

Outer Van-Allen Radiation Belt Response to VLF Plasma Waves: Application of Granger Causality Method to Radiation Belt Storm Probe-A Data

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Research Article

Keywords: Granger causality analysis, Van-Allen radiation belts, very-low-frequency (VLF) waves, relativistic electrons flux, wave-particle interactions, radiation belt storm probes (RBSP) data.

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2 **Application of Granger Causality Method to Radiation Belt**
3 **Storm Probe-A Data**

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6 **Abstract**

7 Granger causality analysis is applied to very-low-frequency (VLF) waves and relativistic
8 electrons interaction in an outer Van-Allen radiation belt. The VLF wave power and electrons
9 flux data are one hour observations of radiation belt storm probe A (RBSP-A) during 14-3-
10 2015. The considered VLF band comprises of 24 waves and electrons flux data at 8 pitch
11 angles is considered. It is found that every wave is not Granger causing flux gradients for all
12 pitch angles; rather some VLF waves are more efficient than others. Specifically, flux
13 gradients at 69 degrees pitch angles are Granger caused by 14 kHz VLF wave. Contrarily, for
14 90 degrees pitch angles flux gradients are Granger caused by 6 waves. These results may be
15 helpful for understanding the dynamics of outer radiation belt.

16 **Key Words:** Granger causality analysis, Van-Allen radiation belts, very-low-frequency
17 (VLF) waves, relativistic electrons flux, wave-particle interactions, radiation belt storm
18 probes (RBSP) data.

19 **I. Introduction**

20 The plasma particles (electrons, protons) trapping requires gradients parallel to the direction
21 of magnetic field. This condition is naturally fulfilled by Earth's magnetic field, which is
22 dipole-like up to few L-shells. Here L-shell represents the geocentric distance of a point

23 where magnetic field line intersects geomagnetic equator (that is 11 degrees tilted with
24 respect to its geographic equator, for example $L=2.5$ means a point in magnetosphere that is
25 lying on magnetic equator and is at distance of 2.5 Earth's radii from centre of Earth). In
26 1958, the first space mission (Explorer spacecraft) discovered two regions of trapped plasma
27 in Earth's magnetosphere. These are known as Van-Allen radiation belts [1-2]. The inner
28 radiation belt exists between L -shell= 1.5 – 3 and is mostly dominated by protons having
29 energies 0.1 MeV – 40 MeV. The outer radiation belt occupies the region $4 < L$ -shell $\lesssim 6.0$
30 and its electrons have energies in keV – MeV range. These radiation belts are separated by a
31 slot region (L -shell=3-4), that is devoid of high intensity plasma particle radiations under
32 normal conditions [3-4]. Electron fluxes in the outer radiation belt fluctuate by several orders
33 of magnitude but its dynamics are yet not fully understood. Some intense storms can lead to
34 the formation of a temporary third radiation belt (also known as storage ring) in slot region,
35 which can survive for four weeks [5-6].

36 The space missions have confirmed the existence of many plasma waves in Earth's
37 magnetosphere [7-12]. These include very-low-frequency (VLF) waves, which are whistler
38 waves in frequency range 3-30 kHz [13-18], ultra-low-frequency (ULF) waves that are
39 Alfvén waves, electromagnetic ion cyclotron (EMIC) waves in vast frequency range (1Hz –
40 1mHz), their periods range from 1- 1000 seconds [19-20]. The VLF waves can be divided
41 into two categories: upper band chorus (having frequencies $\leq 0.5\omega_c$ (cyclotron frequency)),
42 and upper band chorus waves (with frequencies $> 0.5\omega_c$) [21].

43 The plasma particle sources of radiation belts include wave-particle interactions leading to
44 radially inward diffusion of electrons. Wave-particle interactions also act as loss mechanism
45 either by pitch angle scattering, or outward radial diffusion [22-29].

46 Many studies have focused on the wave-particle interactions in Van-Allen radiation belts and
47 third radiation belt formation in slot region [26-29]. But the studies so far have not yet
48 applied the Granger causality method to understand the relationship of VLF waves and outer
49 radiation belt electrons pitch angle resolved flux. Therefore, this study for the first time
50 applies Granger causality method to radiation belt storm probe A (RBSP) data of 24 VLF
51 waves and 8 pitch angles resolved electrons flux. Using Granger causality method, this work
52 finds answers to two basic research questions. The first research question states, do all 24
53 VLF waves in 10-30 kHz range have similar effects on outer radiation belt flux at pitch
54 angles 16-90 degrees? The second research question states that, can VLF waves in frequency
55 range 10-30 kHz non-resonantly affect the electrons flux? The answers for the first and
56 second research questions are no, and yes respectively. The manuscript is organized as
57 follows, section II is about data and methods. Section III describes the VLF plasma waves
58 and electrons flux observations of radiation belt storm probe A (RBSP-A) for one hour.
59 Section IV is devoted to the discussion of results by applying Granger Causality method. The
60 summary of main findings is presented in section V.

61

62 **II. Data and Methods**

63 The VLF waves and electrons flux data analysed in this work are one hour observations of
64 radiation belt storm probe A (RBSP-A) during 00:00:10.148 - 00:59:57.914, on 14-March-
65 2015. The electrons flux data at 2 MeV energy and varying pitch angles from 16-90 degrees
66 pitch angles and very-low-frequency (VLF) electric field power density data at 10 -30 kHz
67 frequencies is analysed.

