#### Digital Photography I

optics and sensors



#### CSC2529

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\*slides adapted from Gordon Wetzstein, Fredo Durand, Ioannis Gkioulekas, Marc Levoy

#### Announcements

- HW 1 is due Wednesday at 11:59pm
- HW 2 is out (due next Wednesday 4/10)
- Instructor office hours today 4-5pm BA 7228
- TA office hours Tues 1-2pm BA 5256

# Let's say we have a sensor...

digital sensor (CCD or CMOS) real-world object

digital sensor (CCD or CMOS)

What would an image taken like this look like?



digital sensor (CCD or CMOS)

digital sensor (CCD or CMOS)

real-world object

digital sensor (CCD or CMOS)

real-world object

digital sensor (CCD or CMOS)

real-world object

All scene points contribute to all sensor pixels

What does the image on the sensor look like?



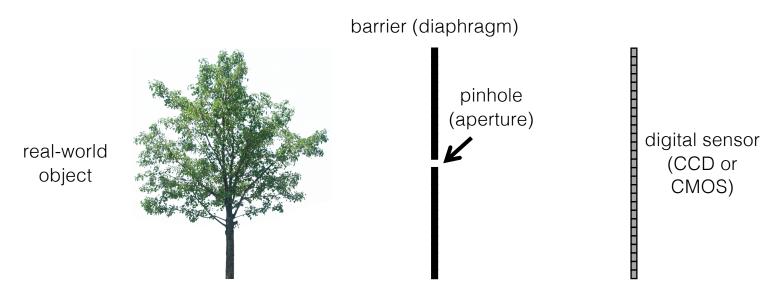
All scene points contribute to all sensor pixels

#### What can we do to make our image look better?

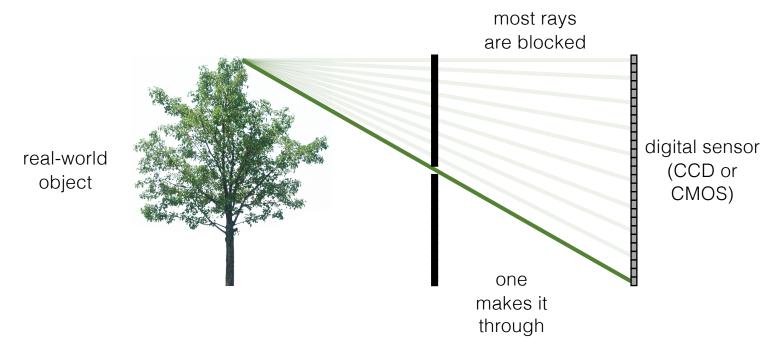


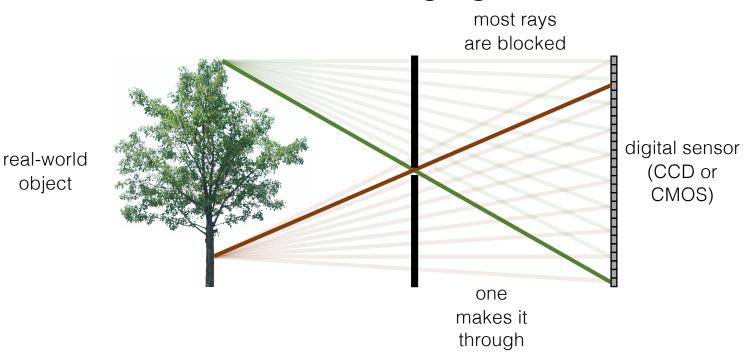
digital sensor (CCD or CMOS)

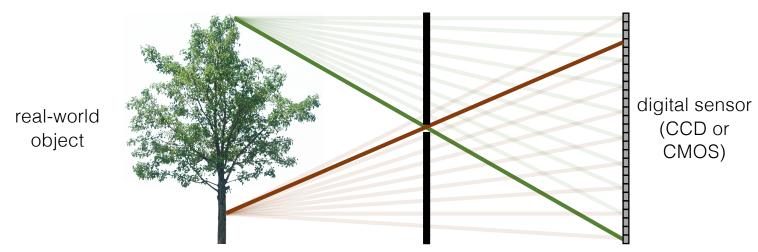
### Let's add something to this scene



What would an image taken like this look like?

