduced as an adaptive attention mechanism in which weights are the probability of a pixel in the image intersecting the sun at a particular forecasting horizon [121].

2.3.5. Satellite image understanding

In solar forecasting from satellite images, the position of the sun is key to defining the ROI because the part of the cloud cover affecting the generation of a photovoltaic system depends on the solar elevation. A proposed trigonometric method identifies the cloud cover ROI based on the angular position of the sun [70]. Additionally, an interpolation of this ROI is possible when the sampling frequency is too low. Another approach to detecting the ROI in satellite images assumes the Eulerian strategy in fluid mechanics as a simplification to the Lagrangian to identify the pixels (i.e., air parcels) that will intersect with the sun at a given forecasting horizon [121]. This method approximates the intersecting probability assuming a rectilinear trajectory [31]. An additional method for determining the ROI is to select the region around the photovoltaic system that is large enough to include all clouds that might move across the FOV [138,139,71]. This ensures that all existing clouds that might affect the radiation are included, other than those that form during the period of interest.

2.3.6. Detrending

Similar to detrending the GSI time series to obtain the CSI, it is possible to detrend the 2-dimensional time series of sky imagers to obtain the diffuse radiation emitted by clouds. Infrared sky images include the additive effect of direct solar radiation, diffuse radiation from the scattering effect produced by clouds and other water molecules, and the reflected radiation emitted by objects on the Earth's surface [95]. The direct radiation effect is stationary, and it can be removed. The diffuse radiation scattered by water molecules in the atmosphere produces cyclo-stationary effects that can be modeled using weather features (atmospheric pressure, relative humidity, and dew point). These weather features are available in standard weather stations [140]. The radiation

emitted by debris on the camera lens (i.e., sediments) is static, and it can be removed as well. The result is a far-infrared image that registers the radiation emitted by the clouds [141] (Fig. 13).

2.4. Feature extraction for physics-informed deep learning

Adding physical features about the atmosphere dynamics (i.e., exogenous variables) is proven to increase the performance of a solar forecast. Deep learning models do not have access to knowledge gained empirically about the nonlinear atmospheric dynamics and how weather features affect the underlying process. This section describes computer vision methods used to derive features from cloud cover observations and weather station measurements.

2.4.1. Cloud albedo

The cloud index (i.e., cloud albedo), derived from satellite images, is a feature that can potentially increase the performance of a satellite-based solar forecast. The cloud index is a relative measure of the opacity of clouds derived from the short wave band [142]. A different feature is the ground albedo, which measures the radiation reflected by the Earth's surfaces. The cloud index can be derived from the ground albedo using the spectral bands provided by the Rapid Scan High Rate SEVIRI imaging instrumentation in a MSG [13]. Additionally, the input tensor can include other information to increase the performance: the pixel's elevation map, longitude, and latitude [38]. The SEVIRI instrument also provides an optical cloud product that gives cloud optical thickness, pressure, phase, and effective radius [143]. The GOES satellites have a similar set of cloud data products available [144].

2.4.2. Cloud height

The cloud height can be directly measured using a ceilometer [145], but this sensor is expensive. Nevertheless, the derivation of the cloud height is possible by detecting the shadows projected by passing clouds at different positions using a cloud speed sensor [146] or a pyranometer grid [147]. Another alternative is to use multiple sky images and apply stereographic methods [123]. This approach has been validated against a ceilometer for solar forecasting applications and has performed accurate estimations [148]. A less exploited approach to approximate the cloud height is using temperature measurements acquired with a far-infrared sky imager [133,149]. The temperature of a cloud can be estimated knowing the rate of temperature decrease of an object floating in the atmosphere (i.e., atmospheric lapses [150]), and the atmospheric lapse can be approximated by the moist adiabatic lapse rate and the dry adiabatic lapse rate when a standard weather station is available [151].

2.4.3. Cloud thickness

The extraction of physical cloud features from ground-based images is challenging since not all sky-imagers provide interpretable measurements. However, recent research in the field proposed a method to build 3D images of clouds, which allows us to approximate the cloud thickness [152], and deep learning methods can extract information in sky images or satellite imagery from cloud optical thickness for cloud segmentation applications [153]. This is important to quantify the aerosol (e.g., clouds) effects in the solar radiation transmissivity and diffuse proportion to estimate solar energy generation [154]. In addition, aerosol optical thickness provides information about the scattering produced in the atmosphere (i.e., atmospheric extinction [155]). It is possible to derive this feature when using satellite infrared images [156] in combination with physical models [157]. Additionally, an alternative metric for visible light imagers to increase the performance of a satellite-based solar forecast is the cloud quality index which quantifies cloud thickness. This index is derived from a grayscale channel (i.e., brightness) and estimates the thickness of the cloud with respect to a set of neighboring pixels [74].

Aerosol optical thickness extraction and quantification methods with direct application in solar forecasting exist [158]. Most research proposed to assimilate data from satellite imagery [159], and it is proven influential in day-ahead solar forecasting based on weather research and forecasting software [160] or hybrid models [161]. However, a study found the effects of aerosols in short-term solar forecasting to be significant as well [162]. The capabilities of deep learning have not been explored yet in great detail for assimilating aerosol thickness in solar forecasts but approaches based on computer vision with deep learning increased in the performances when retrieving it from satellite imagery [163–165]. The product produced by these methods can add information to intra-hour solar forecasting applications that use ground-based imagers.

2.4.4. Cloud coverage

The cloud cover index refers to the fraction of the sky covered by clouds, and it is correlated with other cloud features discussed previously (e.g., cloud index). This feature is known to ameliorate solar forecasts when combined with other meteorological features [166]. This information is provided by standard weather stations [167] or can be directly extracted from ASI [168] or satellite imagery [169]. In fact, a recent study found that the derivation of the cloud cover from the distinct spectral bands of a satellite observation can be slightly different [170], pointing out that characterizing clouds may require fusing multispectral information. When the cloud height feature is known, deriving the cloud area from the cloud cover index is possible.

The extraction of the cloud coverage index requires the detection of pixels with clouds (see Section 3.3.4), and for that task, methods based on deep learning proposed to use spatiotemporal generative adversarial networks (GANs) and [171] or CNNs [172]. Solar forecasts based on deep learning can benefit from using cloud cover in the feature vector [173] and from using binary cloud cover images (i.e., cloud segmentation) in a CNN [174]. Other alternative approaches proposed to forecast the cloud cover to later derive solar generation based on GANs [175], and CNNs for nowcasting applications [176,130].

2.4.5. Cloud motion vectors

The cloud velocity vectors can be approximated using block matching to compute an error metric to find the most similar regions in two consecutive frames [25]. Block matching is computationally expensive, but the implementation of a smart search can potentially speed up the computation [177]. A similar approach to approximate the motion of objects (i.e., clouds) in two consecutive frames is particle image velocimetry (PIV) [178]. PIV computes the cross-correlation or normalized cross-correlation between the pixels in two images in the frequency domain [179]. The phase cross-correlation is an alternative to PIV used in images that is computationally less expensive [24].

The cloud velocity vectors in two consecutive images can also be approximated as the optical flow [180,181]. The Lucas-Kanade method [182] is a simple and efficient method to approximate the cloud velocity vectors [27], although the pyramidal Lucas-Kanade is most commonly applied [183]; however, solving optical flow equations via least-squares suffers the *aperture problem* [184]. The Horn-Schunck method [185] aims to solve the problem by applying a smoothness constraint. In addition, the variational implementation exists [186], and it is feasible for estimating the cloud velocity vectors [26]; however, the “total variation regularization” version of the Horn-Schunck method [187] was found to be the most efficient optical flow solution in cloud motion estimation applications [188]. Additionally, a formulation of weighted pyramidal Horn-Schunck for cloud motion vector estimation was recently proposed for satellite imagery applications [189].

Another alternative to approximate the cloud velocity vectors is the Farnebäck method [190]. This method can use prior knowledge, local weights, and be implemented in a pyramidal manner [191]. It has high performance when approximating the cloud velocity vectors in ground-based sky images [192], but it performs better with satellite images [193,194].

2.4.6. Cloud dynamics

The cloud velocity vectors allow for the derivation of second-order dynamics, such as vorticity and divergence [195]. The vorticity measures the angular velocity of a fluid (i.e., a cloud is rotating). The divergence measures the expansion rate (i.e., negative divergence means a cloud is dissipating and positive forming). Other methods available in computational fluid dynamics literature are still unexplored in solar forecasting applications (e.g., dynamic mode decomposition [196] and proper orthogonal decomposition [197]).

2.5. Open-source data

This section presents a range of publicly available datasets that have been or could be used in computer vision-based solar power modeling. To some extent, evaluating modeling approaches on such datasets offer the opportunity to compare their performance relative to already existing methods.

2.5.1. Ground-based sky images

Different research groups contributed to improving the accessibility of ground-based imagery datasets for solar forecasting and cloud segmentation [34]. In particular, some groups have contributed in extended efforts to collect data from multiple sensors and sites (see Fig. 14). Other groups have developed software to access publicly available datasets [34], such as *SolarData* [198] and *OpenSolar* [199].

- The Solar Radiation Research Laboratory at the National Renewable Energy Laboratory (NREL) developed a baseline measurement system [200]. This system provides multiple years of 10-minute sky images from a TSI (1,920×1,920) since July 2004 and from an ASI (288×352) since September 2017, plus 1-minute resolution GSI and meteorological measurements in Golden, Colorado, USA.
- The SIRT Atmospheric Research Observatory released sky images collected by a TSI (480×640) and a short-exposure ASI (768×1024) during 7 and 8 years, respectively, with a resolution of 1 to 2 minutes, and 1-minute resolution GSI and meteorological measurements from Palaiseau, France [201].
- The UCSD dataset for solar forecasting models combines multiple data sources for Folsom, California, USA [202]. The dataset includes 3 years of high-resolution sky images (1536×1536), GSI measurements in 1-minute resolution, satellite imagery, NWP, and secondary data (i.e., features extracted from images and GSI time series).

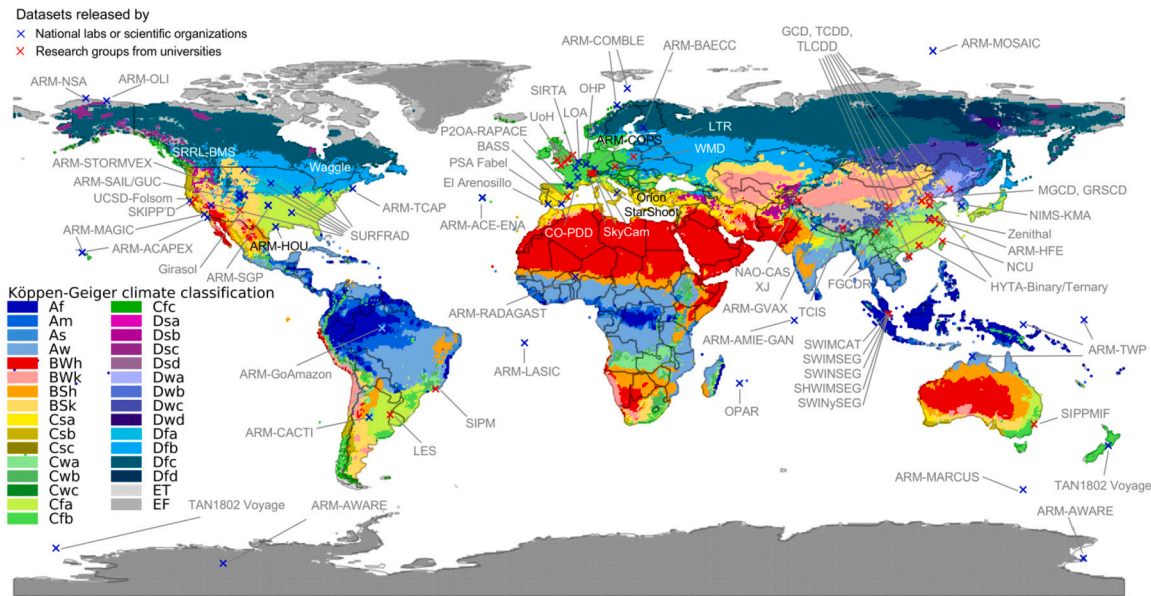


Fig. 14. Köppen-Geiger climate map with geographic locations of open-source sky image datasets annotated by blue (national labs or scientific organizations) and red (universities) crosses. This map highlights the diversity of open-source sky image datasets. Taken from [34].

- *SkyCam* has 1 year of HDR sky images (600×600) acquired using an ASI, paired with GSI measurements (logged every 10 seconds) from three locations in Switzerland (Neuchâtel, Bern, Alpnach) [203].
- *Girasol* provides sun-centered large FOV HDR multi-exposure visible (450×450) and small FOV radiometric infrared (80×60) sky images with 15-second resolution during 244 days. The dataset includes GSI and meteorological measurements (logged every 6 seconds) from Albuquerque, New Mexico, USA [85].
- *SKIPP'D* contains 3 years of ASI and photovoltaic power output data with a 1-minute resolution from Stanford, California, USA [204]. It provides both a processed benchmark dataset that contains pairs of down-sized sky images (64×64) and photovoltaic power output ready to use for computer vision-based solar forecasts as well as raw data containing sky videos, high-resolution sky images (2048×2048), and photovoltaic power output.

There are cloud segmentation datasets composed of ground-based sky images. For instance, *SURFRAD* provides hourly sky images (288×352) from 10 different locations in the USA during 2–10 years (two locations have 30-second sky images) [205]. In addition, there are multiple datasets collected over a year at a 1-minute resolution with cloud labels and corresponding sky images acquired by ASIs located in Singapore. *SWIMCAT* has sky images (125×125) acquired during daytime with 5 cloud classes [206], *SWIMSEG* is a sky image (600×600) dataset with 2 cloud classes acquired during nighttime [207], *SWINYSEG* has a Nychthemeron sky images (300×300) dataset with 2 cloud classes [208], and *SHWIMSEG* is a 2-class HDR sky image (500×500) dataset acquired during nighttime [208].