68

69 The flux data is from relativistic electron proton telescope (REPT), basically from NASA
70 cdaweb (<https://cdaweb.gsfc.nasa.gov/>) at RBSP/ECT REPT pitch angle resolved electron

71 and proton fluxes with electron energies, 2-59.45 MeV with a logical source rbspa_rel03_ect-
72 rept-13. The time resolution of data values is varying from 10-11 seconds. The VLF waves
73 power data comes from radio and plasma wave instrument (EMFISIS-L2, electric and
74 magnetic field instrument level 2) with time resolution 6 seconds at <http://rbsp-ect.sr.unh.edu>.

75

76 The data file for Granger analysis is produced by selecting only those VLF power values
77 which are almost simultaneous with the flux data. The gradients of flux are calculated by
78 subtracting previous value of flux from its current value. A similar procedure is applied for
79 calculating power density gradients in the time series for each of 24 VLF waves. The initial
80 values of flux gradient and VLF power gradients are taken zero.

81

82 This work applies Granger causality tests to study the cause and effect relationship of 24
83 waves in very-low-frequency (VLF) band and 2 MeV relativistic electrons of outer Van-
84 Allen radiation at 8 types of pitch angles.

85

86 The Granger causality test is a statistical procedure of hypothesis testing. There are two types
87 of hypothesis: null and alternate. For this study, the null hypothesis states that temporal
88 gradients of electrons flux time series at given pitch angle are independent of the temporal
89 gradients in the power density time series of very-low-frequency electric field. Contrarily, the
90 alternate hypothesis states that the temporal gradients of electrons flux time series are caused
91 by the temporal gradients of VLF electric field power density gradients. Basically, there is
92 not a single procedure to determine Granger causality. Different methods can be adopted to
93 reject the null hypothesis and prove the causality of VLF waves and electrons flux. The
94 lengthy details are avoided [already presented in Refs. 30, 31] here and a brief description is
95 presented for readers. The causality judgement here is based on P-Value of F-statistics.

96 Basically, the formula of P-value reads as [30-31],

97
$$P - value = 1 - F^{cdf} \quad (1).$$

98 The superscript *cdf* in equation (1) indicates cumulative probability. However, F is

99 calculated by the following relationship [30-31],

100
$$F = \frac{\left[\frac{SSR_r - SSR_u}{N_p} \right]}{\frac{SSR_u}{(N_0 - N_p - 1)}} \quad (2)$$

101 Here SSR_r is sum square of residuals of the restricted model and SSR_u represent sum square
102 of residuals of the un-restricted model, N_p is the number of predictor variables and N_0 is
103 equal to the number of observations. This work takes a confidence interval of 95% which
104 translates to a critical value of 0.05. When equation (1) returns a P-value less than 0.05, then
105 null hypothesis is rejected and alternate hypothesis is accepted. Contrarily, when P-value is
106 greater than 0.05 then null hypotheses is accepted and alternate hypothesis is rejected. Using
107 this procedure the cause and effect relationships of the gradients in time series of VLF
108 electric field power density and gradients in the time series of electrons flux are determined
109 in figures (4)-(5) and explained in the discussion section IV.

110

111 **III. Results**

112 Figure 1 (a) displays the L-shell of radiation belt storm probe A with time. As the outer Van-
113 Allen radiation belt exists between L = 4-6 and the L-shell of RBSP-A changed from L=5.8
114 to 6.4, therefore, the observations are at the outer edge of outer Van-Allen radiation belt and
115 the findings of this study are more significant for understanding the relativistic electrons
116 dynamics in the presence of VLF wave turbulence in that region of Earth's magnetosphere
117 near geostationary orbit (at L=6.6).

118 The electrons in Van-Allen radiation belts can have three adiabatic invariants [33]. This study
119 does not concentrate on the first and third adiabatic invariant and the figure 1 (b) displays
120 second adiabatic invariant (SAI) for electrons. The SAI in figure 1(b) varied from 0.034 to
121 0.024. This indicates a 30 % reduction during the considered time interval. Based on this
122 observation, SAI is not conserved for electrons. Under such conditions one cannot apply the
123 adiabatic theory for finding the source of electrons flux changes. Although we have not
124 provided the figure of first adiabatic invariant (FAI), but a simple argument indicates that
125 FAI cannot be conserved. Because, the observed VLF waves frequencies (10-30 kHz) are
126 comparable to the cyclotron frequencies of electrons (18.87 - 25.84 kHz) in the region of
127 observations. This implies that the phase space density approach is also not applicable,
128 because, it is only valid for conserved first adiabatic (FAI). Consequently, a new approach,
129 Granger causality is adopted in this work to explore the influence of VLF waves on the
130 relativistic electrons flux variations.

131 The power densities in 24 waves of very low frequency band are presented in the figures 2
132 (a)-(c). In figure 2(a) power densities for 8 types of VLF waves at frequencies 10.0, 10.5,
133 11.0, 11.5, 12.1, 12.7, 13.3, 14.0 kHz are presented. It shows that the highest power densities
134 correspond to the 10 kHz waves and power density diminished as frequency increased,
135 smallest power densities occurred for 14 kHz waves. The figure 2(b) displays the power
136 densities for waves at 14.7, 15.4, 16.2, 17.0, 17.8, 18.7, 19.6, 20.5 kHz. Except 19.6, 20.5
137 these VLF waves have frequencies smaller than electron cyclotron and some of these can be
138 recognized as lower band chorus waves when the criteria $f < 0.5f_c$ is satisfied. Here f and
139 f_c denote wave frequency and an electron cyclotron frequency respectively. The VLF waves
140 in figure 2(c) at frequencies 21.5, 22.6, 23.7, 24.9, 26.1, 27.4, 28.7, 30.1 kHz are upper band
141 chorus waves, because these satisfy the criteria $f > 0.5f_c$.