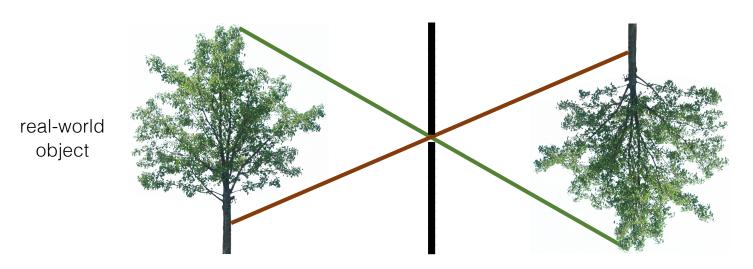






Each scene point contributes to only one sensor pixel

What does the image on the sensor look like?



copy of real-world object (inverted and scaled)

#### Pinhole camera terms

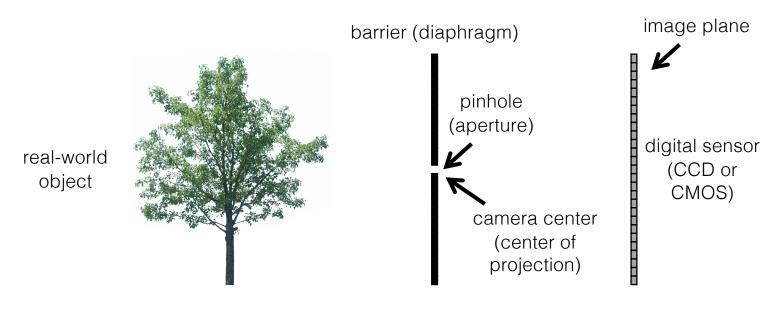
real-world object

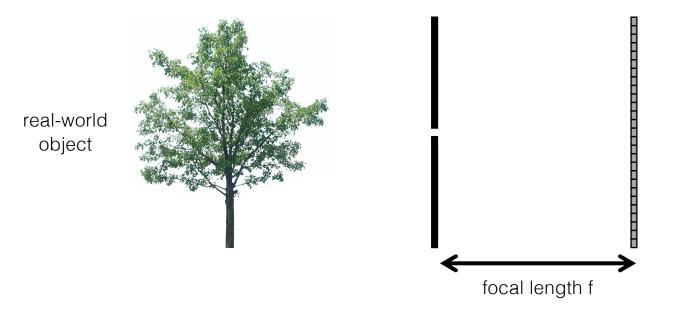
barrier (diaphragm)

pinhole (aperture)

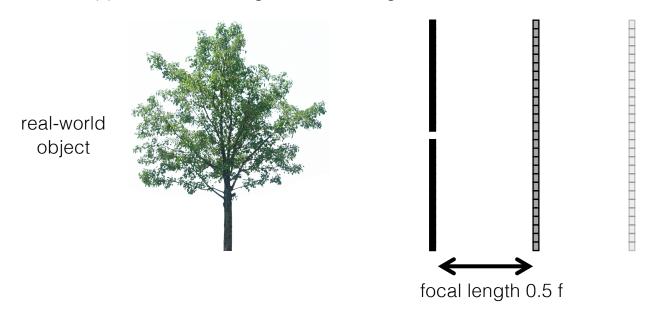
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#### Pinhole camera terms