2.5.2. Irradiance maps

GSI and DNI maps are publicly available for researchers and professionals. These maps provide GSI measurements at the ground level derived from satellite imagery. This information is useful for developing solar forecasting applications based on computer vision and remote-sensing. The publicly accessible databases are described here. The data are accessible through the websites or automatic API requests.

- *Helioclim* is maintained by MINES ParisTech, ARMINES, and TRANSVALOR S.A. [209]. It provides real-time data with 15-minute-ahead resolution measurement maps of GSI, DNI, diffuse, and reflected radiation at normal, tilted, and horizontal planes.

The data are preprocessed images acquired from MSG-11. *Helioclim* contains historic data from 2004, and the maps have a spatial resolution that varies from 3 to 12 km.

- *ERA5* provides estimates of a large number of atmospheric, land, and oceanic climate variables, including direct and diffuse irradiance, to generate weather feature maps [210]. It is maintained by the European Centre for Medium-Range Weather Forecasts (ECMWF). It has a temporal resolution of 1 hour, with historic records from 1950. This dataset has a spatial resolution of 30 km and extends from 137 levels from the surface to a height of 80 km. The data have a delay and are generally available within 5 days of the real-time acquisition.
- A dataset equivalent to *ERA5* is the *MERRA-2* reanalysis climate database managed by the National Aeronautics and Space Administration [211]. This dataset contains measurement maps for research applications from 1980 with a spatial resolution of 50 km.
- The *National solar radiation database (NSRDB)*, managed by NREL, is a GSI map database with a temporal resolution from 5 to 60 minutes [212]. The database contains historic measurements from 1998. The spatial resolution varies from 2 to 4 km. The data were acquired with Himawari, GOES, and MSG satellites.

2.5.3. Satellite imagery

Satellite imagery is also publicly available for end-users and researchers. These satellites provide measurements in multiple spectral channels, different spatial resolutions, and at various temporal resolutions as well. Geostationary satellites cover very large areas of the world constantly but at lower spatial resolution than lower-orbiting satellites such as Landsat. This constant imaging enables geostationary satellite imagery to be useful for short-term forecasting. Non-geostationary satellite imagery is useful for generating high-resolution irradiance maps or for detecting solar sites, but has limited usefulness for short-term forecasting of solar generation because of the long revisit times. In practice, most machine learning methods applied to satellite imagery are essentially interoperable with marginal adaptation to accommodate for different image formats (e.g., number of channels).

- *NOAA* provides satellite imagery covering the continents of North America and South America from 2016 with GOES-16, 17, and 18 satellites equipped 16 spectral channels. The resolutions vary from 0.5 to 2 km spatially, and from 5 to 15 minutes temporally [213].

- The *United States Geological Survey* offers access from multiple satellites, including the Landsat series of satellites. Landsat satellites image in up to 11 spectral bands with between 15-100 meter resolution and cover the entire world, with a revisit time of 8 days [214].
- The *European organization for the exploitation of meteorological satellites (EUMETSAT)*, which is the European operational satellite agency for monitoring weather, climate, and the environment from space, provides both real-time satellite and historic imagery from 1981 [215]. The images are acquired from MSG, Metop, and Sentinel-3 satellites. The satellite images are available over Europe [216] and over the Indian Ocean [217] from 2004 and 2017, respectively. The satellites have 12 spectral channels; one channel has 1-km resolution, and the other 11 have 3-km resolution. The temporal resolution varies from 5 to 15 minutes.
- The *JMA* provides satellite imagery from the Indian Ocean to the mid-Pacific. The images are acquired by Himawari satellite equipped with 16 spectral channels, and it is available from 2015. The temporal resolution ranges from 30 seconds to 10 minutes, and the spatial resolution varies from 0.5 to 2 km [218].

3. Solar power modeling with computer vision

Computer vision-based solar forecasting aims to predict the future solar power output at a location of interest using computer vision to analyze observations of the cloud cover, which accounts for most of the stochastic spatiotemporal solar variability (Fig. 7). Both imagery data (e.g., sky and satellite images) and sensor measurements (e.g., GSI, photovoltaic power output, wind speed, wind direction, temperature, pressure) need to be collected. Researchers can either set up devices to collect their own data (see Section 2) or acquire data from open-source datasets (see Section 2.5). Following this, quality control (Section 4.2.4) and feature engineering methods are applied. After preprocessing, the data are partitioned into training, validation, and testing sets (Section 3.5) for architecture development based on diverse metrics. The model performance is monitored after deployment, and adjustments on the model might be made. Further, online learning might be implemented to incorporate newly collected data and update the model [219].

3.1. Spatial characteristics

The spatial attribute of solar forecasting is crucial for many applications in terms of field of view and resolution. The complementary spatial characteristics of imaging techniques based on sky cameras and satellites facilitate the applicability of modeling techniques to an intermediate spatial extent via information fusion [13].

3.1.1. Sky images

Sky image-based models can offer high-spatial-resolution irradiance maps covering several square kilometers around the camera. In some conditions, local dense irradiance maps can be determined [220], but not all sites offer this possibility. For this reason, the majority of solar forecasting studies based on sky images currently focus on singular value prediction (e.g., GHI or GSI, DNI, photovoltaic power output) [35,119], image prediction [221,131], or cloud mask prediction [222,40].

3.1.2. Satellite imagery

The current geosynchronous weather satellites for cloud cover observations have, at best, an on-the-ground pixel size of 0.5 km – 2 km at the equator, which increases up to 4 km – 6 km per pixel at higher latitudes. Additionally, the pixels do not cover a uniform ground footprint, with pixels closer to the edge of the Earth's disk being more distorted than those near the center of the satellite's FOV. These characteristics mean

that meteorological geostationary satellites do not have the spatial resolution to detect clouds smaller than a few kilometers in size, and how the clouds appear in the image changes depending on where in the image the clouds are located. Nearer the Earth's disk, clouds appear offset from the part of the Earth's surface they cover, whereas near the center of the satellite's FOV, they appear directly over the part of Earth's surface they cover. The region of the cloud cover blocking a solar site can be determined from the angular position of the sun as a preprocessing step [70]. In practice, satellite-based spatial forecasts are more flexible than ASI approaches because irradiance maps at a pixel level can be derived from radiative transfer models or with deep learning from ground-level measurements [75]. For this reason, the types of prediction from satellite observations include *in situ* measurements [40,223] but also cloud-index maps [13,224] and irradiance maps [38].

3.2. Temporal characteristics

Solar energy modeling can be characterized by various temporal variables: temporal resolution of *in situ* measurements or cloud cover observations, spectral resolution, past context, lead time, forecast horizon, and update rate (Fig. 15). Except for the spectral information, these attributes are primarily associated with time-dependent tasks or applications, including solar forecasting or video prediction, as opposed to static tasks, such as image segmentation.

3.2.1. Temporality in solar energy meteorology

Most temporal characteristics of a computer vision task can be split into two groups, past and future, which are optimized or constrained differently. On one hand, future temporal parameters are often set by the application. For instance, in a hybrid power plant optimization scheme, the horizon of the forecast is set by the warm-up time of the fossil fuel backup [225,226]. Further, an algorithm designed for energy trading must comply with market regulations [227] (Section 3.2.2).

In practice, the forecast horizon of sky images is also constrained by the FOV (up to 20/30 minutes, depending on the local cloud speed), whereas satellite observations can infer longer-term forecasts up to several hours. On the other hand, past temporal aspects are partially constrained by the source of data (e.g., fixed temporal resolution of satellite observations ranging from 5 min – 15 min for MSG to 2.5 min – 10 min for MTG), but they can also be tuned to a specific application to maximize the performance of the model [119]. Common hyperparameters defining the past context are the number of lag terms and the temporal resolution. Moreover, for a given computational capacity in operating conditions (e.g., edge computing), a trade-off between spatial and temporal resolutions of sky images can be reached.

Another key temporal aspect of any imaging technology is its spectral information. By capturing signal related to specific wavelengths, sky cameras and satellites can infer various characteristics of clouds, such as their temperature or composition. This facilitates cloud height estimation [133], sun localization [97], optical depth modeling [228] and cloud cover segmentation [229,230].

The diverse time attributes of sensing techniques (e.g., update frequency, spectral bands) represent a significant opportunity for richer detection capabilities (e.g., high temporal resolution, infrared sky cameras, multispectral imaging). However, this heterogeneity challenges the integration of multiple sources of information into a single modeling approach [13]. In that context, the flexibility of sky cameras in terms of temporal resolution can be exploited to match the update frequency of satellite imagery.

3.2.2. Operational solar forecasting

In an operational context, solar forecasting is defined by four time-related terms [231], the lead time, the horizon, the update rate, and the resolution, as shown in Fig. 15. The “boundary” between the short-term and very short-term solar forecasting is vague, due to the multidisciplinary nature of the problem and the involvement of multiple

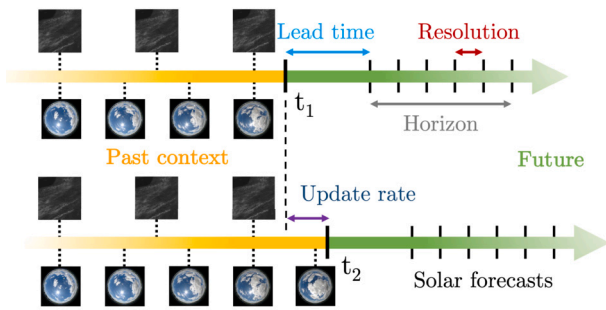


Fig. 15. Operational solar forecasting.

communities. First, different communities define the lead time differently. For example, the World Meteorological Organization refers to very short-term forecasting as forecasts 0-12 h ahead [232]. However, power system operators use day-ahead, intra-day, hour-ahead, or even intra-hour solar forecasts in their operations [233]. Second, even within the same community, the definition can also differ. For example, the energy forecasters differentiate short-term and very short-term solar forecasting by the availability of the numerical weather predictions (NWP) [234] that have various lead times from different models. In power systems, operators define the requirements for forecasts according to their own market, which are usually different from one to another. Third, there are also gaps between the literature and real world practice, since the processing and communication time is usually not considered in scientific publications but is important in power system daily operations.

3.3. Computer vision tasks in solar forecasting

Modern computer vision offers a wide range of solutions to various real-world challenges, from object detection and tracking, to modeling physical phenomena and generating synthetic data. Many studies in solar energy have demonstrated the applicability of vision algorithms to tasks, such as solar panel localization from remote imagery [235,236] or solar cell defect automatic detection [237,238]. Regarding solar forecasting, computer vision is key to modeling the complexity of the cloud cover spatiotemporal dynamics. From a general integrated learning task, such as predicting the future solar output from past sky observations, computer vision-based solar forecasting can also be split into several subtasks addressable by computer vision, such as cloud classification or solar estimation. We describe some of these tasks in the next subsections.

3.3.1. Solar estimation and radiative transfer modeling

Solar estimation is essentially building a mapping, f , from the image, I_t (captured at time t), to the concurrent sensor measurement, \mathcal{M}_t , either GSI or photovoltaic power output, i.e., $f : I_t \mapsto \mathcal{M}_t$, which is similar to the computer vision task of estimating the age of people based on their facial images (Fig. 16). In situ measurements of irradiance and power output of a photovoltaic system are usually costly in devices and human labor for regular maintenance of the devices. Solar estimation models, in contrast, only require imagery data for training at the beginning, and they can then be applied with minimum cost; therefore, they can be used as an alternative to in situ measurements.

Depending on the spatial coverage, different imagery data are required for developing the solar estimation models. As shown in Fig. 2, ground-based sky images are suitable for a spatial extent below 1000 m, whereas satellite imagery is more useful for larger spatial coverages above 1000 m.

Numerous studies have examined the effectiveness of using end-to-end deep learning models, mostly CNN, for solar estimation [239]. In 2018, Sun et al. [36] developed for the first time a CNN model for solar estimation, called SUNSET, by using ground-based sky images, and

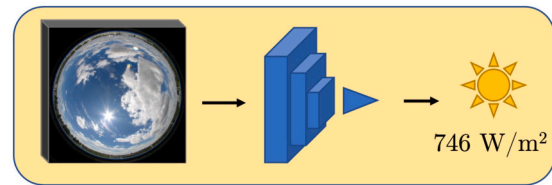


Fig. 16. Deep learning model for solar estimation (radiative transfer modeling) from cloud cover observation.

the model shows promising performance in correlating sky images to photovoltaic power output. The estimated value almost perfectly aligns with the ground truth in clear-sky days, whereas it struggles a bit with cloudy days. To deal with the challenge of solar estimation in cloudy days, Jiang et al. [240] built a CNN model for GSI estimation based on only the cloudy-day data, which achieved a relative RMSE of 8.7% on the test set. Nie et al. [241] developed a classification-prediction framework that first classifies input sky images into different sky conditions, before the classified images are sent to specific, fine-tuned CNNs for concurrent photovoltaic power output estimation.