142 Figure 3 (a) shows the scatter plot of the 2 MeV relativistic electrons flux at 8 pitch angles,
143 16, 26, 37, 47, 58, 69, 79, and 90 degrees. The flux is decreasing with the time for electrons
144 at these considered pitch angles. To highlight extra information, figure 3 (b) presents a
145 contour plot of the 2 MeV relativistic electrons with time and pitch angles. The highest flux
146 peaks occurred for 37 degrees and 47 degrees pitch angles. Temporal trend shows that flux at
147 all pitch angles is diminishing with time. The presence of VLF waves in figures 2 and flux
148 changes in figure (3) can be linked together. One can assume that flux changes in figure (3)
149 are due to VLF wave interactions with electrons. Such possibilities are explained in the
150 following section.

151

152 **IV. Discussion of Results:**

153 Radiation belt storm probe A (RBSP-A) simultaneously observed VLF waves (figure 2) and
154 relativistic electrons flux reduction (figure 3), this implies that the flux changes are triggered
155 by VLF waves. However, there are some basic research questions, such as, are the electron
156 flux changes at 16 degrees pitch angles are caused by all the 24 waves or by only one wave?
157 Similar, research questions can be asked for electrons flux changes at other pitch angles and
158 combinations of waves. The RBSP-A observations in figures (2) and (3) cannot provide
159 answers to such research questions. Therefore, further analysis of data in figures (2) and (3) is
160 required. This work applies the Granger causality tests to explore the roles of considered VLF
161 waves. The critical value 0.05 is used for judgement. When P-value is smaller than 0.05, then
162 we reject the null hypothesis and alternate hypothesis is accepted. Otherwise, when P-value is
163 greater than 0.05, then null hypothesis is accepted and alternate hypothesis is rejected. In this
164 study, the null hypothesis states, the electron flux gradients at pitch angle (= 16, 26, 37, 47,
165 58, 69, 79, 90 degrees) are independent of the gradients in powers of VLF wave (10, 10.5,

166 11.0, 11.5, 12.1, 12.7, 13.3, 14.0, 14.7, 15.4, 16.2, 17.0, 17.8, 18.7, 19.6, 20.5, 21.5, 22.6,
167 23.7, 24.9, 26.1, 27.4, 28.7, 30.1 kHz). The alternate hypothesis states that flux gradients are
168 dependent on the gradients of the wave power. These hypotheses are tested with Granger
169 causality test and relationships between waves and flux changes at considered frequencies,
170 and pitch angles are determined. Figure 4 shows P-values of Granger causality tests at 4-lags
171 for the 2 MeV electrons at pitch angles 58 degrees (4a), 69 degrees (4b), 79 degrees (4c), and
172 90 degrees (4d). In figures 4 (a) –(d) the horizontal red line shows a critical level of 0.05 and
173 any point that lies under this level leads to rejection of null hypotheses and all points above it,
174 accepts the null hypothesis. The P-values in figure 4(a) show that 8 waves Granger caused the
175 flux changes for 58 degrees pitch angle electrons. Contrarily, gradients in power of only one
176 wave (at 14 kHz) Granger caused (resulted in P-values less than 0.05 for first and second lag)
177 flux gradients at 69 degrees pitch angles, as shown in figure 4(b). The flux changes in figure
178 4 (c) for 79 degrees pitch angle electrons are Granger caused by two waves. Figure 4 (d)
179 shows that power gradients of 6 waves Granger caused flux variations of 90 degrees pitch
180 angle electrons. Figure 5 shows that the flux gradients of 16 (figure 5a), 26 (figure 5b), 37
181 (figure 5c), and 47 degrees (figure 5d) pitch angle electrons are Granger caused by 2, 5, 3,
182 and 5 waves respectively. The results in figures (4) and (5) also suggest that VLF waves also
183 interact non-resonantly with outer Van-Allen radiation belt electrons, because, all these
184 waves does not satisfy the cyclotron resonance condition. For example, consider the case of
185 90 degrees pitch angle in figure 4 (d). In this case, 18.87 kHz VLF wave frequency is
186 comparable to electron cyclotron frequency (18.87 - 25.84 kHz) in region of observation,
187 therefore, the flux changes due 18.87 kHz VLF wave can be recognized as cyclotron resonant
188 effect. Because, the cyclotron resonance condition reads [32-33],

189
$$\omega - kV_{parallel} = n\omega_c \quad (3).$$

190 For 90 degrees pitch angle, parallel velocity $V_{Parll} = 0$, and resonance condition (1) reduces
191 to $\omega = \omega_c$ when $n = 1$. The effects of 20.5, 23.7 kHz waves can also be attributed to
192 cyclotron resonance for $n = 1$ because, these waves are bounded in the domain of electron
193 cyclotron frequencies (18.87 - 25.84 kHz). However, the VLF waves at 12.1, 27.4 kHz and
194 28.7 kHz does not satisfy the resonance condition for any $n (= \pm 1, \pm 2, \pm 3, \dots)$ and therefore,
195 these waves non-resonantly Granger caused flux gradients for the 90 degrees pitch angle
196 electrons.