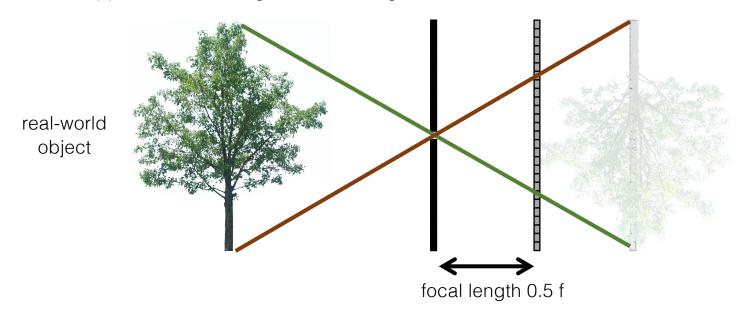


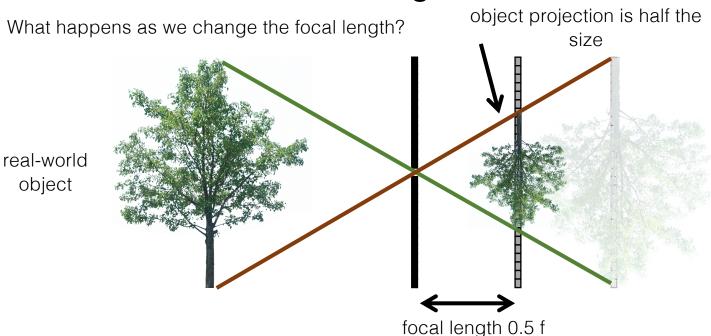


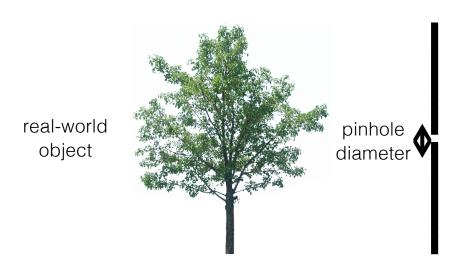
What happens as we change the focal length?



What happens as we change the focal length?



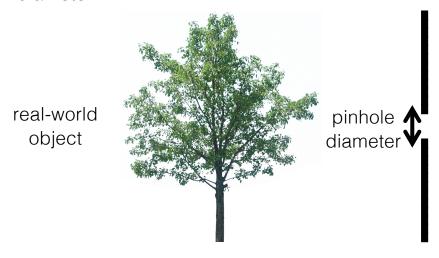




Ideal pinhole has infinitesimally small size

In practice that is impossible.

What happens as we change the pinhole diameter?

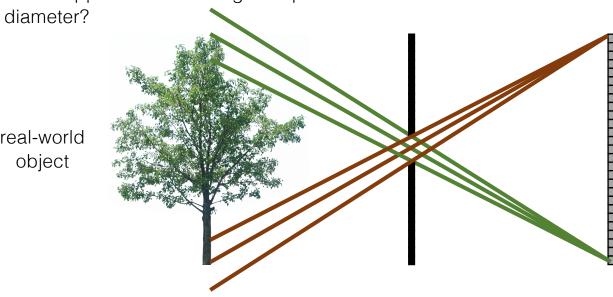


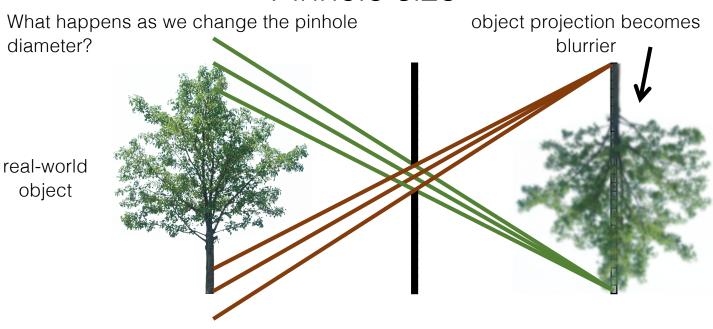
What happens as we change the pinhole diameter?

real-world object

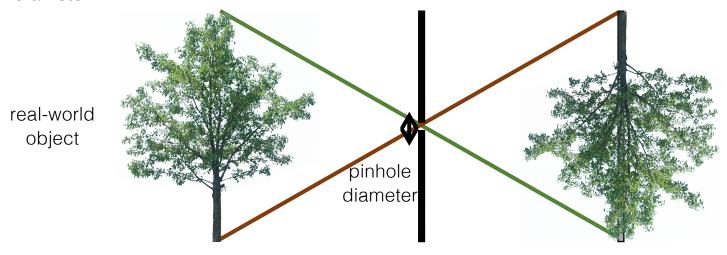
What happens as we change the pinhole

real-world object



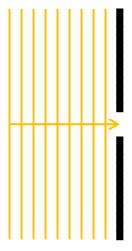


What happens as we change the pinhole diameter?

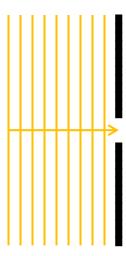


Will the image keep getting sharper the smaller we make the pinhole?