For satellite imagery, one of the main non-deep learning approaches are the HelioSat methods [242-245], which map from geosynchronous satellite imagery to GSI using physics-based models. In recent years, new satellite-based retrieval methods using machine learning/deep learning have been used to directly infer solar irradiance from satellite images. Jiang et al. [246] investigated the spatial-scale effects on the accuracy of surface solar radiation retrieval using multivariate linear regression, artificial neural network and CNN models and found that scale effects could significantly influence the accuracy of satellite-based solar radiation estimation. In another study, Verbois et al. [247] examined the practical efficacy of using machine learning methods to directly infer the surface solar irradiance from satellite images compared to state-of-the-art physical retrieval methods. A multilayer perceptron model was trained using various spatial and temporal configurations of the development and evaluation datasets comprising radiance levels and colocated ground measurements. The authors concluded that the performance of the data-driven model is very dependent on the training dataset. The estimation of photovoltaic system output from satellite images can be performed via solar panel segmentation methods to infer the local solar capacity [248].

The seasonality of GSI, both over the course of the year and diurnally, can cause issues for solar forecasting models, so generally they should be taken out before giving the data to the model via the CSI [10]. One of the best methods for removing the seasonality is to use a clear-sky model [249-251]. This is especially important for satellite-based modeling approaches because sky image-based forecasting models can learn to correlate the CSI with the position of the sun in the sky.

3.3.2. Solar forecasting and nowcasting

In the past five years, computer vision-based solar forecasting has seen an increase in using end-to-end deep learning models to correlate the future irradiance level or photovoltaic power output with the sequence of past cloud cover observation (Fig. 17). A typical setup is to predict the T -minute-ahead future GSI or photovoltaic power output, referred to as measurement \mathcal{M} , based on an image sequence, I , covering the past H minutes, and possibly together with concurrent historical measurements. Mathematically, it can be expressed as learning a mapping $f : (I, \mathcal{M})_{t-H:\delta:t} \mapsto \mathcal{M}_{t+T}$, where δ is the interval between two lag terms.

CNNs are commonly used in these data-driven frameworks, either solely [119,120,252,253] or as a backbone, generally hybridized with recurrent neural network architectures, such as long short-term memory networks (LSTM) [101,39,40]. For example, Sun et al. [119] modified the CNN model architecture they developed for the solar estimation task [36] to solar forecasting by injecting sky images in the past 15

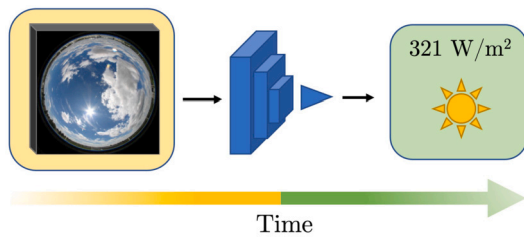


Fig. 17. Solar power forecasting with deep learning.

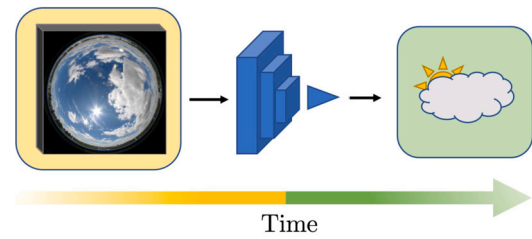


Fig. 18. Deep learning approach for event prediction.

minutes together with the concurrent photovoltaic power output history to predict photovoltaic power output 15 minutes ahead into the future. The study extensively explored the optimal model input and output configurations. The results suggest that using the same length of history as the forecasting horizon, i.e., $T = H$, is a good heuristic, and using hybrid input, i.e., images plus the photovoltaic power output history, could achieve better performance than using solely images or the photovoltaic power output history. Paletta et al. [39] established a benchmark for short-term GSI forecasting by comparing four commonly used deep learning architectures — CNN, LSTM, 3D-CNN and ConvLSTM — with different forecasting horizons from 2 to 30 minutes. The results show that these deep learning models can outperform by approximately 20% the reference persistence model for a 10-minute forecasting horizon; however, these models have similar limitations in anticipating fundamental events, causing large irradiance changes, such as clouds obscuring the sun. Furthermore, these four neural networks are shown to provide comparable predictions and forecast errors, with the best models marginally improving over the standard CNN architecture.

For solar nowcasting with satellite imagery, the task involves determining the amount of solar power being generated at the current time and a few hours into the future, corresponding to some amount of time steps into the future, for example, 48 time steps for a 4-hour forecast at 5-minute resolution. The task is generally framed as a regression task, trying to predict the exact amount of solar energy within the area or site of interest, and is an extension of the previous section on future times. Because satellite images are simply multichannel images, common deep learning models — such as ResNets, ConvLSTMs, and U-Nets — can and have been applied to this problem [254,73,255]. Yeom et al. [256] used a ConvLSTM model to predict short-term changes in solar radiation over South Korea, validated by ground pyranometers. In [257], satellite and NWP data were combined to forecast the future irradiance. Both ConvLSTM and 3D convolutional models were investigated for forecasting GSI up to 6 hours ahead in Italy in [258], with the ConvLSTM winning out over the smart persistence and the 3D convolutional model. The input to these models is usually a stack of satellite images covering the area surrounding the region of interest (e.g., a solar power plant), potentially much larger than the area of interest to better capture clouds that might move over the area of interest within the forecast horizon [138,139]. To work with models not designed for time series, such as U-Nets and ResNets, one common approach is to combine the time and channel dimensions of the inputs [255].

Predicting irradiance is different than predicting photovoltaic power output. Irradiance predictions are generally dense predictions for every pixel in the input images, predicting the irradiance at those points, whereas photovoltaic power output forecasting tends to be more sparse, predicting the actual power output for specific sites in the FOV. Photovoltaic forecasting is also quite directly linked to the specificities of the photovoltaic panels (e.g., size, orientation, soiling) and site-specific data (e.g., shadowing effects), rather than irradiance predictions, which can be adapted to different types of panels and site-specific data after the fact.

3.3.3. Event prediction

Passing clouds obscure (or show) the sun and significantly attenuate (or increase) the irradiance reaching the ground in the short term, consequently leading to a sudden drop (rise) of the power output of photovoltaic solar panels. These significant power output changes of local photovoltaic systems within a short time period are often referred to as ramp events, which are crucial information for grid operation and management. The cloud dynamics can be captured by ground-based sky imagers or satellites; however, existing deep learning models used in solar forecasting are usually lacking in their ability to anticipate these events, evidenced by recurrent temporal lags in predictions [39].

Several studies have tried to address these issues by either directly predicting the ramp events or predicting the future states of the sky using sky images as a way of assisting solar forecasting (Fig. 18). Pothineni et al. [259] used CNNs to predict the 5-minute-ahead sky condition to be either clear or occluded. The work experimented with different neural network architectures and concluded that an 18-layer ResNet performed the best, with 91%–93% accuracy, though no follow-up experiments further showed the usage of this information for solar forecasting. Abuella and Chowdhury [260] developed a post-processing adjusting approach to improve the ramp event forecasting capability of hour-ahead ensemble solar forecasting models. The ramp event forecasting can be formulated as a classification problem, where four types of ramp events are defined based on the direction and rate of the ramp: ramp-up high rate, ramp-up low rate, ramp-down high rate, and ramp-down low rate. The results demonstrate the efficacy of this adjusting approach. Leelarui and Teerakawanich [261] combined image processing techniques with a CNN based on ResNet for alerting of trigger events before the sun cover happens, 1 to 2 minutes in advance. The results show a prediction error of 27.50%. On the other hand, Paletta et al. [40] developed a deep learning model called ECLIPSE, which consists of spatial and temporal encoders to extract features from the past sky image sequence to predict the future segmentation map (i.e., as a means of indicating whether the sun is obscured by the clouds) and the future irradiance at the same time. Besides minimizing the irradiance prediction error, a regularization term associated with the segmentation map prediction is included in the loss function of the deep learning model.

Commonly used evaluation metrics — such as MAE, mean square error (MSE), and RMSE — are not able to capture the fluctuations of irradiance or power output time series, and they are not suitable metrics for assessing the model performance on detecting ramp events. To bridge this gap, Vallance et al. [262] proposed a new ramp metric, called the ramp score, which compares the ramps extracted from the forecast with those extracted from the measurements (details can be found in Section 3.5 and in [39]).

3.3.4. Image segmentation

Cloud cover observations are widely used to perform accurate segmentation maps. Such accurate detection of clouds assists remote-sensing analysts and meteorologists in performing automatic computations of cloud cover values. Traditionally, the cloud cover values are reported via *oktas*; however, with the introduction of sky cameras, it is now possible to perform automatic and accurate segmentation of clouds, and thereby report the cloud cover value. Researchers in the field have used various statistical image processing techniques over

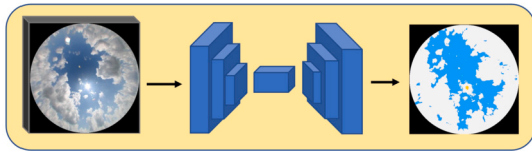


Fig. 19. Encoder-decoder neural network architecture for cloud cover segmentation.

the last decade in the segmentation of sky/cloud images, including superpixels [263], threshold methods [263], graph cuts [264,265], as well as probabilistic approaches [266]. Apart from classic statistical methods, supervised [149] and unsupervised [267] fully probabilistic approaches and kernel learning methods were used for cloud segmentation in ground-based infrared images [268]. In addition, unsupervised fully probabilistic methods enable cloud semantic segmentation [141].

Recently, deep learning techniques have also been used to improve the segmentation accuracy of the various approaches (Fig. 19). Researchers in [208] proposed a lightweight deep learning architecture, called CloudSegNet, for efficient segmentation of sky/cloud images. It is the first to accomplish state-of-the-art results on open datasets by combining daylight and nighttime (also known as nychthemeron) image segmentation in a single framework. Hasenbalg et al. [67] compared six benchmarking cloud segmentation algorithms on images captured by a standard security camera: a color-channel threshold-based algorithm; a Clear Sky Library based approach, a region-increasing algorithm, the Hybrid Thresholding Algorithm (HYTA), and an improved HYTA+ development. Additionally, a CNN is modified for this issue via transfer learning. Further, Xie et al. [269] proposed SegCloud, a new deep CNN model for precise cloud segmentation based on ground-based observation. The SegCloud architecture has a symmetric encoder-decoder structure in terms of architecture. While the decoder network restores the high-level cloud feature maps produced by the encoder network to the same resolution as the input images, the encoder network combines low-level cloud features to create high-level, low-resolution cloud feature maps. The SegCloud approach can automatically partition whole-sky images captured by an ASI and has a powerful cloud-discriminating capacity. Transformer-based cloud segmentation approaches are also recently explored in [270].

In satellite imagery, cloud segmentation is used to compute the cloud cover and to identify pixels that are covered by clouds [271]. Pugazhenthil and Kumar [272] proposed an improved fuzzy c-means and an improved k-means clustering algorithms that segment far-infrared images of the INSAT-3D satellite into regions with low-, middle-, and high-level clouds as well as areas without clouds. Further, Francis et al. [273] examined the issue of cloud detection in visible and multispectral satellite data from two high-resolution sensors, Carbonite-2 and Landsat 8. In addition to binary segmentation, Wieland et al. [274] provided a multiclass, data-driven method for semantic segmentation of clouds and cloud shadows in single-date images based on a modified U-Net CNN. On a large-scale library of Landsat images, they trained the network to classify objects into five categories: shadow, cloud, water, land, and snow/ice. These algorithms are often modified to run on onboard edge devices. Bahl et al. [275] proposed low-power neural networks for satellite image semantic segmentation [275]. These networks are designed to function on edge devices with limited resources, such as phones, drones, or satellites. Other types of computer vision-based detection tasks consist of identifying rooftop solar panels via remote-sensing imagery and estimating some of their main characteristics (e.g., capacity, tilt angle, azimuth) for modeling purposes [276–279].

In short, Table 3 provides a summary of the various available deep networks for the purpose of cloud image segmentation.

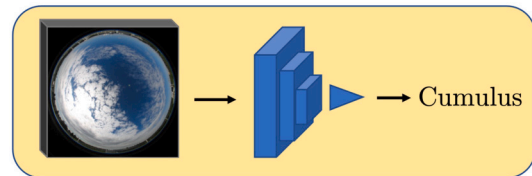


Fig. 20. Deep learning architecture for cloud classification.

3.3.5. Cloud classification

In more recent times, deep learning classifiers have been trained to produce superior outcomes without any need for feature identification prior to training (Fig. 20). A CNN was implemented to extract features that are then processed and sent to a support vector machine classifier for classification [281]. The network's average accuracy was evaluated across the six classes and claimed an accuracy of 89.5% on its dataset. Later, to categorize cloud images, Zhang et al. [282] developed AlexNet [283] and published the CloudNet architecture. The author-released the cirrus cumulus stratus nimbus dataset and the SWIMCAT dataset were both used to test the model. Recently, Wang et al. [284] introduced CloudA, in which they trained a unique deep CNN using a support vector machine classifier.

In summary, the task of cloud classification has been attempted by researchers using both machine-learning and deep-learning methods. Some of these methods offer good accuracy in a handful of annotated cloud classification datasets; however, we feel that there is still a lot of scope for improvement in this area, primarily in the accuracy of the cloud classification task and the availability of open-access cloud classification annotated datasets.

3.3.6. Cloud motion prediction

In video prediction, the objective is to predict the next set of frames into the future given the last N frames as history (Fig. 21). In relation to solar forecasting, the main application of video prediction is to predict where clouds will move in the future and therefore how clouds visible at the inference time will affect the output from solar energy systems.

Defining what a good video prediction is remains difficult. With metrics such as RMSE, and MAE to a lesser extent, deep learning models tend to blur details, causing clouds to become more of a white fog as the forecast time lengthens [13,38,221,285,286] (Fig. 22). Although this blurriness is rarely representative of the actual future scenario, it can be beneficial to some operational applications because it describes the uncertainty associated with predictions. A resulting drawback is the low variance of the forecasts, which mitigates steep solar ramps [39]. To force more realistic outputs, deep learning models such as GANs have been shown to create much more realistic forecasts of precipitation [287] or cloud displacement [288] compared to non-generative models trained with RMSE or related metrics. Alternatively, structure-preserving loss functions have been applied to produce more realistic predictions by also measuring nonstructural distortions, such as luminance and contrast changes [224].

