Antimatter in the lab

Lecture 2 Jack Devlin CERN 25/7/2023 Recap

Lecture 1

What is antimatter? 1.

2. Why study antimatter?



- $\Psi_{2} = A \begin{pmatrix} 1 \\ 0 \\ \frac{-p}{E_{\mu} + m} \end{pmatrix} e^{-ip^{\mu}x_{\mu}} \qquad \Psi_{4} = A \begin{pmatrix} 0 \\ \frac{p}{E_{p} + m} \\ 0 \\ 1 \end{pmatrix} e^{-ip^{\mu}x_{\mu}}$
- 3. How do we make antimatter at CERN?





In a magnetic field

e-



Figure adapted from "The Essential Cosmic Perspective", by Bennett, Voit, and Donahue

Overview

Lecture 2: Experiments at the Antimatter Factory

- 1. Catching and storing antimatter
- 2. Measurements on antiprotons
- 3. Spectroscopy of anti-atoms
- 4. Gravity and antihydrogen
- 5. Taking antimatter out of the lab (and into another lab..)



Low energy physics

- High energy experiments (direct search)
 - Produce and detect new particles in high energy particle collisions
 - High energies needed to make non-virtual particles

p

• Precision experiments: (indirect search)

$$\frac{g_{electron}}{2} = 1 + a_{QED} + a_{\mu,\tau} + a_{weak} + a_{hadrons} + \frac{a_{New \, physics}}{2}$$

So far, all low energy CPT tests have been consistent with no CPT breaking



Other efforts

Many groups studying antimatter. Some notable ones are:

Positron-electron magnetic moment University of Washington/Harvard, new effort at Northwestern Van Dyke& Dehmelt / Gabrielse Gabrielse



Reidar Hahn

Muon-antimuon magnetic moment Fermilab, formerly Brookhaven

Positronium, muonium, bound states of electrons/positron and antimuon+ electron: many groups

Kaons, B mesons, other collider searches

D. Hanneke et al. Phys. Rev. Lett. **100** (2008)
R.S. Van Dyke, Jr., P.B. Schwinberg, and H.G. Dehmelt, Phys. Rev. Lett. **59**, 26 (1987)
G.W. Bennett et al., Phys. Rev. Lett. **89**, 101804 (2002)
B. Abi et al. (Muon g–2 Collaboration) Phys. Rev. Lett. **126**, (2021)



Why many efforts?

Minimal Standard Model Extension (SME)

$\mathcal{L}^{ ext{CPT-even}}_{ ext{lepton}} = rac{1}{2} i (c_L)_{\mu u AB} \overline{L}_A \gamma^\mu \stackrel{ ightarrow}{D^ u} L_B$	$\mathcal{L}_{\text{gauge}}^{\text{CPT-even}} = -\frac{1}{2} (k_G)_{\kappa\lambda\mu u} \text{Tr}(G^{\kappa\lambda}G^{\mu u})$
$+\frac{1}{2}i(c_R)_{\mu\nu AB}\overline{R}_A\gamma^{\mu}\overrightarrow{D}^{\nu}R_B , \qquad (9)$	$-rac{1}{2}(k_W)_{\kappa\lambda\mu u}{ m Tr}(W^{\kappa\lambda}W^{\mu u})$
	$-\frac{1}{4}(k_B)_{\kappa\lambda\mu\nu}B^{\kappa\lambda}B^{\mu\nu} . \tag{16}$
${\cal L}^{ m CPT-odd}_{ m lepton} = -(a_L)_{\mu AB} \overline{L}_A \gamma^\mu L_B - (a_R)_{\mu AB} \overline{R}_A \gamma^\mu R_B ,$	
(10)	$\mathcal{L}_{\text{gauge}}^{\text{CPT-odd}} = (k_3)_{\kappa} \epsilon^{\kappa \lambda \mu \nu} \text{Tr}(G_{\lambda} G_{\mu \nu} + \frac{2}{3} i g_3 G_{\lambda} G_{\mu} G_{\nu}) + (k_2)_{\kappa} \epsilon^{\kappa \lambda \mu \nu} \text{Tr}(W_{\lambda} W_{\mu \nu} + \frac{2}{3} i g W_{\lambda} W_{\mu} W_{\nu})$
$\mathcal{L}^{ ext{CPT-even}}_{ ext{quark}} = rac{1}{2} i (c_Q)_{\mu uAB} \overline{Q}_A \gamma^\mu \stackrel{ ightarrow}{D^ u} Q_B$	$+(k_1)_{\kappa}\epsilon^{\kappa\lambda\mu\nu}B_{\lambda}B_{\mu\nu}+(k_0)_{\kappa}B^{\kappa} . \tag{17}$
$+rac{1}{2}i(c_U)_{\mu u AB}\overline{U}_A\gamma^\mu \stackrel{ ightarrow}{D^ u} U_B$	
$+\frac{1}{2}i(c_D)_{\mu\nu AB}\overline{D}_A\gamma^{\mu}\stackrel{\leftrightarrow}{D^{\nu}}D_B , \qquad (11)$	
$\mathcal{L}_{ ext{quark}}^{ ext{CPT-odd}} = -(a_Q)_{\mu AB} \overline{Q}_A \gamma^\mu Q_B - (a_U)_{\mu AB} \overline{U}_A \gamma^\mu U_B$	
$-(a_D)_{\mu AB}\overline{D}_A\gamma^{\mu}D_B . \tag{12}$	
$\mathcal{L}_{ ext{Yukawa}}^{ ext{CPT-even}} = -rac{1}{2} ig[(H_L)_{\mu u AB} \overline{L}_A \phi \sigma^{\mu u} R_B$	
$+(H_U)_{\mu u AB}\overline{Q}_A\phi^c\sigma^{\mu u}U_B$	
$+(H_D)_{\mu\nu AB}\overline{Q}_A\phi\sigma^{\mu\nu}D_B] + \text{h.c.} $ (13)	
$\mathcal{L}_{ ext{Higgs}}^{ ext{CPT-even}} = rac{1}{2} (k_{\phi\phi})^{\mu u} (D_{\mu}\phi)^{\dagger} D_{ u}\phi + ext{h.c.}$	
$-rac{1}{2}(k_{\phi B})^{\mu u}\phi^{\dagger}\phi B_{\mu u}$	
$-\frac{1}{2}(k_{\phi W})^{\mu\nu}\phi^{\dagger}W_{\mu\nu}\phi , \tag{14}$	
$\mathcal{L}_{\text{Higgs}}^{\text{CPT-odd}} = i(k_{\phi})^{\mu} \phi^{\dagger} D_{\mu} \phi + \text{h.c.} . \tag{15}$	