A consequence of the wave nature of light

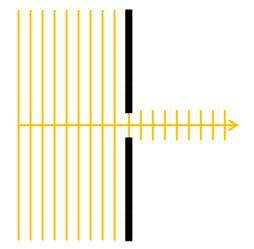


What do geometric optics predict will happen?

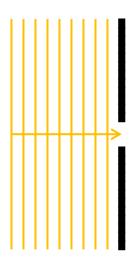


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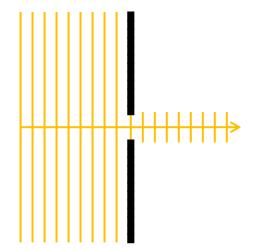


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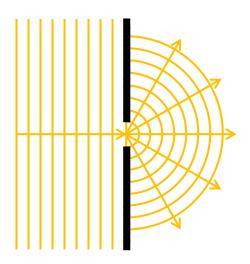


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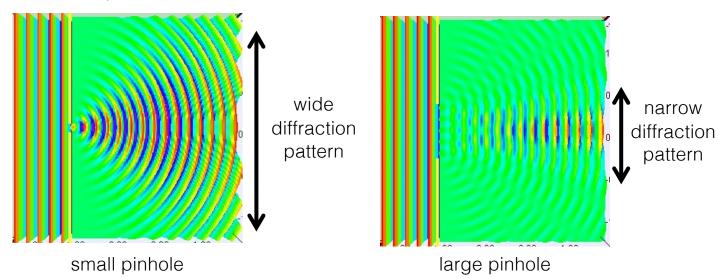
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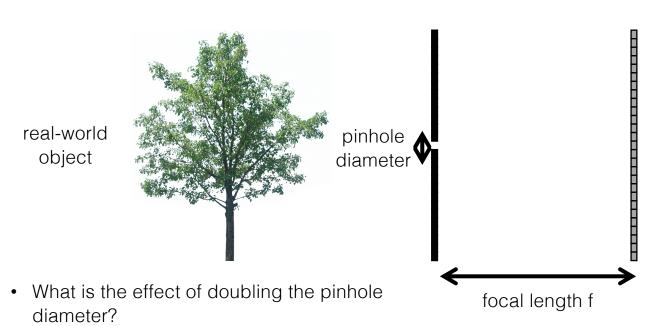
What do wave optics predict will happen?

Diffraction pattern = Fourier transform of the pinhole.

- Smaller pinhole means bigger Fourier spectrum.
- Smaller pinhole means more diffraction.

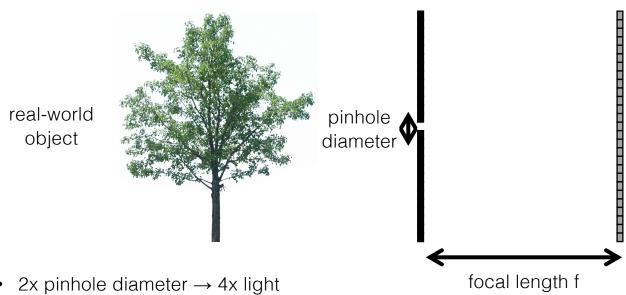


# What about light efficiency?



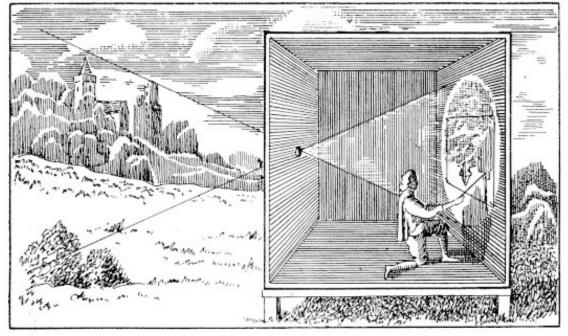
What is the effect of doubling the focal length?