In relation to satellite-image video-based prediction, each frame is taken between 5 and 15 minutes apart and covers extended areas. This large FOV means that satellite-based methods can see clouds that are hundreds of kilometers away from the site of interest, potentially allowing for more accurate long-term forecasts of cloud movement than sky images, with their more limited FOV. Current works build neural networks trained on satellite observations to predict the future cloud cover. These sequence-to-sequence architectures take raw and preprocessed images (e.g., irradiance or cloud-index maps) as well as auxiliary data encoded in additional channels (e.g., irradiance values, time spatial coordinates) as input [13,38].

Video prediction with sky images exploits similar approaches: predicting future sky observations from a context of several past frames (Fig. 23). In practice, sky cameras offer much higher spatial and temporal resolutions of the local cloud cover, revealing clouds that cannot

Table 3
Summary of various deep learning networks for cloud image segmentation.

Network	Image Type	Advantages	Disadvantages
DeepLabV3+ [280]	Ground-based	High accuracy	Computationally expensive
CloudSegNet [208]	Ground-based	Lightweight architecture, efficient segmentation	Limited performance on complex scenes
SegCloud [269]	Ground-based	Precise cloud segmentation	Symmetric encoder-decoder structure limits scalability
Transformer-based models [270]	Ground-based	Captures long-range dependencies	Higher training complexity
Pugazhenthil and Kumar [272]	Satellite	Multi-label segmentation	Dependent on thresholds
Francis et al. [273]	Satellite	Works on multispectral data	Low performance on snow or icy terrain
Wieland et al. [274]	Satellite	Five-class classification	Variance in performances across sensors and locations
Bahl et al. [275]	Satellite	Designed to function on limited resources	Does not work in multi-class image segmentation

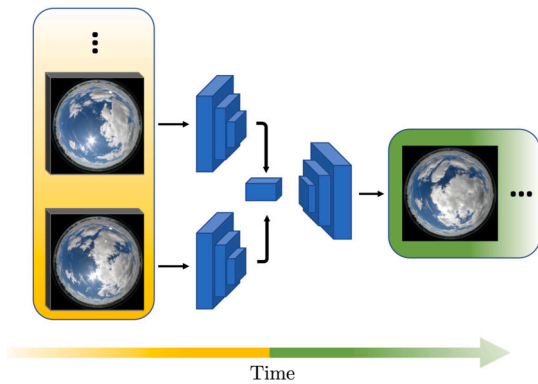


Fig. 21. Deep learning network for cloud motion prediction.

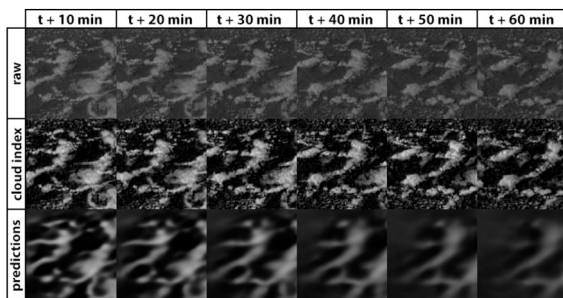


Fig. 22. Prediction of future satellite observations by a deep learning model. The longer the horizon, the fuzzier the predictions from the model. Taken from [13].

be observed from satellites; however, the FOV of a unique camera remains limited to several square kilometers centered on the device. This limits the forecast horizon of methods based on a single sky camera to the amount of time taken by clouds to hide the sun from the moment they were visible in the image. Typically, clouds cross the FOV within 20 to 30 minutes, depending on the wind speed. For this reason, deep learning models usually observe a 10-minute context window to predict future frames from 5 to 20 minutes in the future.

A range of deep learning architectures have been applied to video prediction from sky images. Siddiqui and Bharadwaj [222] learns semantic masks of both past and future frames using ResNet to generate a representation tensor for each context frame. Future masks and corresponding irradiance levels are obtained from spatial attention mechanisms combined with ConvLSTM layers. To address the fuzziness of sky image predictions, a method based on a GAN was proposed by [288]. A generator is trained to predict the next frame from a context of four sky images while the discriminator tries to distinguish real sequences of five consecutive real frames from sequences resulting from the generating process (i.e., four real frames plus the corresponding prediction of the

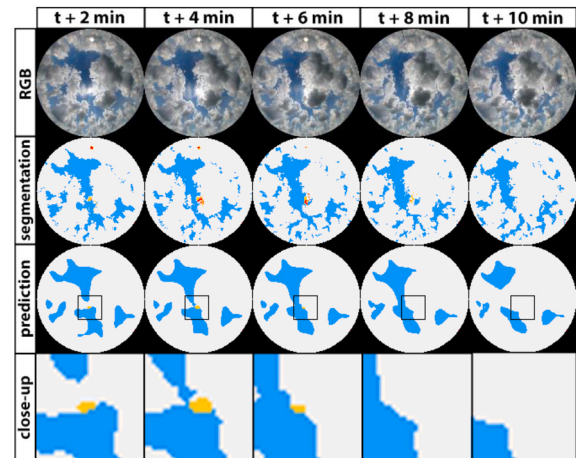


Fig. 23. Cloud motion prediction with sky images. A close-up of the circumsolar area shows how the model predicts that the sun is going to be hidden by a cloud within the next 10 minutes. Taken from [40].

generator). Observing that physical phenomena such as cloud displacement can be described by advection-diffusion equations, another model, called PhyDNet, aims to constrain predictions with partial differential equations in the latent space [221]. PredNet [289], which uses the error associated with previous generated images to improve the subsequent frame prediction, was also used to predict irradiance from the learned representations of future sky images [285]. Another method iteratively predicts cloud masks frame by frame with a CNN [290]. A deep learning model, named ECLIPSE, generates a spatiotemporal representation of the past sequence of sky or satellite images using 2D and 3D convolutional layers [40]. A recurrent convolutional module iteratively predicts future states that can then be decoded into various signals: irradiance level, irradiance distribution, or cloud mask. Because fish-eye lenses induce a strong distortion, some approaches first unwrap sky images as a preprocessing step prior to feeding a deep learning model [40,131]. This improves the shape consistency of clouds as they move across the FOV.

3.4. Open-source code and models

Increasingly, solar forecasting models are being made open source, either only the code or the code and the trained model weights. These releases allow for anyone to easily build upon the existing work, better compare approaches, reproduce the original results, or improve the open code. Platforms for machine learning model hosting, such as Hugging Face, are also sometimes used to publish trained models and reduce the barriers for others to build upon and use the models.

For satellite-based solar forecasting, some examples of open code and model weights include PVNet [291], which takes historical satellite data, NWP, and live photovoltaic system readings to predict regional

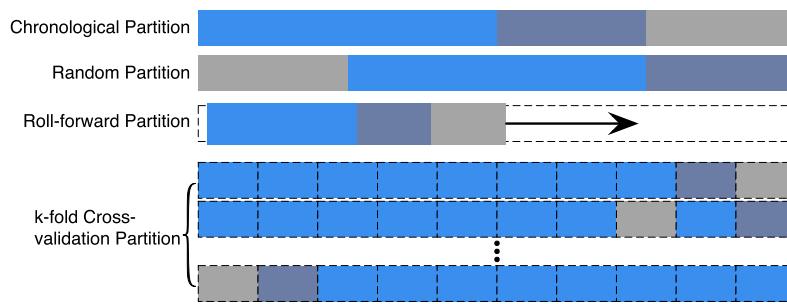


Fig. 24. Data partitioning strategies. Blue, purple, and gray indicate training, validation, and testing datasets, respectively.

photovoltaic power output for the United Kingdom, or transformer-based models, such as Power Perceiver [68], which takes similar data and predicts both site-level solar output and regional forecasts and builds on the work of the more general Perceiver and PerceiverIO architectures [292,293]. Showing the benefits of open code, other, related open-source models for rain forecasting have been adapted for solar forecasting [138,139,287]. These models were trained on open datasets [64,216] and released on Hugging Face [65] for use by other researchers and practitioners.

A combination of sky image, satellite, NWP, and photovoltaic power output dataset and baseline model code was made open in [294], providing a set of simple benchmark models for comparing approaches. Some simple open models include [295], which uses an LSTM to forecast photovoltaic power output. SUNSET [296] is an open-source CNN model that uses both imagery and photovoltaic power output history to predict short-term solar generation.

The International Energy Agency PVPS collaboration is working to evaluate current and emerging resources for solar forecasting, including both models and data. The group is planning on releasing their findings and both data and model products by the end of 2023 [297].

3.5. Quantitative and qualitative evaluation

With decades of development, a consensus has been reached on how to validate a proposed solar forecasting model [298,299]. When applied to computer vision-based solar forecasting, the same standards and practices should be followed. Generally, a data-driven model should be selected, constructed, and optimized using in-sample period data, during which hyperparameters and parameters are tuned with the training and validation datasets. Then, the model will be tested on the testing (or forecasting) dataset in the out-of-sample period. Evaluation metrics are required in both processes to guide the training or verify the performance. The proposed method is proven to be superior if it has better performance compared to the benchmarks based on the same metrics and testing dataset; however, there are several common issues in data partition, metric selection, and competitive model selection in the computer vision-based solar forecasting evaluation and validation. Therefore, we will discuss the three critical components in evaluating a new model: the data partitioning, benchmark setting, and evaluation metric selection.

The intention of data partitioning in solar forecasting is to simulate the real-world situation where only historical data are available for training and updating the model; therefore, data partitioning normally separates the data into training, validation, and testing subsets. Solar time series is chronological and has evident diurnal and seasonal trends, which makes the data partitioning more important in the model evaluation.

The most natural partitioning method for time-series prediction is chronological partitioning, where the training dataset consists of the most ancient samples, the testing dataset has the most recent samples, and the validation dataset contains the remaining samples. This data partitioning method requires multiple years of data to ensure that the

evaluation is unbiased. For example, six years of sky image data were used in [117], where the testing dataset included two years of data with more than 52,000 samples. In a CNN-based satellite image solar forecasting work, three years of data were used for training, and the last year of data was used for testing [300]. The second widely used data partitioning method is random partitioning, which leaves one or multiple period(s) for testing and the rest for training and validation. This partitioning method does not require the temporal orders of the three subsets, which means the testing period can be sampled anywhere in the time series. The most famous example is the global energy forecasting competitions [301]. In [302], two years of GOES-13 satellite images data were used to develop irradiance forecasts up to 5 hours ahead, where the 2017 data were reserved for algorithm training and the 2016 data were used for evaluation. This method can ensure fair comparisons, but it cannot reflect the real-world forecasting practices and might miss the long-term trend patterns and heteroscedasticity in the forecast time series.

These two methods are usually nonoverlapping partitioning methods. Two overlapping partitioning methods are widely used in computer vision-based solar forecasting, shown in Fig. 24. The first is roll-forward partitioning, which uses a fixed-length moving window to partition the dataset with updated information. The forecasting model is retrained in each step with the newly constructed subsets. This method is usually used in traditional statistical methods that take less time to train. For example, the Bayesian model averaging model was combined with a moving window to generate probabilistic forecasts by post-processing the deterministic solar forecast scenarios [303]. The second overlapping partitioning method is the k -fold cross-validation. It splits the data into k groups, holds one group of data for testing, and uses the rest of the data for training/validation. The procedure is repeated k times to cover the entire dataset in the testing, and the results are evaluated and summarized. k -fold cross-validation is a resampling method usually applied to limited data. For example, a 10-fold cross-validation method was used in a study on cloud motion speed estimation to augment a dataset with only 500 images [304]. Although k -fold cross-validation can mitigate a lack of data to some extent, this method is only powerful when applied to datasets with decent sizes. All four methods can be applied to validate the computer vision-based solar forecasting models. But there are several bottom-line rules: (i) models should be tested with out-of-sample data, (ii) the same data partitioning method should be applied to the proposed forecasting model and the competitive models for fair comparisons, (iii) at least one year of data should be included in the testing dataset to maintain seasonal patterns and various weather conditions, and (iv) the data partitioning method should be described in detail and should be reproducible.

Another important aspect when evaluating a new model is benchmarks, which are a collection of competitive models for comparison. Benchmark models should be a mix of classic baseline models, state-of-the-art models, and similar models in the immediate family. For example, in sky image-based very short-term solar forecasting, the persistence methods are usually used as the baseline method. Traditional time-series methods, such as the autoregressive integrated moving aver-

age method (ARIMA), and standard shallow machine learning methods — such as neural network, SVR, random forest, and GBM — are always used as benchmark methods. Taking transformer models as an example, their immediate competitive models include CNN without an attention mechanism and CNN with different attention mechanisms. With the advances in open-source research, a number of easily implemented models are available to compare with computer vision-based solar forecasting models [39,305]. Detailed descriptions of these models can be found in Section 3.4. In addition to the benchmark selection, model construction and optimization are also critical for fair comparisons. It was criticized in [54] that model parameters were manipulated to allow the competing models to be dominated by the proposed models; hence, both the proposed model and the benchmark models should be designed and optimized using the same strategy and method. It is also important to show the optimization process and results for reproducibility.