197 **V. Summary and Conclusions:**

198 The radiation belt storm probe (RBSP-A) simultaneous observations of 24 VLF waves and
199 diminishing fluxes of 2 MeV electrons at 8 pitch angles suggest that flux changes may be due
200 to VLF waves. By applying, the Granger causality tests, we find an answer to the basic
201 research question, does all 24 VLF waves equally contributed to electrons flux changes? The
202 answer is that all these waves do not contribute equally to electron flux changes in outer
203 radiation belt at the considered pitch angles. This work also answers another research
204 question, does VLF wave only Granger caused flux gradients when cyclotron resonance
205 condition is satisfied. It is found that in some cases flux changes are Granger caused by VLF
206 waves satisfying cyclotron resonance condition but not always. This implies that VLF waves
207 also interacted non-resonantly with electrons and caused flux changes. The findings of this
208 study can be helpful for understanding the dynamics of outer radiation belt.

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210 measured electrons flux data. We are also thankful to C. Kletzing for EMFISIS-L2 data of
211 VLF wave electric field power density data.

212 **Data availability statement:** The data is available upon request to corresponding author at
213 masifkf@gmail.com.

214 **Conflict of Interest:** The author has no conflict of interest.

215 **References**

216 ¹Van Allen, J. A., & Frank, L. A. Radiation Around the Earth to a Radial
217 Distance of 107,400 km. *Nature* **183**, 430-434 (1959).

218 ²Van Allen, J. A. Radiation belts around the Earth. *Scientific American*, **200**, 39-
219 47 (1959).

220 ³Summers, D., Mann, I. R., Baker, D. N., & Schulz, M. G. (Eds.).
221 (2013). *Dynamics of the Earth's radiation belts and inner magnetosphere* (Vol.
222 199). John Wiley & Sons.

223 ⁴Friedel, R. H. W., G. D. Reeves, and T. Obara. Relativistic electron dynamics in
224 the inner magnetosphere—A review. *Journal of Atmospheric and Solar-*
225 *Terrestrial Physics* **64**, 265-282 (2002).

226 ⁵Baker, D.N., Kanekal, S.G., Hoxie, V.C., Henderson, M.G., Li, X., Spence,
227 H.E., Elkington, S.R., Friedel, R.H.W., Goldstein, J., Hudson, M.K. and Reeves,
228 G.D., 2013. A long-lived relativistic electron storage ring embedded in Earth's
229 outer Van Allen belt. *Science*, *340*(6129), pp.186-190.

230 ⁶Shprits, Y.Y., Subbotin, D., Drozdov, A., Usanova, M.E., Kellerman, A.,
231 Orlova, K., Baker, D.N., Turner, D.L. and Kim, K.C., 2013. Unusual stable
232 trapping of the ultrarelativistic electrons in the Van Allen radiation belts. *Nature*
233 *Physics*, *9*(11), pp.699-703.

234 ⁷LaBelle, J. and Treumann, R.A., 1988. Plasma waves at the dayside
235 magnetopause. *Space Science Reviews*, *47*(1), pp.175-202.

236 ⁸Shawhan, S.D., 1979. Magnetospheric plasma wave research 1975–
237 1978. *Reviews of Geophysics*, *17*(4), pp.705-724.

238 ⁹Shawhan, S.D., 1977. *Magnetospheric plasma waves* (No. TRITA-EPP--77-07).
239 Kungliga Tekniska Hogskolan.

240 ¹⁰Yumoto, K., 1988. External and internal sources of low-frequency MHD waves
241 in the magnetosphere-A review. *Journal of geomagnetism and*
242 *geoelectricity*, 40(3), pp.293-311.

243 ¹¹Amatucci, W.E., 2006. A review of laboratory investigations of space plasma
244 waves. *URSI Radio Science Bulletin*, 2006(319), pp.32-66.

245 ¹²Lanzerotti, L.J., Fukunishi, H., MacLennan, C.G. and Cahill Jr, L.J., 1976.
246 Observations of magnetohydrodynamic waves on the ground and on a
247 satellite. *Journal of Geophysical Research*, 81(25), pp.4537-4545.

248 ¹³Hashimoto, Kozo, Isamu Nagano, Masayuki Yamamoto, Toshimi Okada, Iwane
249 Kimura, Hiroshi Matsumoto, and Hidetaka Oki. "EXOS-D (AKEBONO) very
250 low frequency plasma wave instruments (VLF)." *IEEE transactions on*
251 *geoscience and remote sensing* 35, no. 2 (1997): 278-286.

252 ¹⁴Cohen, M. (2013, May). Very Low Frequency Remote Sensing of the
253 Ionosphere and Magnetosphere. In *AGU Spring Meeting Abstracts* (Vol. 2013,
254 pp. SM24A-01).

255 ¹⁵Kaiser, T. R., D. Orr, and A. J. Smith. "Very low frequency electromagnetic
256 phenomena:'whistlers' and micropulsations." *Philosophical Transactions of the*
257 *Royal Society of London. B, Biological Sciences* 279, no. 963 (1977): 225-238.

258 ¹⁶Záhlava, J., F. Němec, Jean-Louis Pincon, O. Santolík, I. Kolmašová, and
259 Michel Parrot. "Whistler influence on the overall very low frequency wave
260 intensity in the upper ionosphere." *Journal of Geophysical Research: Space*
261 *Physics* 123, no. 7 (2018): 5648-5660.

262 ¹⁷Sazhin, S. S. "VLF emissions in the earth's magnetosphere." *Geomagnitnye*
263 *Issledovaniia* 18 (1976): 24-53.

264 ¹⁸Kulkarni, V. H., and Jharana Das. "Very low frequency (VLF) chorus
265 emissions: A topical survey." *Surveys in geophysics* 13, no. 1 (1992): 35-46.

266 ¹⁹Dungey, J. W., & Southwood, D. J. (1970). Ultra-low frequency waves in the
267 magnetosphere. *Space Science Reviews*, 10(5), 672-688.