# What about light efficiency?

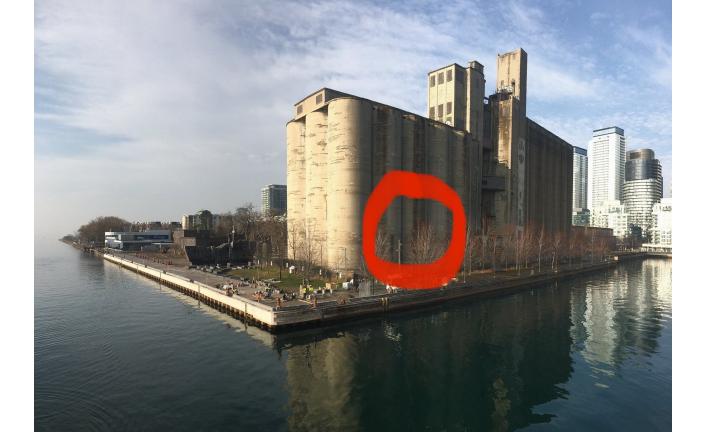


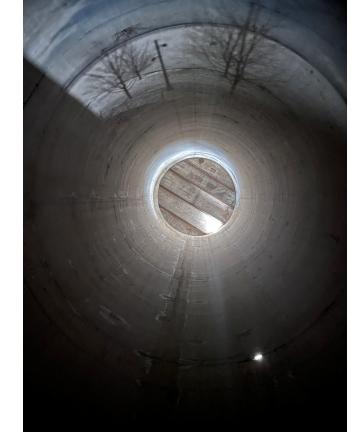
•  $2x \text{ focal length} \rightarrow \frac{1}{4}x \text{ light}$ 

#### Pinhole Camera / Camera Obscura



Mo-Ti (Chinese Philosopher) 470-390 BC







Fun discovery - a small crack in the eastern facade of the Canada Malting Co silos has created a perfect pinhole camera. The result: real time projection of Toronto's waterfront on the silo's interior curved surfaces. An unplugged projection show!



9:37 AM · Jan 27, 2022 · Twitter for iPhone

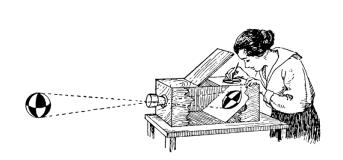
656 Retweets	<b>70</b> Quote Tweets	<b>2,836</b> Likes		
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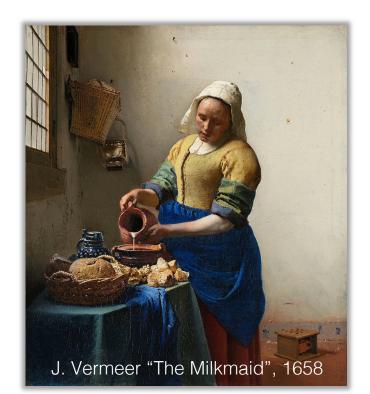






## Pinhole Camera / Camera Obscura





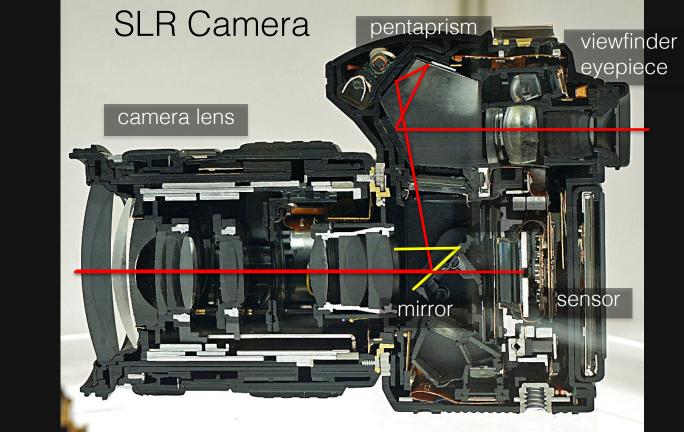


©Rijksmuseum Amsterdam. Purchased with the support of the Vereniging Rembrandt

# Digital Photography - Overview

- optics
- aperture
- · depth of field
- field of view
- exposure
- noise
- color filter arrays
- image processing pipeline





# Camera Optics



1826 8h exp

## Daguerrotype





- invented in 1836 by Louis Daguerre
- lenses focus light, better chemicals!



exposure 10-12 mins

#### Lenses

- focus light
- magnify objects



Nimrud lens - 2700 years old

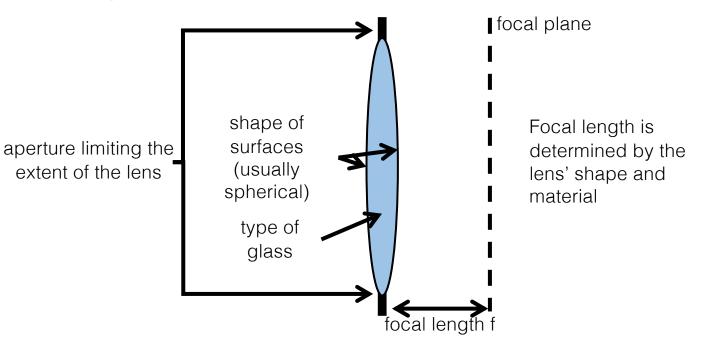
## What is a lens?