The two important types of goodness defining a good forecast are quality and value [298,306]. The former measures the correspondence between forecasts and observations, and the latter measures the benefits of accurate forecasts to end users. Most computer vision-based solar forecasting research uses one or more quantitative evaluation metric(s) to assess the quality of forecasts. There are a redundant number of review articles discussing the quality evaluation metrics in solar forecasting [57,298,307]; however, only a limited number of articles focus on both algorithm development and benefit assessment in computer vision-based solar forecasting. For example, sky image-based solar forecasts were developed to mitigate the photovoltaic power fluctuations through a battery smoothing control method [308] and a battery-less proactive ramp down strategy [309], respectively. Similarly, sky image-based solar forecasts were used for photovoltaic power ramp-rate control with a higher control rate [310]. A possible reason for a lack of forecasting articles favoring the value of forecasts is that the forecast value highly depends on the end users, which need additional power system modeling and simulations. It is also challenging to set up a unified standard to assess the benefit and value of forecasts; hence, in this review, we only briefly introduce the quality metrics and give a few examples of evaluating the value of forecasts.

Solar forecasts can be evaluated both qualitatively and quantitatively. Depending on the objective in the forecasting methodology workflow and the output format, the quantitative evaluation metrics can be divided into regression metrics and classification metrics, or deterministic metrics and probabilistic metrics. Table 4 lists a collection of metrics in each group. The most widely used metrics are forecasting errors, which quantify the discrepancies between observations and forecasts, including MAE, RMSE, MSE, mean bias error (MBE), and their variants [307]. Error metrics are also used to compare the developed model with the baseline to show the improvement, which is defined as the forecasting skill. RMSE is usually used in the solar forecasting literature [298]. The coefficient of determination [76] and the correlation coefficient [77] are also used to assess the accuracy of the forecasts. In probabilistic solar forecasting, reliability and sharpness are the two major qualitative characteristics to assess [61]. Reliability refers to the correspondence between the forecast probability and the true probability. Sharpness measures the uncertainty density or the predictive distribution narrowness. The former can be assessed by the prediction interval coverage probability, the average coverage error, and the latter one can be quantified by the prediction interval averaged width [299,311]. The continuous ranked probability score (CRPS) is a metric to jointly measure the reliability and sharpness; it can be complemented by several other scores, including the Winkler score [312], quantile score, interval score, and ignorance score [299].

The regression predictions can also be qualitatively validated. The most widely used is by plotting the time series of the actual values, forecasts, and errors, which directly shows the discrepancies between the actual and forecast values. Another common approach is plotting the distributions. For example, the histogram, quantile-quantile plot, and the cumulative distribution plot of the deterministic forecast errors

are used to assess the goodness of the solar forecasts or to investigate the correlation between solar and wind power output forecast errors [313–315]. Another way of using distribution to time-independently verify the forecasts is the joint distribution of observations and forecasts, which can illustrate the bias corresponding to the actual target values. Probabilistic forecasting performance can also be visually evaluated. For example, a rank histogram is used to assess the dispersion of ensemble forecasts by looking at the flatness. A probability integral transform histogram is used to evaluate the calibration property of cumulative predictive distributions of the probabilistic forecasts. The prediction interval coverage probability (PICP) is a metric of the reliability of a forecast, while the prediction interval normalized averaged width (PINAW) measures the sharpness. Reliability and sharpness diagrams demonstrate the reliability and sharpness indices for different confidence levels. These metrics are commonly used, but they are not the most appropriate for solar irradiance forecasting. The definitions and discussions of these probabilistic evaluation plots can be found in [316].

Solar forecasting also involves some classification problem, such as event prediction, cloud segmentation, and cloud classification. Satellite images or meteorological data are used to analyze the presence of clouds or their types. Evaluation of the classification results is different but could also be performed quantitatively and qualitatively. The most commonly used quantitative evaluation metrics are confusion matrix-based metrics, including accuracy, precision, recall, specificity, and error rate. These metrics can be used for binary classification problems in solar forecasting, such as ramp detection and event existence prediction, and also for multiclass classification problems, such as cloud type classification. Three more comprehensive metrics that can also be derived from the confusion matrix are the F-score, the J-statistic, and the Matthews correlation coefficient. The first measures the harmonic mean of precision and sensitivity, the second uses sensitivity and specificity, and the last is the correlation between the observed and predicted class considering the class imbalance impact. The receiver operating characteristic (ROC) curve is qualitative that measures the misclassification rate of one class and the accuracy of the other. The corresponding quantitative metric is the area under the ROC curve that measures the quality of binary classifiers.

Some metrics are especially useful to evaluate possible class imbalances in the cloud segmentation problem, such as the intersection over union, the Dice coefficient, and the Jaccard index. In the literature, there have been a few cloud classification methods [317,318]. The experimental results show that adding the cloud classification step to solar forecasting improves the forecasting performance [122,149,121]. The overall performance of the forecasting model with the cloud classification step can be evaluated based on the metrics in Table 4.

Solar variables have diurnal and seasonal patterns and are highly impacted by weather conditions; therefore, it is always encouraged to separate the evaluation by hour, by month, and by weather condition to take a closer look at the forecasts. As an unsupervised process, clustering is used to distinguish and identify the type of each time period in the training data. Due to the customization of data labels, evaluation of the clustering results is challenging. The three deterministic metrics that are commonly used as quantitative metrics to assess the clustering performance are connectivity, silhouette, and the Dunn's index.

3.6. Applications and end users of vision-based forecasting

Forecasts of solar and wind power have been traditionally used by power plant operators for economic and system reliability benefits. In surveys on the adoption of forecasts of variable energy generation in the Western Interconnection conducted in 2012 and 2014, many surveyed balancing areas reported that they regarded forecasts as a necessity for electrical system reliability and effectiveness [319,320]. These surveys demonstrated economic benefits in which a 1% reduction of forecast MAE could be worth up to millions of dollars per year in some op-

Table 4
Different metrics used to evaluate computer vision-based cloud segmentation, sky (i.e., atmospheric) condition classification, and solar forecasts with deterministic and probabilistic deep learning models.

Evaluation	Deterministic	Probabilistic
Quantitative	MAE, RMSE, MSE, MBE, NMAE, NRMSE, MAPE, RMSPE, R ² , forecasting skill	Skewness, kurtosis, PICP, PINAW, negative log-predictive density, CRPS, pinball loss Winkler, quantile, interval and ignorance scores
Qualitative	Confusion matrix, ROC curve, J-statistic, F-score, Matthews correlation coefficient, Jaccard index, Dice coefficient	Bayesian and Akaike information criterion, entropy, reliability, sharpness, integrated complete likelihood

erating entities [320]. Previous studies also show the economic and flexibility benefits of improved forecast accuracy in electric power systems [321,322]. Forecasts can also be used for other purposes, such as scheduling and operations of electricity markets, procurement of operating reserves, determination of optimal market bids for power plants, and improvement of situational awareness [323–326].

Electricity market participants include producers, such as power plants that provide electricity to the grid, and consumers, i.e., end-use customers who consume electricity. Unlike market operators, the operations of market participants are profit-oriented, and their goals are typically maximizing their revenues from electricity markets by bidding into energy and/or ancillary markets. Solar forecasts can be used in the determination of optimal operating schedules of market participants. These models usually represent solar power uncertainty and market price volatility with probability distributions, scenarios, or uncertainty sets, and they optimize their bidding strategies using stochastic optimization, robust optimization, and chance-constrained optimization [327–333]. Satellite data, such as measurements of GSI, can be used as input to produce short-term solar power forecasts and to inform the optimal bids into electricity markets for both photovoltaic and CSP plants. Previous studies have shown both financial and reliability improvements in different markets by using forecasts from satellite data [334,335]. In addition to energy markets, solar forecasts can add additional value to solar power plants in ancillary service markets. For example, [336] demonstrated that photovoltaic can provide frequency regulation when satellite-derived solar power forecasts were used. On the demand side, electricity consumers can adjust their load to mitigate load imbalance or to improve economic performance in response to time-varying electricity rates. Demand response actions can be realized by simply selecting the on/off settings [337] or by dynamically varying levels of end-use electric load, such as HVAC (heating, ventilating, and air conditioning), based on satellite-derived solar power forecasts [338,339].

Operating reserves are key ancillary services that serve to stabilize power system frequency, which is an indicator of the balance of real power [340]. Although the names, types, and functions of these operating reserves might vary with markets, most markets procure a variety of operating reserves with response times ranging from approximately 5 minutes (e.g., regulating reserves) to 30 minutes (e.g., nonspinning reserves) [341]. Forecasting plays an increasingly important role in the determination of the requirements of operating reserves because of increasing renewable penetrations in modern power systems and because too high or too low requirements can sacrifice cost-effectiveness or system reliability, respectively. Methods of determining requirements of operating reserves can be categorized into endogenous and exogenous methods [342]. Endogenous methods usually rely on market dispatch models to co-optimize the operations of energy and operating reserves, and they represent renewable and load uncertainties explicitly using a variety of techniques. These models give the amount of required reserves to achieve predefined reliability levels, which usually comply with market standards [343–345]. Despite their popularity in the literature, these methods are rarely adopted by real-world market operators because they are computationally prohibitive. Exogenous methods, in contrast, are less computationally ex-

pensive and are more widely used in the real world. These methods determine the requirements of operating reserves based on uncertainty levels of net load (i.e., total electric demand minus non-dispatchable power, such as wind and solar). Typically, distributions of net load are constructed from historical data [346]; however, because of increasing renewable penetration, historical data might fail to reflect the latest system states or weather conditions; therefore, the latest studies propose addressing this issue by using forecasts [347,348]. Because net load is a linear combination of gross load, wind power, and solar power, distributions of net load forecasting errors are determined by distributions of all constituent components. Although these components can be assumed to be independent [116], studies suggest that they often lead to conservative results [346,349,350]; therefore, some studies use copula to account for correlations across variables [351–354]. In addition, forecasts of solar power, including sky images and satellite data, can be further processed to estimate distributions of net load forecasting errors using a variety of techniques, such as time-series models [355], k-nearest neighbors [347], regressions [348,356], and neural networks [116,357]. Previous studies suggest better economic and reliability performance from using weather-informed requirements of operating reserves [347,355,358–361].

Because of their dependence on meteorological conditions, both solar and wind power are subject to power ramps, which can significantly impact grid security; therefore, utilities usually rely on active power curtailment or energy storage to reduce ramp rates and smooth solar power output [362–366]. Ramp forecasts enable operators to take informed actions prior to the occurrence of solar ramps and thus improve grid reliability and economic performance [367]. For example, Wen et al. [364], Saleh et al. [365] made solar power forecasts using sky images, and they developed ramp control strategies that comply with grid regulations based on their ramp forecasts. The improved ramp control strategies can ameliorate system cost-effectiveness, and they can sometimes even reduce the battery degradation of a hybrid photovoltaic and battery system by reducing the energy throughput over a battery's lifetime [367]. Although most studies use deterministic forecasts, probabilistic forecasts can also be used to inform ramp control strategies [363].

4. Deep learning techniques for vision-based solar energy meteorology

This section presents diverse deep learning methods used in vision-based solar energy modeling. These techniques focus on various aspects of the general modeling pipeline including architecture tuning, data processing, model training, interpretability and uncertainty quantification.

4.1. Architecture tuning

Implementing deep learning models often involves numerous architecture design choices (e.g., number of layers, size of convolutional filters, and number of neurons) or different learning setups (e.g., learning rate and batch size). Tuning these hyperparameters to improve the performance of deep learning models is a nontrivial job. Normally, for

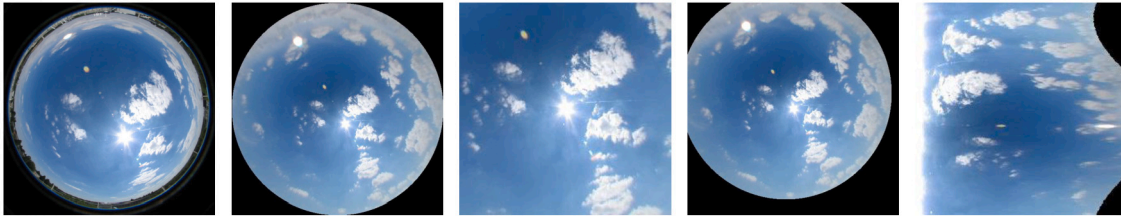


Fig. 25. Different scene representations of a cloud cover observation [136] (source: SIRTAs [129]). From left to right: 1. raw (distorted) image of the sky taken with a fish-eye camera, 2. undistorted sky image, 3. close-up on the circumsolar area, 4. sky image centered on the sun's position, and 5. polar coordinates.

each hyperparameter, a range of possible values are defined. An exhaustive search (i.e., grid search) including every combination of these parameters is often impractical given the constraints on computational resources and time. Further, hyperparameters could be dataset dependent [368], i.e., the optimal hyperparameter settings obtained for a given dataset might not be optimal for another one; thus, grid search is not efficient for hyperparameter searching. In view of these issues, sampling methods (i.e., random search) are usually employed. Instead of a discrete set of values, a statistical distribution is defined for each hyperparameter from which the values can be randomly sampled. In most cases, hyperparameters are not equally important [368]. Random search has more exploratory power and can focus on finding the optimal values for the important hyperparameters, whereas grid search might waste time in exploring unimportant parameters [368]. In addition to random search, Bayesian optimization has recently emerged as a powerful tool for hyperparameter tuning. Unlike random search, which needs a lot of parallel but isolated experiments to find the optimal hyperparameter settings, Bayesian optimization is a machine learning algorithm that can be used to automatically search for the optimal hyperparameter settings based on multiple iterations [369]. The Gaussian process is often used to estimate the prior probability of model performance across the hyperparameter space [370]. The results from previous iterations can be used to determine the sampling method for the next iteration based on an acquisition function until it converges to an optimum [369].