Non-minimal Standard Model Extension (SME)

18 pages to write down

Don't know a priori where to look!

CPT breaking coefficients proportional to (energy)^(-order)

Measurement	Energy scale	Fractional precision	Measurement in energy units
$K_0 - \overline{K}_0$ mass difference	Mass of two Kaons ~1 GeV	4.8x10 ⁻¹⁹	4.8x10 ⁻¹⁹ GeV
\overline{H} 1S-2S	~2500 THz	2x10 ⁻¹²	2x10 ⁻²⁰ GeV
$ar{p}$ magnetic moment	Larmor frequency ~81 MHz at 1.95 T	1.5x10 ⁻⁹	5x10 ⁻²⁵ GeV

Catching and storing antimatter

Final step of slowing

Final ELENA energy 100 keV, need a final step to reach trappable energies

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Catching antiprotons

Strong magnetic and moderate electric fields used



Measurements on antiprotons

Properties

Measurements

CPT Tests



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Frequency measurements in a Penning trap



Voltages applied to ring-shaped electrodes





Axial $v_z = 674 \text{ kHz}$ numbers

Modified cyclotron

 $v_{+} = 28.9 \text{ MHz}$

 $\sqrt{\nu_z^2 + \nu_+^2 + \nu_-^2} = \nu_c = \frac{q}{2\pi m} B_e$

Orbit is sum of three normal modes

Measure frequencies and get access to charge-to-mass ratio and magnetic field

Charge-to-mass ratio comparisons



Gravity

Clock comparison between matter and antimatter clocks in a gravitational potential

Relies on assumptions about CPT



b.)

Measuring the magnetic moment



Weak radiofrequency magnetic field most likely to flip antiproton spin at ω_L



G. Schneider et al., Science. 358 6366 (2017)

Method



Measure ω_L and ω_c and spin states in the same trap

Works well for electrons, but large quadratic B field adds temperature broadening, limits measurements



Double trap method

Separate spin state identification from measuring ω_L and ω_c



H Häffner et al., The European Physical Journal D **22** 2 (2003) DiSciacca, J. & Gabrielse, G. *Phys. Rev. Lett.* **108**, 153001 (2012) H. Nagahama et al. Nature Communications **8** (2017) Hanneke et al., Phys. Rev. Lett. **100** (2008)

Van Dyck, R.S., Jr.; Schwinberg, P.B.; Dehmelt, H.G. Phys. Rev. Lett. 59 (1987)

Spectroscopy of anti-atoms

What about the positrons?