268 ²⁰Menk, F. W. (2011). Magnetospheric ULF waves: A review. *The dynamic*
269 *magnetosphere*, 223-256.

270 ²¹Sazhin, S. S., and M. Hayakawa. "Magnetospheric chorus emissions: A
271 review." *Planetary and space science* 40, no. 5 (1992): 681-697.

272 ²²Rae, I.J., Murphy, K.R., Watt, C.E., Halford, A.J., Mann, I.R., Ozeke, L.G.,
273 Sibeck, D.G., Clilverd, M.A., Rodger, C.J., Degeling, A.W. and Forsyth, C.,
274 2018. The role of localized compressional ultra-low frequency waves in energetic
275 electron precipitation. *Journal of Geophysical Research: Space Physics*, 123(3),
276 pp.1900-1914.

277 ²³Su, Z., Zhu, H., Xiao, F., Zong, Q.G., Zhou, X.Z., Zheng, H., Wang, Y., Wang,
278 S., Hao, Y.X., Gao, Z. and He, Z., 2015. Ultra-low-frequency wave-driven
279 diffusion of radiation belt relativistic electrons. *Nature communications*, 6(1),
280 pp.1-8.

281 ²⁴Rankin, R., Wang, C.R., Wang, Y.F., Zong, Q., Zhou, X.Z., Degeling, A.W.,
282 Sydorenko, D. and Whittall-Scherfee, G., 2020. Ultra-Low-Frequency Wave–
283 Particle Interactions in Earth's Outer Radiation Belt. *Dayside Magnetosphere*
284 *Interactions*, pp.189-205.

285 ²⁵Zong, Q., Hao, Y. and Wang, Y., 2009. Ultra low frequency waves impact on
286 radiation belt energetic particles. *Science in China Series E: Technological*
287 *Sciences*, 52(12), pp.3698-3708.

288 ²⁶Thorne, R.M., 2010. Radiation belt dynamics: The importance of wave-particle
289 interactions. *Geophysical Research Letters*, 37(22).

290 ²⁷Watt, C.E., Rae, I.J., Murphy, K.R., Anekallu, C., Bentley, S.N. and Forsyth,
291 C., 2017. The parameterization of wave-particle interactions in the Outer
292 Radiation Belt. *Journal of Geophysical Research: Space Physics*, 122(9),
293 pp.9545-9551.

294 ²⁸Xiao, F., Liu, S., Tao, X., Su, Z., Zhou, Q., Yang, C., He, Z., He, Y., Gao, Z.,
295 Baker, D.N. and Spence, H.E., 2017. Generation of extremely low frequency
296 chorus in Van Allen radiation belts. *Journal of Geophysical Research: Space*
297 *Physics*, 122(3), pp.3201-3211.

298 ²⁹Allison, H.J., Shprits, Y.Y., Zhelavskaya, I.S., Wang, D. and Smirnov, A.G.,
299 2021. Gyroresonant wave-particle interactions with chorus waves during extreme
300 depletions of plasma density in the Van Allen radiation belts. *Science*
301 *advances*, 7(5), p.eabc0380.

302 ³⁰Shojaie, Ali, and Emily B. Fox. "Granger causality: A review and recent
303 advances." *Annual Review of Statistics and Its Application* 9 (2022): 289-319.

304 ³¹Faes, L., Nollo, G., Stramaglia, S., & Marinazzo, D. (2017). Multiscale granger
305 causality. *Physical Review E*, 96(4), 042150.

306 ³²Litvak, A.G., Sergeev, A.M., Suvorov, E.V., Tokman, M.D. and Khazanov,
307 I.V., 1993. On nonlinear effects in electron-cyclotron resonance plasma heating

308 by microwave radiation. *Physics of Fluids B: Plasma Physics*, 5(12), pp.4347-
309 4359.

310 ³³Baumjohann, W., & Treumann, R. A. (2012). *Basic space plasma physics (revised*
311 *edition)*. World Scientific Publishing Company.

312 **Figure Captions**

313 **Figure 1** shows (a) L-shell of radiation belt storm probe A (RBSP-A) during one hour
314 observations (00:00:10.148 - 00:59:57.914) on 14-March-2015. (b) Exhibits temporal
315 variations in the electrons second adiabatic invariant.

316 **Figure 2** shows very- low-frequency (VLF) waves electric field power densities for (a) eight
317 waves at 10.0, 10.5, 11.0, 11.5, 12.1, 12.7, 13.3, 14.0 kHz (b) eight waves at 14.7, 15.4,
318 16.2, 17.0, 17.8, 18.7, 19.6, 20.5 kHz (c) eight waves at 21.5, 22.6, 23.7, 24.9, 26.1, 27.4,
319 28.7, 30.1 kHz. These are observations of RBSP-A for same time interval as in figure 1.

320 **Figure 3** shows (a) scatter plot of 2 MeV energy electrons pitch angle resolved flux at eight
321 pitch angles, 16, 26, 37, 47, 58, 69, 79, 90 degrees respectively (b) displays contour plot of
322 the 2 MeV electrons flux with pitch angle and time. The flux measurements are based on the
323 relativistic electron proton telescope (REPT) on board RBSP-A during one hour observations
324 (00:00:10.148 - 00:59:57.914) on 14-March-2015.

325 **Figure 4** shows the P-values of Granger causality test, determined for gradients in VLF
326 waves powers and gradients of electrons flux at pitch angles (a) 58 degrees (b) 69 degrees (c)
327 79 degrees, and (d) 90 degrees. When P-value is less than critical value 0.05 (highlighted
328 with red horizontal line) then null hypothesis is rejected and flux gradients are Granger
329 caused by power gradients of VLF waves. Otherwise, null hypothesis is accepted and flux
330 gradients are thought to be independent of power gradients of VLF wave.