For solar forecasting using deep learning models, few studies have shown systematic experiments on architecture tuning with different searching strategies. For categorical hyperparameters, a uniform probability distribution was used, whereas for continuous hyperparameters, the Latin hypercube sampling strategy was applied [371]. In [122] the authors used Bayesian optimization instead of grid search for cross-validating the hyperparameters.

4.2. Data-centric methods

In addition to model-centric approaches aiming to improve the model architecture or the training strategy, some works try to facilitate modeling by improving the data. These data-centric techniques cover a range of aspects, such as feature engineering and data augmentation.

4.2.1. Feature engineering and scene representation

Feature engineering aims to use domain knowledge to extract new features from raw data. For instance, a common preprocessing step in vision-based solar forecasting is to unwrap sky images taken by fish-eye lenses, which cause a strong distortion of the scene (see panels 1 and 2 in Fig. 25). This can be done by projecting clouds on a rectangular grid assuming, for example, median cloud height [178,372]. The resulting scene representation aims to retrieve the actual shape and relative size of the clouds distorted by the lens, hence resembling satellite observations. In deep learning studies, this method was applied to provide object trajectory consistency (sun and clouds) [97] or to facilitate the prediction of distant cloud displacements [131]. In contrast, the original distortion imposed by the lens was also shown to improve solar predictions by providing more detail on the center of the image where

the sun and surrounding clouds are visible in high solar flux conditions (high solar zenith angle) [40]. Providing more details on the point of interest (the sun in sky images or a solar plant/meteorological station in satellite observations) can also simply be performed by magnifying the ROI. This was shown to improve shorter-term forecasts by enabling the model to focus on nearby clouds or by facilitating the estimation of the current irradiance level from a more detailed circumsolar area in sky images (third panel in Fig. 25) [136]. Similarly, centering the whole image on the point of interest (fourth panel in Fig. 25) provides information on its location even if it is hidden by a cloud. In sky images, for instance, the position of the sun is a crucial indicator of the level of clear-sky irradiance.

Computer vision-based solar irradiance forecasting is approximately invariant by rotation around the point of interest because future irradiance shifts are independent from the direction from which clouds hide the sun (in sky images) or a solar plant (in satellite observations). To leverage this property, the data can be augmented with rotations (see subsequent section on data augmentation) or with the use of polar coordinates [373], which turn the rotation invariance of the problem into a shift invariance (fifth panel in Fig. 25). The resulting shift invariance of the scene can be better exploited by convolutional architectures with pooling operations, which are translational invariant (Fig. 25): a given sun-cloud (or solar plant-cloud in satellite images) spatial configuration can be learned from a single image without having to augment the data [136]. In addition, the polar representation magnifies the ROI, leading to improved short-term irradiance predictions. This approach can be further augmented by centering the polar representation on the incoming flow of clouds, which can be determined by an optical flow algorithm [374].

4.2.2. Data augmentation

Data augmentation aims to artificially increase a dataset by generating modified versions of existing samples. Instead of collecting additional data, which can be expensive or impracticable, data augmentation tries to better exploit available data to improve a model's generalizability. In practice, synthetic data can be obtained from generative models, such as GANs, or by modifying existing samples with visual transformations, such as noise addition, cropping, rotations, or flipping.

Various data augmentation approaches have been applied to vision-based solar forecasting (see Fig. 26). For instance, noise injection was performed by adding Gaussian noise to pixel values [375]. A different color space transformation technique consists of randomly changing the intensity level of the three RGB channels independently (colorcasting) [375,376]. Alternatively, the pixel values of the image can be increased or decreased by a constant value across all three channels simultaneously (brightness adjustment) [375]. Applied to solar modeling from sky images, these approaches have been shown to benefit the now-casting task of correlating an image with the corresponding solar value, whereas the impact on the forecasting task is not significant [375].

Another set of data augmentation techniques involves mixing images. The synthetic minority over-sampling technique [377] for regression [378] creates data by fusing an observation with a randomly selected neighbor from the same set (see Fig. 26). The corresponding

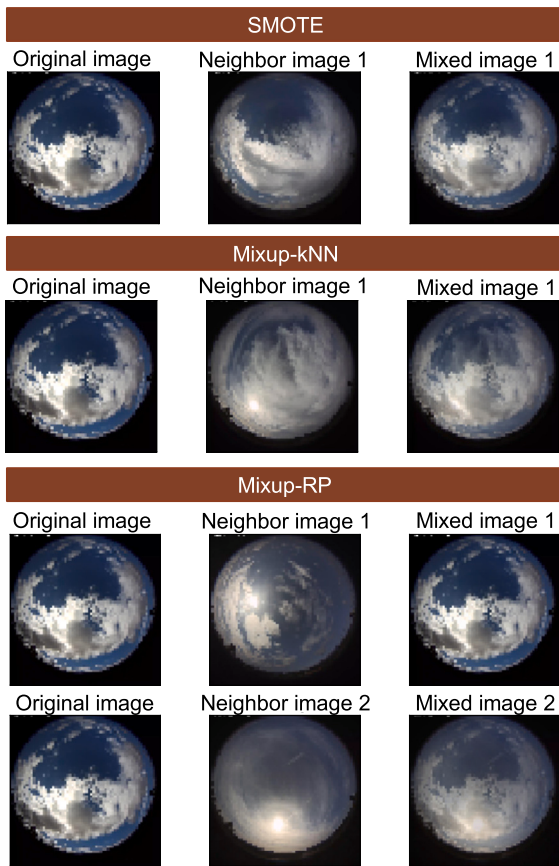


Fig. 26. Demonstration of different data augmentation techniques applied to sky images. Source: [375].

target variable is defined by the distance between the generated sample and each of the two seed samples [375]. Alternately, the *mixup* method [379] constructs virtual samples by taking a convex combination of example pairs and their labels [375]. Original images can be mixed with a random sample or with one of its k -nearest neighbors. Similar to the color space transformation methods presented in the previous paragraph, mixing images was shown to facilitate nowcasting predictions the most [375].

Image transformation — such as rotations, cropping, or vertical/horizontal flip — have also been used in recent works [136,380–384]. For instance, random cropping is a common strategy in satellite-based machine learning approaches that deal with very large Earth observations covering countries or continents [38]. In addition, rotations applied to sun-centered sky images or satellite images centered on a solar site leverage the rotational invariance of vision-based solar irradiance forecasting [136]. Note that this method might be less effective with photovoltaic power output forecasting because the spatial distribution, tilt, and orientation of the panels impede this rotational invariance. Further, a polar representation can be augmented with vertical translations that correspond to rotations of the original image. Finally, temporal flip consists of randomly flipping image sequences to predict the past instead of the future, hence producing more diverse temporal dynamics [136].

Although augmenting the data by applying various transformations provides some additional sample variability during training, its potential remains constrained by the diversity of cloud patterns in the dataset. A physics-based simulator of sky camera observations has been developed, but it has not been exploited in a data-driven approach [385]. Alternatively, Nie et al. [375] present cloud image samples generated by a Deep Convolutional GAN (DCGAN) (Fig. 27); however, the co-generation of the images and corresponding continuous prediction tar-

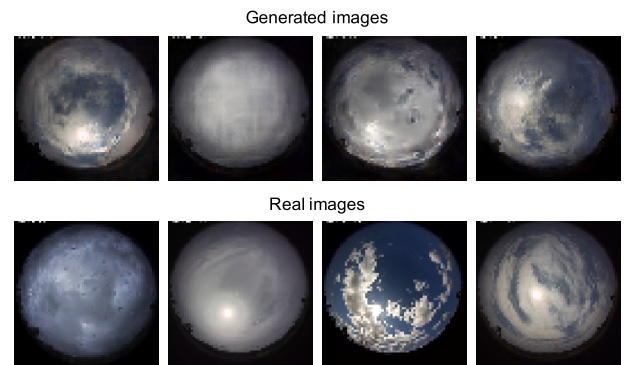


Fig. 27. Samples of cloudy images generated from DCGAN (the first row) using random Gaussian noise and real cloudy images used for training DCGAN (the second row). Note that the generated images are from random seed and do not correspond to the real images. Adapted from [375].

gets, i.e., photovoltaic power output or GSI, remains a key challenge of GAN-based data augmentation for solar forecasting.

In this regard, cloud classification and segmentation are more suitable for the application of GAN-based data augmentation because the labels can be either inherited from the original data (e.g., in cloud category classification, the synthetic cloud images share the same labels as the data used for training the GAN), generated by humans or estimated by algorithms (e.g., in cloud segmentation, the synthetic images can be segmented with human knowledge or algorithms). For example, Jain et al. [386] demonstrate the effectiveness of using GANs to augment datasets for the cloud segmentation task with the corresponding binary segmentation maps estimated by an unsupervised clustering algorithm.

4.2.3. Sampling strategies

Like many real world datasets, satellite observations and sky image datasets are unbalanced in different ways. For instance, the distribution of cloud types strongly depends on the weather conditions at a given geographic location. This results in some cloud patterns being under- or overrepresented in the original dataset. If mini-batches are uniformly sampled from a training set representative of the dataset, the distribution learned by the model will be site-specific, which might impact its performance in different climate conditions. To address this limitation, Nie et al. [375] presents different sampling approaches to overrepresent the minority class (cloudy sky images), and hence best match the target distribution (test set). The experiments conducted in the article show that the proposed sampling strategy mainly benefited the nowcasting task. Similarly, Pothineni et al. [382] augmented the minority class with various data augmentation techniques.

4.2.4. Quality control

In addition to the size of a dataset, its quality can have a large impact on the performance of the model and its generalization properties. For instance, wrongly labeled samples or irradiance mismeasurements can be detrimental to the model training by inducing an erroneous supervision signal and to the model exploitation if it operates in unknown conditions. For these reasons, quality control is an essential aspect of machine learning projects. In vision-based solar power monitoring, the causes of anomalies include equipment failure (e.g., temperature response, shade ring misalignment, spectral selectivity) and operational disturbances (e.g., dust, snow, dew, water droplets, or bird droppings on the camera) [387]. Several methods, detailed here have been developed to accordingly detect these abnormalities.

Sky cameras are exposed to various external hazards impairing the collection of cloud cover observations: animals, soiling, snow, or rain. Inspecting outliers in a dataset is a common strategy to filter these anomalies. For example, by measuring the distance between samples in a feature space, one can define outliers as samples that significantly

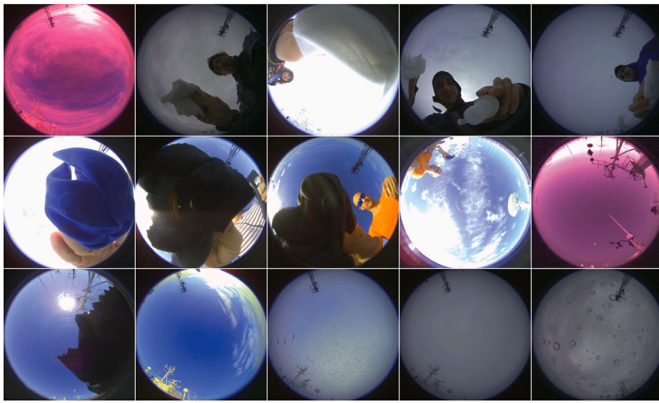


Fig. 28. Examples of sky images marked as outliers by a CVAE [388].

differ from the average representation; however, comparing sky images in the large original RGB feature space is limited because of the curse of dimensionality. To overcome this difficulty, a filtering strategy based on a convolutional variational autoencoder (CVAE) model was proposed by [388]. An autoencoder was trained to reconstruct sky images from a low dimension feature space representation. Outliers derived from this low-dimensional representation are then manually checked to differentiate actual anomalies (e.g., birds or human beings) from visually uncommon sky images (e.g., rare or complex cloud patterns); however, this filtering method might not catch common anomalies, such as rain or dust in sandy regions (Fig. 28).

Quality control of ground-based measurements, such as GSI or photovoltaic power output, entails different signal analysis approaches. A common strategy involves the comparison of measurements with a reference to flag anomalies. For instance, extraterrestrial and simulated GSI levels under clear-sky or overcast conditions help detect abnormal solar flux values [389–393]. Successive measurements or measurements from neighboring stations can also be subject to statistical or spatial coherence tests [394]. In addition, the comparison of the relative duration of sunshine hours with daily global solar radiation provides indications on systemic and nonsystemic errors [395]. Another strategy consists of comparing satellite estimates with *in situ* measurements [396]. More recently, visual tools were developed to support the various quality control approaches [387,397].

4.3. Data fusion

Data fusion is an important area of research in machine learning that aims to integrate data from different sources to produce more consistent and accurate information. It is based on the expectation that multiple data sources are more informative than a single data source. For computer vision-based solar forecasting, several scenarios of data fusion have been explored.