A piece of glass manufactured to have a specific shape

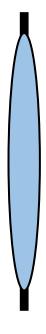


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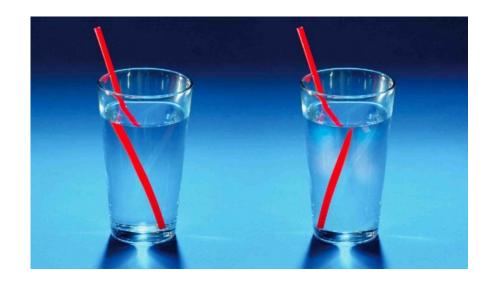


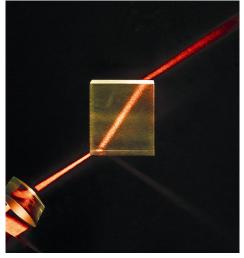
## How does a lens work?



## Refraction

Refraction is the bending of rays of light when they move from one material to another





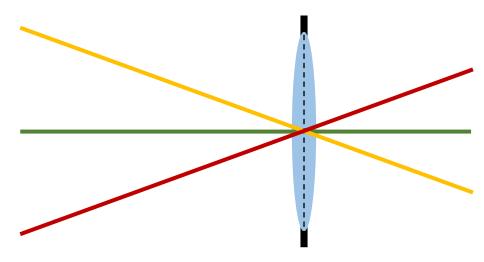
## How does a lens work?

Lenses are designed so that their refraction makes light rays bend in a very specific way.

Simplification of geometric optics for well-designed lenses.



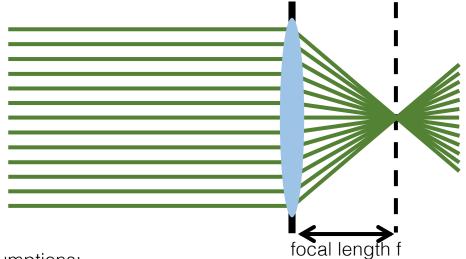
Simplification of geometric optics for well-designed lenses.



#### Two assumptions:

1. Rays passing through lens center are unaffected.

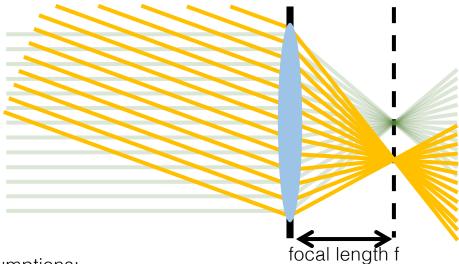
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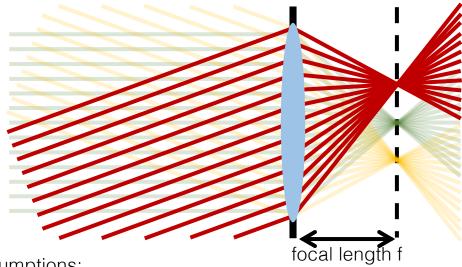
- 1. Rays passing through lens center are unaffected.
- 2. Parallel rays converge to a single point located on focal plane.

Simplification of geometric optics for well-designed lenses.