331 **Figure 5** shows the P-values of Granger causality test determined for gradients in VLF waves
332 powers and gradients of electrons flux at pitch angles (a) 16 degrees (b) 26 degrees (c) 37
333 degrees, and (d) 47 degrees.

334

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339

Figure 1

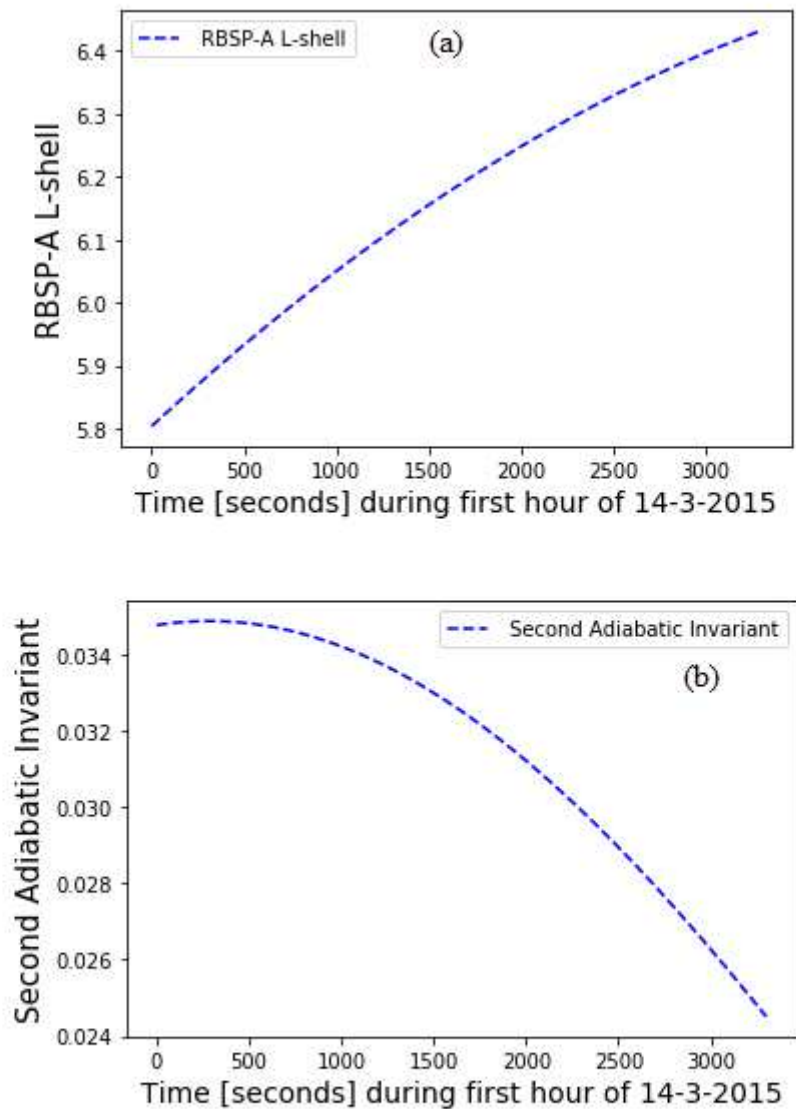


Figure 2

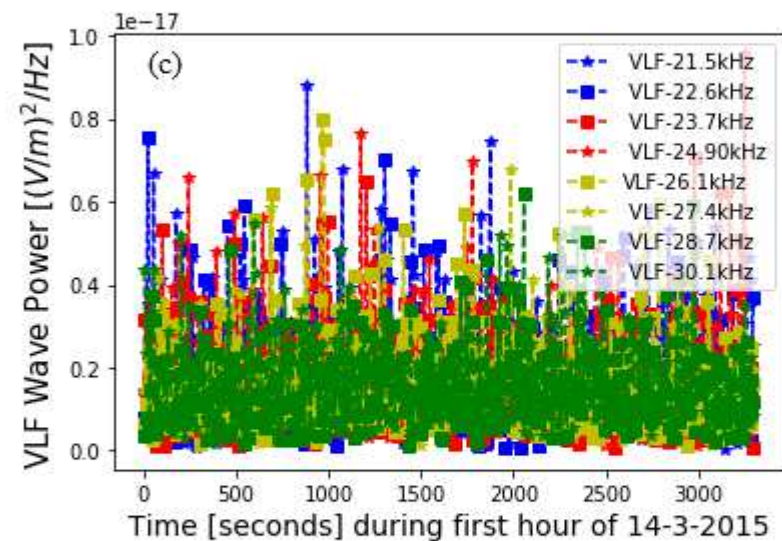
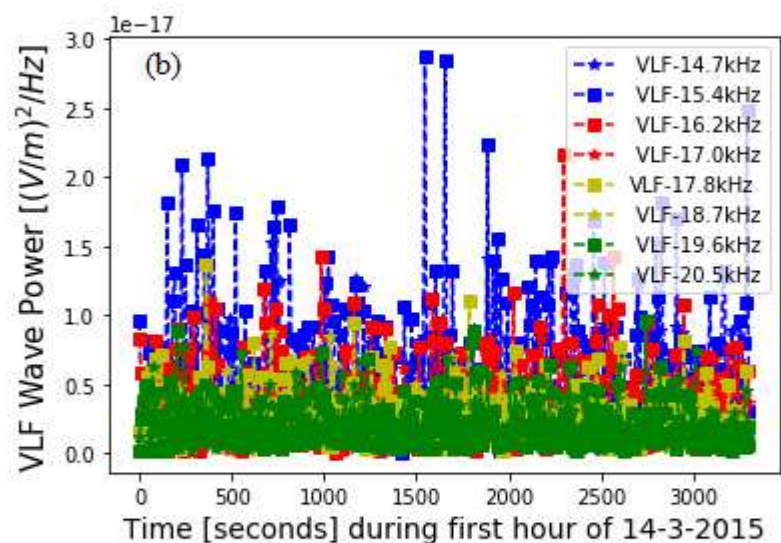
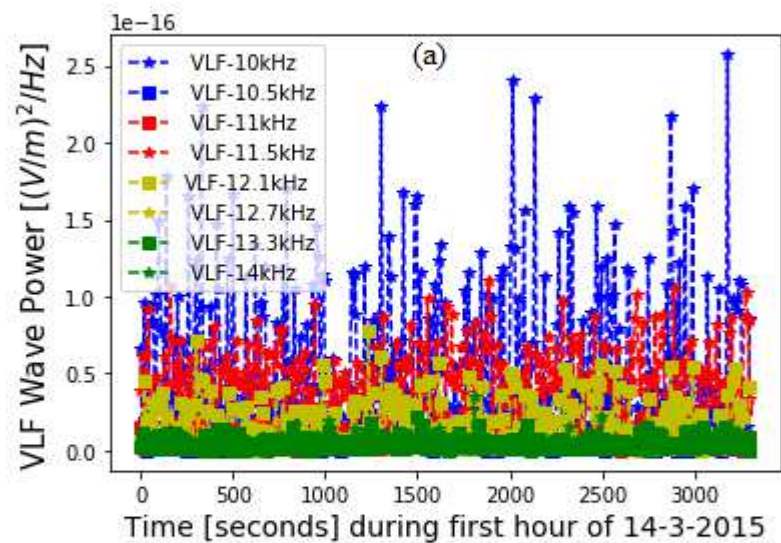