4.3.1. Images and auxiliary data

Computer vision-based solar forecasting often involves heterogeneous input. Besides cloud cover observations, diverse sensor measurements (e.g., GSI, photovoltaic power output, wind speed, wind direction, sun angles) provide crucial local information on the atmospheric and operating conditions of a solar site. In particular, the thermal environment (i.e., air temperature, wind speed, humidity, and air pressure) plays a key role in the overall efficiency of solar panels [398]. In addition to the different scales of data, another challenge is how to make the deep learning system pay due attention to each type of data. The problem arises because imagery data are often high dimensional, i.e., an array with hundreds of thousands of entries, whereas sensor measurements might only be a vector of tens of entries. It is likely that the deep learning model pays more attention to the high-dimensional imagery information than the low-dimensional sensor measurements. This is an

active field of research in applications such as self-driving [399] and robotics [400], but very limited studies have investigated this problem in solar forecasting. In 2019, Venugopal et al. [401] first systematically investigated 28 methods of fusion (MOF), which belongs to 4 families: mix-in; activation map combination; activation map stacking; and two-step, for integrating the hybrid input of a sky image sequence and photovoltaic power output history to forecast 15-minute-ahead photovoltaic power output by using CNN as the backbone. The results show that a two-step autoregressive CNN MOF has the best performance, closely followed by a mix-in MOF that performs feature expansion and reduction to give appropriate importance to the two types of inputs. Similarly, Ajith and Martínez-Ramón [103] developed a multimodal fusion network based on CNN and LSTM for studying GSI forecasts by using infrared images and past GSI data with forecasting horizons ranging from 15 to 150 s. The extracted spatial and temporal features are fused using a fully connected neural network. The proposed model shows up to 46% improvement compared with benchmark machine learning and deep learning models using solely irradiance data. Another approach investigated by [402,40] consists of augmenting images by stacking additional channels containing concomitant auxiliary information, such as the sun distance or the irradiance level. More recently, Terrén-Serrano and Martínez-Ramón [122] explored fusing information from infrared sky images with various other sensors — including a pyranometer, a solar tracker, and a weather station — using a multitask deep learning architecture based on RNNs. The accuracy of an 8-minute-ahead forecast is improved by nearly 20% in terms of forecasting skill.

4.3.2. Camera network

Networks of ground sky cameras are starting to be deployed to offer high-spatiotemporal observations of the cloud cover [403]. A preliminary study was conducted to evaluate the benefit of integrating observations from three cameras collocated at the Abengoa Solar Platform of Solúcar near Seville [402]. Results show that combining the three sources of information with a shared encoder leads to a performance gain.

4.3.3. Satellite and sky images

Often used separately, satellite observations and sky images offer two complementary points of view on the cloud cover. For this reason, some approaches try to combine both sources of images into a single modeling framework [404] or try to evaluate in which weather conditions one prevails over the other [405,406]. In deep learning, both sources of data can be independently processed by parallel spatial encoders prior to being recombined into a single state representation for solar forecasting [13]. Although the method shows valuable improvement over single-vision approaches (sky images only or satellite observations only), some challenges remain, such as the integration in the modeling of observations with different temporal resolutions.

4.4. Transfer learning — domain adaptation

Transfer learning is a commonly used strategy in machine learning to help improve performance or more quickly solve new problems on target domains by transferring the knowledge learned from different (but related) source domains [407]. For vision-based solar forecasting, a solar site could benefit from knowledge carried by models trained on images collected in other locations to improve the accuracy of its site-specific algorithm. Indeed, some common features learned from other sites, such as the sun position or the cloud distribution, can be shared across different locations. This could be especially beneficial for new solar facilities that have limited data collection at the beginning. A recent study by [408] suggests that transferring learning from a large and diverse source dataset to a local target dataset can save up to 80% of the training data while achieving comparable performance for 15-minute-ahead solar forecasting. With more and more sky image datasets being open source in recent years, pretraining a general solar

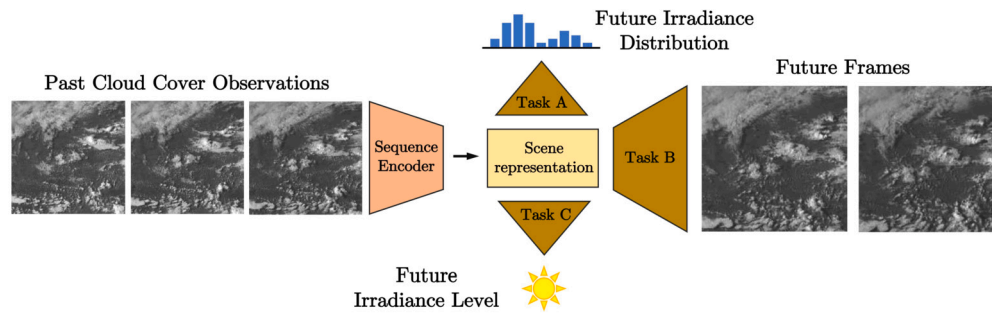


Fig. 29. Neural network architecture for MTL.

forecasting model based on a large-scale centralized dataset with massive and diverse sky image samples is facilitated. Similar to pretraining strategies based on ImageNet, transferring the knowledge from such a model to local sites can potentially save data and jump-start local model development; however, the various types of heterogeneity observed in data collected from different locations remain challenging for transfer learning approaches: different prediction variables (i.e., photovoltaic power output versus GSI), different measurement distributions given local weather conditions, and different camera types (i.e., TSI versus ASI) and orientations. Processing and reconciliation of these datasets is thus a critical step for building a large-scale dataset for knowledge transfer.

4.5. Multitask learning

A Multitask learning (MTL) model aims to address multiple related tasks at once (Fig. 29). Intra-task information can be used to increase the performance of a model and ameliorate prediction consistency such as inter-task consistency (e.g., cloud masking and cloud index mapping, deterministic and probabilistic solar forecasting) and inter-horizon consistency (e.g., short- and long-term horizons). This is generally done using a single model to learn multiple tasks instead of using multiple models to learn each task individually [409]. In deep learning, the multiple tasks do not need to be directly combined using a single model (i.e., architecture). A model can share representations (i.e., hidden space) between its different structures to improve the overall performance with respect to task-specific models.

When MTL is applied to forecast time series, the output space is that of a classic statistical functional model, but in deep learning models, exogenous variables are added to the input space. In addition, performance is improved by including information about past cloud cover from ground-based sky imagers (i.e., tensor signals) in the input space [40,221] or satellite images [13]. From the perspective of the input and output spaces, MTL can be divided into two categories: multi-input multi-output (MIMO) or single-input multi-output (SIMO) [410]. Depending on the inter-task correlation level, the tasks can be considered independent (i.e., uncorrelated), with directional correlation (e.g., time domain), or fully correlated. A MIMO model is a multitask model made of multiple models. Each model can potentially have a different input feature vector (or tensor). If tasks are independent, a multitask model can be assembled from several output-specific models. In that case, the feature vector can also be different (i.e., independent multitask model). These types of MTL models are common in solar forecasting applications [36,103,253].

A SIMO model corresponds to single models (i.e., architecture) that generate multiple outputs from a feature vector or tensor. When the output variables have a time structure, using the sequential correlation between outputs is beneficial. A chain of models is commonly used in deep learning models for language processing. These are known as recursive architectures [411]. In solar forecasting applications, they are proven to improve performance with respect to MIMO models [122]. Chains of deep learning models can form SIMOs, sharing representations from sequential forecasts and exploiting the directional structure in GSI with

an *attention mechanism* [223]. In addition, deep learning-based SIMOs can have encoded representations between decoding structures in a sequential (i.e., directional) manner to forecast CSI and future spatial sky conditions (i.e., cloud cover maps) from consecutive past satellite [13] or ground-based images [40].

A common method to develop SIMO models in deep learning is to define the last dense layer as having multiple outputs. This approach was proposed in solar forecasting applications, using cloud features extracted from multiple images [122] or using batches of multiple sky images [70] to forecast a batch of sky images. In solar forecasting applications, it is possible to implement a SIMO deep learning model to forecast tasks in different output spaces, such as a future sky image and a GSI measurement [412]. This forecasting problem can be solved by using the same architecture with attention mechanisms [222] or by dividing it into multiples ones [221]. The resulting output space is the same, but the forecasting application is different [413]. A SIMO model trained using a single probabilistic loss independently weighting the terms that form each model [414] can be considered as an additional class of MTL models.

4.6. Interpretable deep learning

Often described as *black boxes*, deep learning models suffer from a lack of transparency that hinders their applicability to many sensitive domains, such as transport, healthcare, or justice. Interpretable AI (or explainable AI (XAI)), aims to make such models more intelligible to humans [415]. This research domain helps to justify, control, discover, or improve the behavior of a data-driven method [416]. In practice, model-agnostic interpretability approaches fall into four categories: visualization (visualization of the mapping learned by the model), knowledge extraction (extraction of the knowledge acquired during training and encoded as an internal representation), influence methods (evaluation of the influence of an input change on the model prediction), and example-based explanation (selection of specific instances of the dataset to interpret the prediction of the model) [416]. In addition, note that in some cases, improving the explainability of a deep learning model impacts its performance [415].

In solar energy, the International Energy Agency Task 16 alerts on the lack of interpretability of neural network: “Yet the lack of transparency that usually accompanies artificial intelligence models (i.e., retracing the path that the model took to reach its conclusion) has drawbacks: the difficulty of translating the information generated into basic principles (or physics) and interpreting the information to determine what is happening in the natural world and apply the knowledge gained at a particular project site to other regions. Consequently, an appropriate strategy could benefit from the optimized combination of both physical and artificial intelligence-based models.” [417]. Moreover, increasing the transparency of deep learning solar modeling would have various benefits. By developing trust in AI systems, XAI would increase their acceptance by policy makers or grid operators. In addition, operational actions would be better supported by model insights. Developers

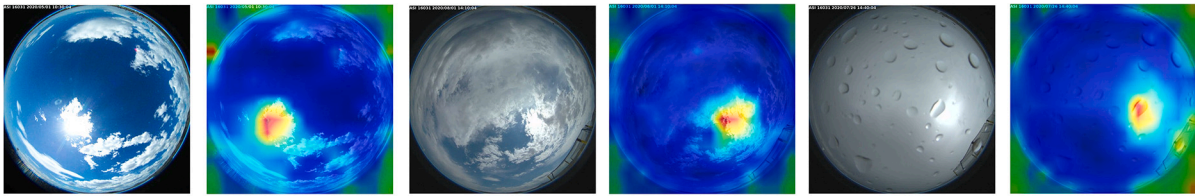
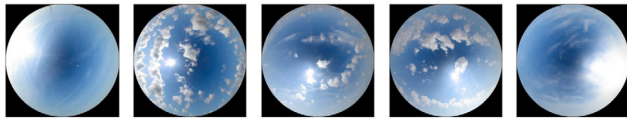


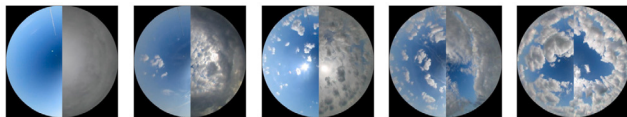
Fig. 30. Visualization of sky images and their corresponding attention maps by a transformer-based nowcasting network on the TSI880 dataset (top) and ASI16 dataset (bottom) from the National Renewable Energy Laboratory (NREL). Taken from [418].



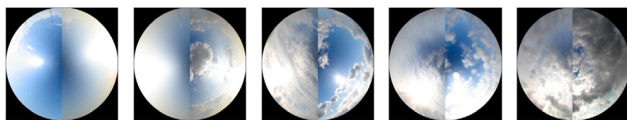
(a) PC1: Extent of the cloud cover.



(b) PC2: Horizontal position of the sun in the sky.



(c) PC3: Spatial variability of the cloud cover: from fully cloudy or fully sunny to partly cloudy. Each image illustrates two neighboring samples from the distribution.



(d) PC4: Vertical position of the sun in the sky. Each image illustrates two neighboring samples from the distribution.

Fig. 31. Four principal components (PCs) of the spatiotemporal representation of past sky images encoded by a video and irradiance prediction deep learning model. Taken from [40].

would gain understanding on the capabilities but also limitations of the AI system, which might trigger further improvements.

Several works in solar energy have presented experiments that aim to improve our understanding of neural network for computer vision. For example, observing feature maps corresponding to particular instances of the dataset provides insights into the types of features extracted by convolutional layers [252,364,380]. In practice, some filters seem to focus on specific elements in the image, such as clouds, the sun, the sky, or the solar tracker [120,380,419]. In addition, the specific visual pattern recognized by a filter can be generated by tuning the input image to maximize the average response of that filter [252]. Another approach aims to generate *hyper columns* (the vector of activations of all CNN units above an input pixel) [420] to derive a heat map that indicates the focus of CNNs in sky images [413]. For transformer-based models, attention maps clearly show the focus on the circumsolar area (see Fig. 30). A principal component analysis of the learned spatiotemporal representation of the past sequence of sky images highlighted the most significant features used by a deep learning model: the extent of the cloud cover, the horizontal and vertical position of the sun, the spatial variability of the cloud cover (from fully cloudy or fully clear-sky to partly cloudy) [40] (Fig. 31). This method achieved similar results when applied to the polar representation of sky images [136].

In addition to these purely explainable AI approaches, some training objectives provide additional interpretability and transparency for deep learning modeling. For instance, providing the uncertainty associated

with a forecast informs on the confidence of the model in its prediction [38,40]. Similarly, predicting sky or satellite observations as well as insolation levels provides visual feedback on the future cloud cover expected by the model. From that perspective, the blurrier the predicted cloud cover, the higher the associated uncertainty.

4.7. Probabilistic forecasting

In a recent review of the field of AI applied to solar power forecasting, the investigation of probabilistic methods was seen as a critical research direction [421]. Advancing the quantification of prediction uncertainties would benefit power systems [422] while increasing the interpretability of deep learning modeling (see Section 4.6). In mathematical terms, a probabilistic model aims to predict the probability distribution, $P(Y|X)$, of a predicted value, Y (e.g., future photovoltaic power output or GSI value, cloud type) conditioned by some observations X , (e.g., past sky or satellite images, historical measurements).

In solar forecasting, the variability observed at a short-term timescale mainly comes from the formation, dissipation, and displacement of clouds; thus, vision-based solar forecasting could play a major role in improving the quantification of this type of uncertainty (Fig. 32). A common strategy in deep learning consists of predicting a histogram instead of a single value (Fig. 29). This probabilistic framework can be applied to local or regional solar forecasting [38,13,40].