Need positrons to make antihydrogen



Radioactive source and moderation with frozen noble gas





~ 5 million slow e+ per second





(CERN/GBAR/Comini)

~ 40 million slow e+ per second e- accelerated 10 MeV onto a water-cooled tungsten target to form positrons by pair production, moderated by tungsten mesh

Producing antihydrogen

1) Recombination $\overline{p} + e^+ \rightarrow \overline{H} + \text{UV photon}$ 2) Three body recombination $\overline{p} + e^+ + e^+ \rightarrow \overline{H} + e^+$

rate 2)>> rate 1) typically



ALPHA: 2.6 ± 0.2 detected \overline{H} trapped per minute

3) Charge transfer \bar{p} + positronium $\rightarrow \bar{H} + e^{-}$



AEGIS: 0.021(5) \overline{H} per attempt, ~15 minutes per attempt

M. Ahmadi, Nature Communications **8** (2017) C. Amsler et al., Nature Communications Physics **4** (2021) F Robicheaux J. Phys. B: At. Mol. Opt. Phys. **41** (2008) <u>arXiv:2306.15801</u> [hep-ex] (2023)

Antihydrogen production in ALPHA



(anti)hydrogen



Analytically calculable energy levels, high precision hydrogen measurements (4.5 ppt for 1S-2S) to compare to antihydrogen

C. G. Parthey, et al. Phys. Rev. Lett. 107, 203001 (2011)

Think like a precision measurer



What properties to we want in a transition?

What do we want to control about the atom?

What do we want to control about the environment?

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Laser spectroscopy of Antihydrogen 1S-2S



ALPHA collaboration

60 hr antihydrogen storage

 $f_{d-d} = 2,466,061,103,079.4(5.4)$ kHz (measured) $f_{d-d} = 2,466,061,103,080.3(0.6)$ kHz (predicted)



2 ppt

M. Ahmadi, Nature 557, 71-75 (2018) C. J Baker, Nature 592, 35–42 (2021)

RF spectroscopy hyperfine splitting



Lamb shift $2S_{1/2} - 2P_{1/2} \overline{H} 0.99 \pm 0.11$ GHz, H 0.9098717(32) GHz Ground state splitting: \overline{H} 1,420.4 ± 0.5 MHz, H 1 420 405.751 766 7(9)



2.7 ppb in Hydrogen

M. Ahmadi, Nature **548**, 66-69 (2017) M. Ahmadi, Nature **578**, 375–380 (2020) ASACUSA Collaboration. Report No. SPSC-P-307 Add. 1 CERN-SPSC-2005-002 (2005) N. Ramsey, Hyp. Interactions **81** (1993)

Antiprotonic helium



Effect of gravity on antihydrogen

Types of measurement





Freefall

Clock comparison

High precision needed- If proton is any guide, antiquark masses only ~1% of the antiproton





GBAR



- 1. Make \overline{H}^+ by reacting $\overline{H}^- + e^+$
- 2. Sympathetically laser cool \overline{H}^+ with Be⁺ to 20 μ K
- 3. Photoionise \overline{H}^+ to \overline{H}
- 4. Drop and measure effect of gravity

Initial goal 1%, final goal 0.1% with quantum measurement

Also, long term possibility to produce \overline{H}_2^+ molecules and do antimolecular spectroscopy!



D. P. van der Werf, Antimatter and Gravity (WAG 2013) **30** (2014) GBAR status report CERN-SPSC-2022-003 / SPSC-SR-302 (2022) arXiv:2306.15801 [hep-ex] AEgIS



Kellerbauer et al., NIM B 226 3 (2008)

A.

Taking antimatter out of the lab

PUMA

Take 10^9 antiprotons from the antimatter factory to ISOLDE

Some nuclei at the limits of N>Z have a neutron halo a where one or more neutrons are found far outside the nucleus. Other's have neutron skins, where the density of neutrons is larger than

protons at the nuclear surface – antiproton annihilation study these effects









Nuclear density

p annihilation
 <R> p annihilation
 <R> nucleon removal

BASE-STEP







Antimatter bombs: Could antimatter weaponry wipe out all life on Earth? Expert weighs in Wed 21 Jul 2021

NO!

Professor Robson said: "The idea you can produce masses of antimatter and make a bomb from it or something is not realistic. It's not something anyone needs to worry about."

One billion antiprotons annihilating

 $E = mc^2 = (2 \times 1 \times 10^9 \times 1.6 \times 10^{-27})c^2 = 1 \text{ nJ}$

1nJ can heat 3 picogram of water from 20->100 °C



About enough to make an espresso for an amoeba

(2015). "An update on *Acanthamoeba* keratitis: diagnosis, pathogenesis and treatment". *Parasite* **22**: 10.



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And thank you for listening

Storing antiprotons

BASE holds record for antiprotons stored from 03.11.2015 – 22.12.2016





- Storage of antiprotons for more than one year: **405.5 days**
- Extraction of single particles by a potential tweezer scheme

Inversion of the baryon asymmetry:

Antibaryon density: ~ $10^8/cm^3 V < (50 \ \mu m)^3$ Baryon density: ~ $1 \ / \ cm^3 p < 10^{-16} Pa$

> C. Smorra et al., Int. J. Mass Spectr. **389**, 10 (2015). S. Sellner et al., New J. Phys. 19, 083023 (2017).