Figure 3

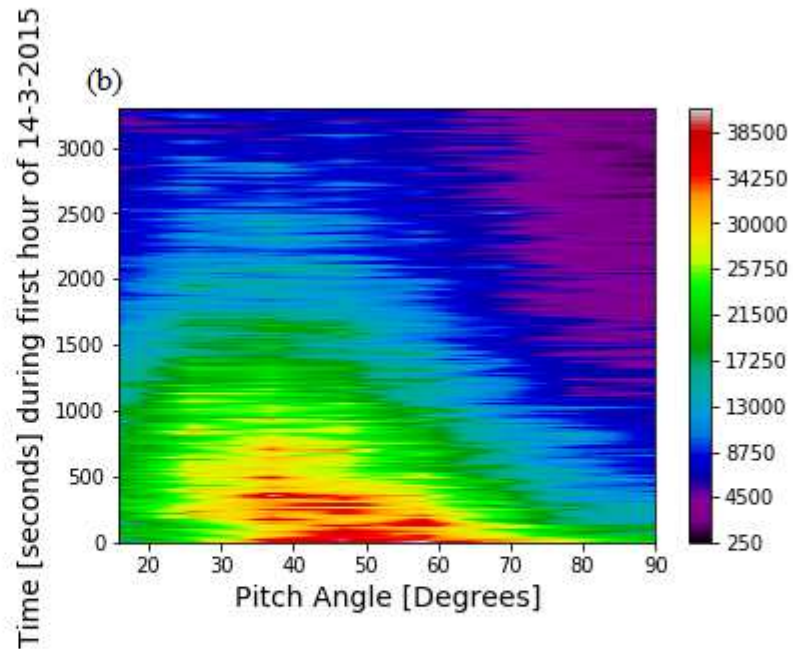
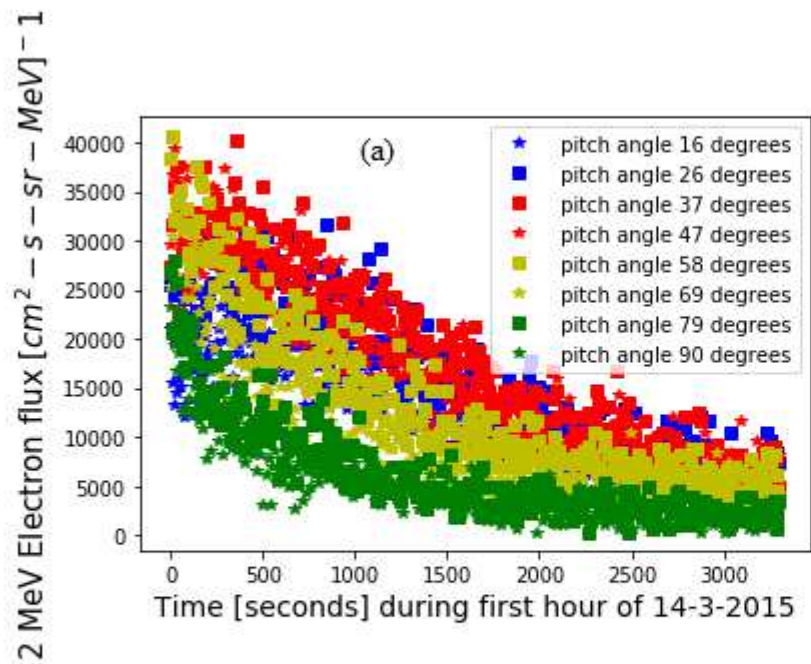