Likewise, probabilistic modeling is commonly used in other computer vision tasks involving classification, including segmentation [423–425] or cloud type classification [426,427]. The output of the model is passed through a Softmax activation function (or eventually a Sigmoid activation function for binary classification) to retrieve the probability of each category. Alternatively, the quantile loss function associated with a quantile q , which is

$$QL(\hat{y}, y) = \max[q(\hat{y} - y), (q - 1)(\hat{y} - y)], \quad (1)$$

where \hat{y} the model prediction and y the ground truth, has been applied to probabilistic radiative transfer modeling [428] and solar power forecasting [429].

5. Discussion

This review study highlights the considerable advances observed in recent years in the field of vision based-solar energy modeling. Despite this rapid progress, many technical limitations and implementation challenges have still to be overcome to fully benefit from the development of deep learning. The following subsections describe some of the research areas that require further investigation, and discuss potential solutions to facilitate the operational adoption of these rapidly advancing techniques.

5.1. Road map for future research

The demand for cloud cover-informed solar predictions has motivated an increased interest toward computer vision-based solar forecasting approaches. In this context, the increasing availability of open-source sky and satellite image datasets has paved the way to the application of data-driven methods to address this challenging task. In particu-

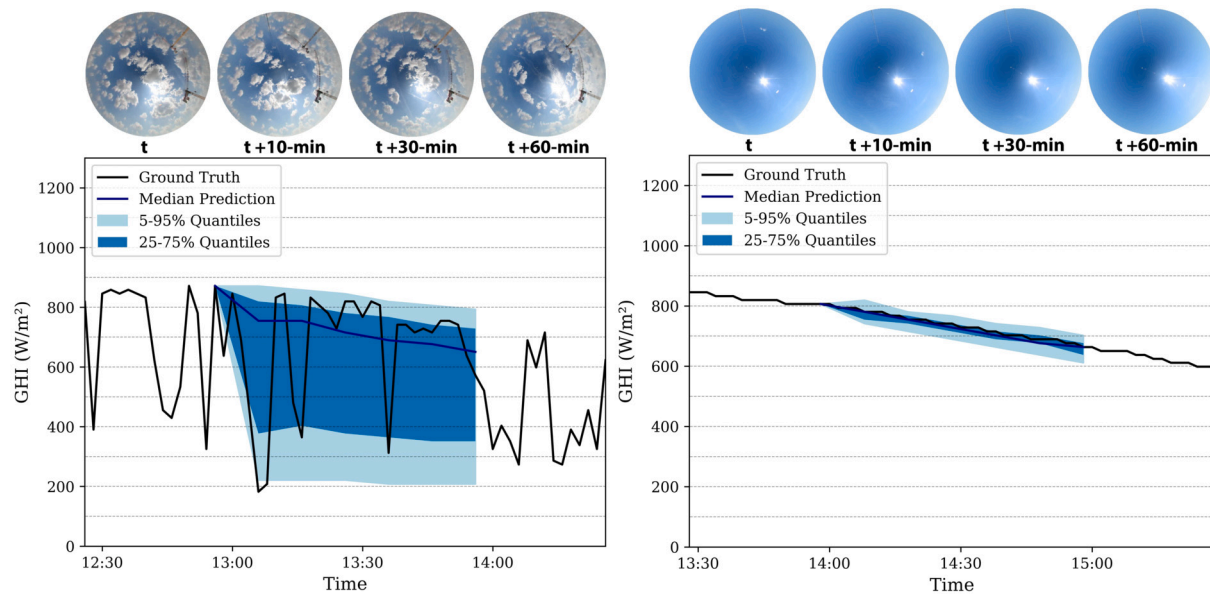


Fig. 32. Probabilistic predictions based on past sky and satellite images in diverse weather conditions. The uncertainty predicted by the model adapts to the situation: high uncertainty in partly cloudy conditions, and low uncertainty in clear-sky days. Taken from [13].

lar, neural network is capable of (i) modeling large-scale spatiotemporal solar data, (ii) handling multimodal data, and (iii) providing accurate predictions with significantly less computing time. In addition, this rapidly evolving field benefits from the maturity of some computer vision research areas (e.g., image classification or image segmentation) while remaining open to new advances (e.g., physics-informed neural network or video prediction). Despite promising results, some theoretical and practical challenges listed here remain or could be further investigated to facilitate the applications of solar forecasting in power system management and operations.

5.1.1. Forecast verification

Forecast verification is crucial to evaluating and comparing the performance of solar forecasting models. There exists a wide range of methods and metrics (see Section 3.5) to expand beyond the limited evaluation scheme of assessing a model's performance based on a metric corresponding to the training loss (e.g., RMSE and L_2 loss). In particular, few works have tried to propose or apply more advanced loss functions adapted to particular objectives (e.g., decreasing temporal distortion, prediction of sharp solar power changes). Further, contributions on the topic of value-oriented solar forecasting verification and training would strongly benefit many downstream applications.

5.1.2. Benchmark datasets

Existing solar forecasting models are mostly developed and evaluated based on different datasets, making it hard to consistently compare the performance of these models. Although there are studies benchmarking certain types of solar forecasting models based on a single dataset [39], there are currently very limited studies cross-comparing different types of solar forecasting models, i.e., deep learning-based models (e.g., CNN, LSTM, ConvLSTM, etc.), machine learning-based models (e.g., extreme grading boosting, random forest, SVR), time-series models (e.g., autoregressive, ARIMA), and physics deterministic models (e.g., cloud motion vector, advection) on standardized datasets. Another challenge is to reproduce the models published in existing studies. Erroneous implementation of the existing models can lead to false comparisons. With more and more open-source datasets and models available in recent years (see descriptions in Sections 2.5 and 3.4), we highly encourage the researchers to report their model performance by comparing with open-source models (e.g., SUNSET [119], PhyD-Net [221]) and standardized datasets (e.g., UCSD [202], SKIPP'D [204],

SIRTA [201], Girsol [430]), and we further call on the efforts to build a large-scale sky image dataset by assembling suitable open-source datasets for solar forecasting to benchmark different types of models and to accelerate solar forecasting methods development.

5.1.3. Probabilistic forecasts

Probabilistic modeling can provide valuable feedback on the uncertainties associated with solar power forecasting, including the stochastic nature of the cloud cover dynamics, the lack of information on distant clouds, such as their location and spatial configuration. To date, few studies have investigated the potential of the probabilistic framework to ameliorate the prediction of deep learning models for solar forecasting [13,431]. An application of a probabilistic forecast is stochastic electric grid planning, which requires generating random realizations of future renewable energy generation outcomes [432]. This is especially important to consider the risk introduced by certain generators in the realization of a potential extreme weather scenario. This application requires a fully probabilistic instead of a probabilistic approach. A fully probabilistic approach allows for drawing random samples from a predictive probability density [122].

5.1.4. Explainable artificial intelligence

A major downside of current deep learning methods is the lack of interpretability of their predictions [417]. Although probabilistic deep learning approaches can provide some insights on the predictions of a network, a stronger focus on more diverse explainable AI techniques will foster the acceptance for deep learning-based solar power forecasts. Moreover, a deeper understanding of solar modeling with neural networks is likely going to clarify some limitations and trigger new technical development.

5.1.5. Site adaptation

The limited generalization skills of current deep learning models are a strong limitation to their widespread implementation. Recent work on sky images has shown that domain adaptation strategies could significantly facilitate the transfer of knowledge from large datasets to new solar sites [408]. The main challenge to address in that context is how to provide valuable solar forecast, with no or limited local data. This difficult problem could be tackled by building more robust and general foundation models based on the large number of diverse open-source

datasets that are capable of being adapted to a wide range of downstream tasks [34].

5.1.6. Edge computing and privacy-preserving algorithms

Coupling solar forecasting with edge computing and distributed sensing to facilitate the fusion of global and local information is another critical research direction. Further, designing data-sharing mechanisms and privacy-preserving deep learning algorithms that protect privacy while improving power and energy system efficiency is a valuable emerging topic in solar forecasting.

5.1.7. Synthetic data generation

Generating synthetic data to guide a deep learning model training is an appealing strategy to address edge cases, improve generalization properties, pretrain deep learning models, and adapt a training to specific environments or weather conditions. A solution to that problem would be to develop realistic simulations of the cloud cover dynamics and its physical properties, as in [385]. Alternatively, GANs or VAEs are powerful tools to produce new observations from an approximated data distribution [375]. These models could help generating realistic sequences of cloud cover observations as well as corresponding solar values.

5.1.8. Physics informed deep learning

Introducing prior knowledge (e.g., physics laws) into deep learning models can potentially increase their performance while improving their interpretability by (re)discovering realistic physical equations from observations. For instance, PhyDNet [221] constrains a neural network to learn partial differential equations that best describe the cloud cover dynamics to predict future cloud displacements. Streamlines are the paths that floating clouds trace in the atmosphere. Machine learning methods based on computer vision are capable of visualizing streamlines in satellite images [433] and ground-based sky images [134] but in a suboptimal manner. In this context, a physics informed deep learning method has the potential to optimally visualize fluid mechanics field lines.

5.1.9. Morphology-informed solar power modeling

Integrating the solar panel spatial configuration and local shadowing effects caused by vegetation, buildings or terrain variation, has the potential to ameliorate the accuracy of solar forecasts. This information could be inferred from irradiance measurements [387], object detection from sky images [434] or high resolution remote sensing observations [435,279], LiDAR data [436], and other *in situ* 3D mapping measurements [437]. Although local shadowing and morphological effects could be learnt by the model providing enough cloud cover observations and corresponding solar data, approaches based on transfer learning and multilocation dataset fusion might benefit from the development of morphology-informed methods.

5.2. Industry adoption and implementation

In addition to improving the technique, advancing its operational implementation is vital to reduce the uncertainties associated with weather variability and to contribute toward an effective integration of clean energy into energy systems. Some companies in the industry and private sector have adopted satellite imagery and ground-based sky imagers solar forecasting. For example, Steadysun and Reuniwatt provide intra-day and day-ahead solar forecasts combining machine learning with sky images and geostationary satellite imagery. Similarly, Solcast (recently acquired by DNV in January 2023) uses sky images, weather satellite imagery, machine learning, and computer vision for day-ahead solar forecasting models. Open Climate Fix, a UK-based startup, relies on satellite imagery, NWP, vertical cloud profiles, geographic information, and machine learning to predict several hours-head photovoltaic power output for UK's National Grid. SolarAnywhere fuses data from

multiple sources, including satellite imagery, using deep learning in their forecasting models.

However, industry adoption of solar forecasting techniques and tools remains slow compared to the pace of development and innovation introduced by research teams. This can be explained by several factors:

1. Lack of resources to assimilate the innovations and deploy them in the operational tools.
2. Lack of computing capacity to implement the methods, monitor the forecasts, and evaluate the added-value.
3. Difficulty to integrate new and potentially complex tools in existing operational processing chains, including to build the links with the necessary input data (e.g., satellite image fluxes in real time).
4. Inadequacy and lack of adaptability of operational tools to integrate forecasting models. For instance, some forecasting tools are still based on old models, including at least, in part, the use of spreadsheets and macros, which are not ready to interface with Python-based models (or other advanced programming languages).
5. Lack of skills in the workforce to handle and assimilate AI/deep learning methodologies.

The main limitation to the wide adoption of innovative computer vision-based solar forecasting using deep learning can be summarized as a lack of coordination between providers and end users at the utility level (i.e., grid operators). This is a significant barrier that can, however, be overcome by putting effort into understanding the needs and processes of the user (i.e., grid operators) on the one side and developing the products with the users on the other side. Further, grid operators might need to upgrade their skills via dedicated training to gain the required knowledge to understand and use recent computer vision-based forecasting models.

In summary, there is a strong need for developers to interact with grid operators at the concept stage of the development of new products. From that context, some insight might come from a weather and climate service ecosystem, which, if developed, could provide an efficient provider/user interaction processes based on stakeholder engagement and co-production principles (co-design, co-development, co-evaluation, etc.) [438].

6. Conclusion

Solar energy meteorology using computer vision is essential to address the variability of solar generation caused by changing cloud cover, and thus facilitate its integration into the electric grid. In this context, machine learning offers a valuable modeling framework capable of analyzing and combining large amounts of complex data originating from Earth observations (i.e., remote-sensing), ground-based sky imagers, and *in situ* measurement devices. An extensive review of the resulting computer vision tasks and associated deep learning approaches for solar forecasting is presented in this article. Additionally, accessible resources, validation methods, and promising research topics — such as probabilistic modeling, video prediction, transfer learning, and explainable artificial intelligence — are thoroughly described. Despite several ongoing challenges, including improving the generalization skill of the models and facilitating the adoption of the techniques by the industry, this detailed analysis shows that machine learning has the capacity to improve the modeling of solar energy generation.

CRediT authorship contribution statement

Quentin Paletta: Conceptualization, Funding acquisition, Investigation, Project administration, Supervision, Visualization, Writing – original draft, Writing – review & editing. **Guillermo Terrén-Serrano:** Funding acquisition, Investigation, Project administration, Visualization, Writing – original draft, Writing – review & editing. **Yuhao**

Nie: Funding acquisition, Investigation, Visualization, Writing – original draft, Writing – review & editing. **Binghui Li:** Funding acquisition, Investigation, Writing – original draft. **Jacob Bieker:** Funding acquisition, Investigation, Writing – original draft, Writing – review & editing. **Wenqi Zhang:** Funding acquisition, Investigation, Writing – review & editing. **Laurent Dubus:** Funding acquisition, Investigation, Writing – original draft. **Soumyabrata Dev:** Funding acquisition, Investigation, Writing – original draft. **Cong Feng:** Conceptualization, Funding acquisition, Investigation, Project administration, Supervision, Visualization, Writing – original draft, Writing – review & editing.

Declaration of competing interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

Data availability

Data will be made available on request.

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