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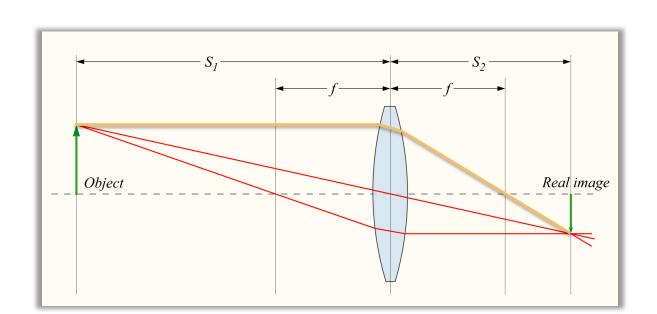


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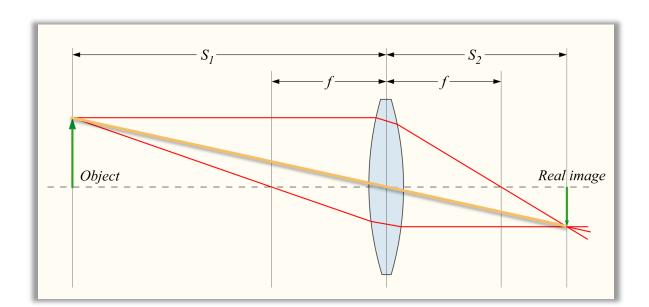
#### Ray tracing example

Parallel rays map to the focal plane



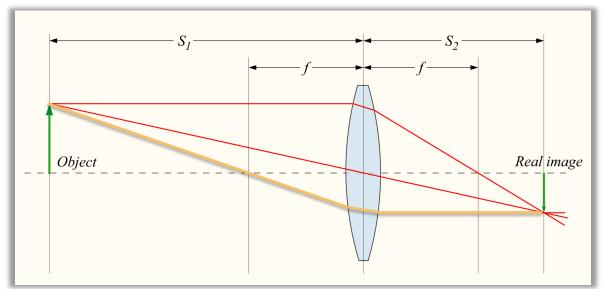
#### Ray tracing example

- Parallel rays map to the focal plane
- The chief ray passes straight through the center

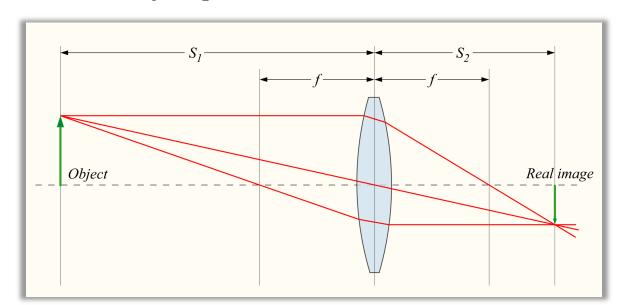


#### Ray tracing example

- Parallel rays map to the focal plane
- The chief ray passes straight through the center
- The ray that passes through the near focal plane becomes parallel

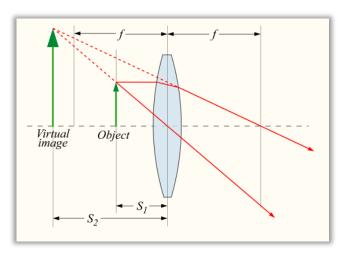


Thin lens 
$$\frac{1}{f} = \frac{1}{S_1} + \frac{1}{S_2}$$
 magnification:  $M = -\frac{S_2}{S_1} = \frac{f}{f - S_1}$ 



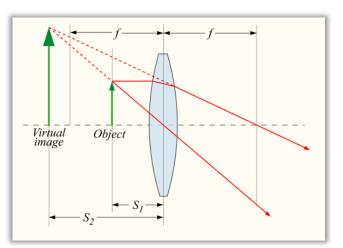
## Lenses

S1<f: magnifying glass

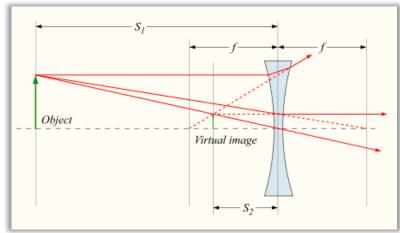


## Lenses

S1<f: magnifying glass



#### minification



Yes, but...

#### Thin lenses are a fiction

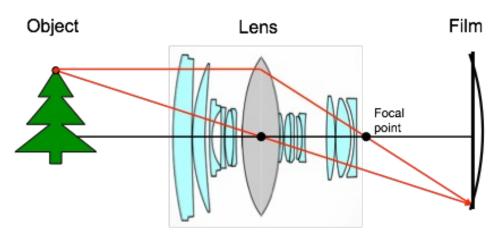
The thin lens model assumes that the lens has no thickness, but this is rarely true...