Figure 4

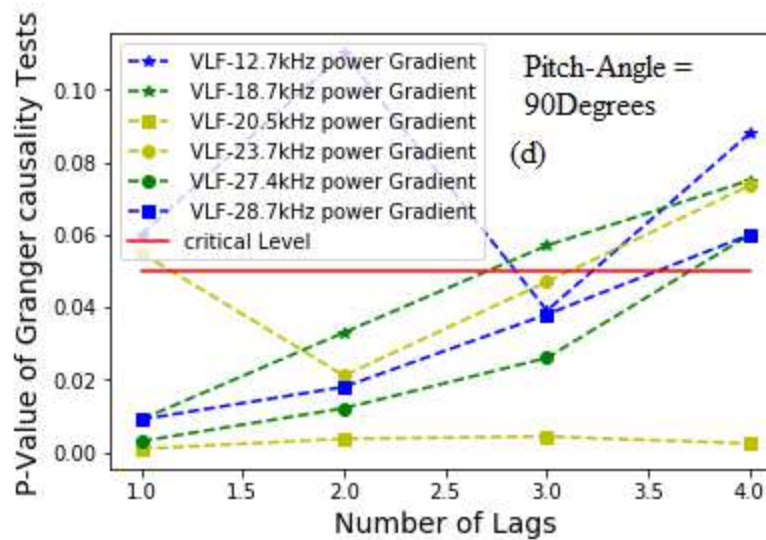
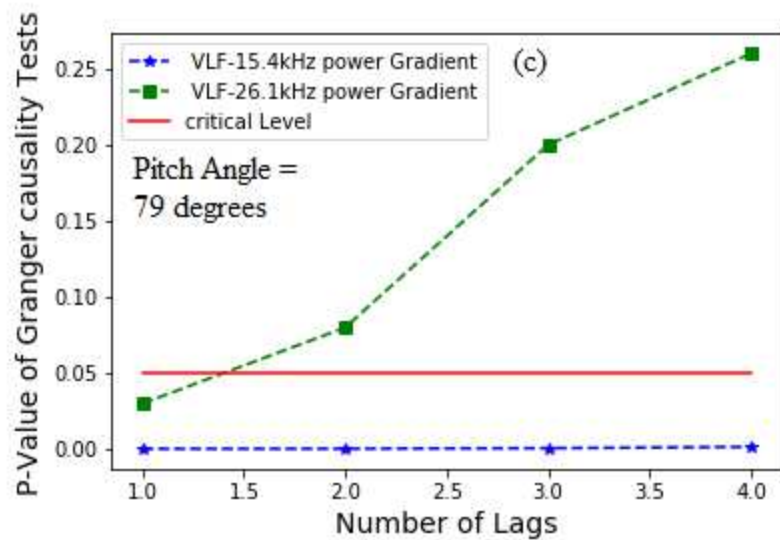
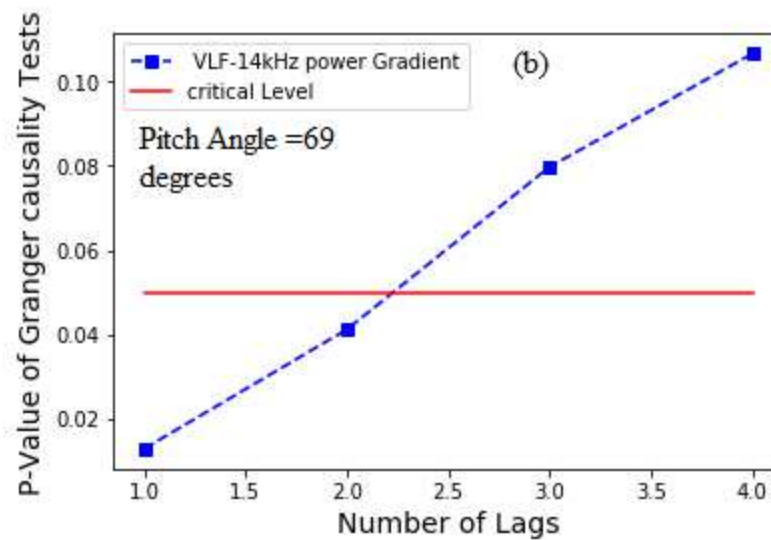
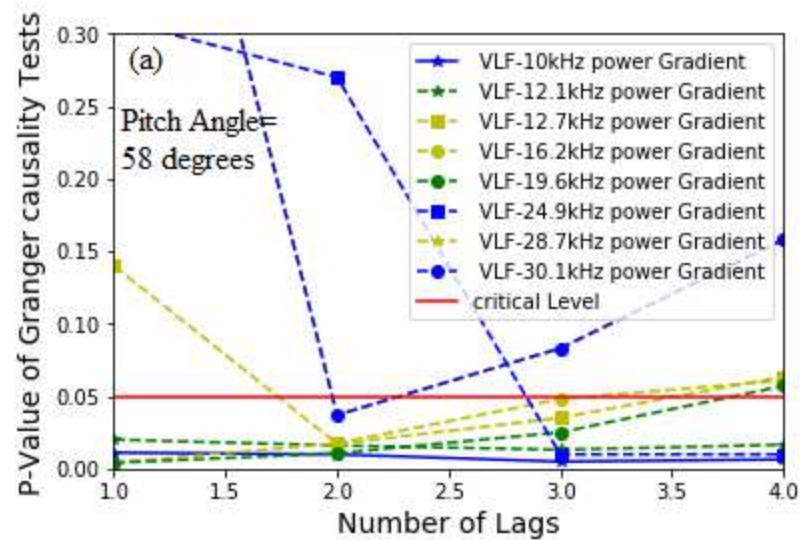


Figure 5