To make real lenses behave like ideal thin lenses, we have to use combinations of multiple lens elements (compound lenses).

#### Thin lenses are a fiction

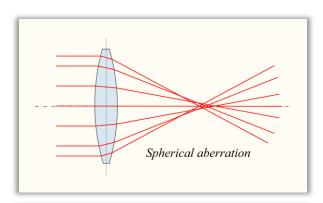
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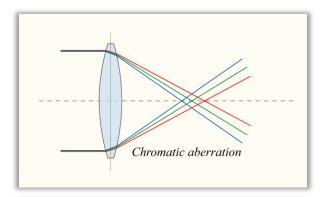


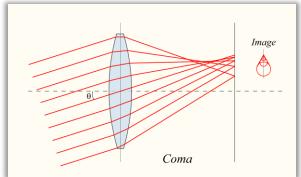
 Even though we have multiple lenses, the entire optical system can be (paraxially) described using a single thin lens of some equivalent focal length and aperture number.

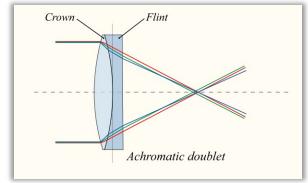
To make real lenses behave like ideal thin lenses, we have to use combinations of multiple lens elements (compound lenses).

## Lenses - Aberrations

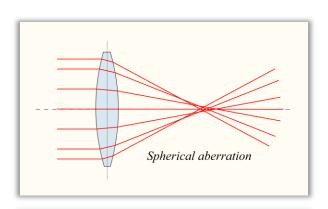


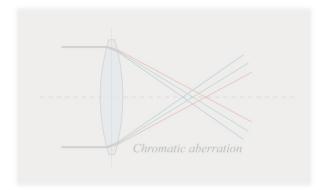


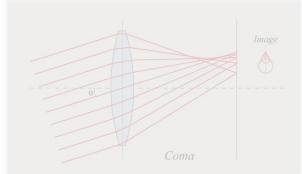


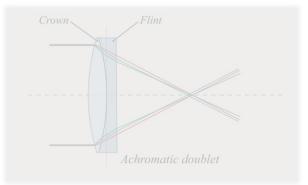


## Lenses - Aberrations



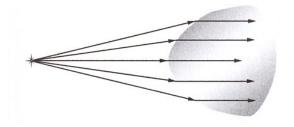






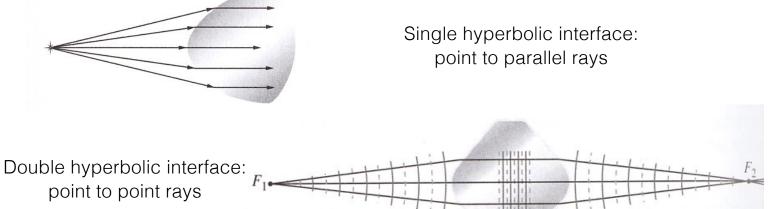
# Refraction at interfaces of complicated shapes

What shape should an interface have to make parallel rays converge to a point?



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What shape should an interface have to make parallel rays converge to a point?



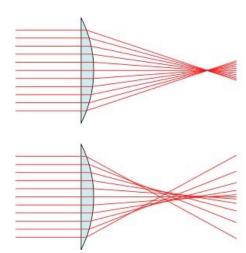
Therefore, lenses should also have hyperbolic shapes.

# Spherical lenses

In practice, lenses are often made to have spherical interfaces for ease of fabrication.

Two roughly fitting curved surfaces ground together will eventually become spherical.





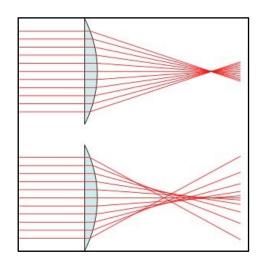
Spherical lenses don't bring parallel rays to a point.

- This is called spherical aberration.
- Approximately axial (i.e., paraxial) rays behave better.

#### Aberrations

Deviations from ideal thin lens behavior (e.g., imperfect focus).

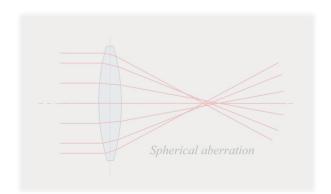
• Example: spherical aberration.

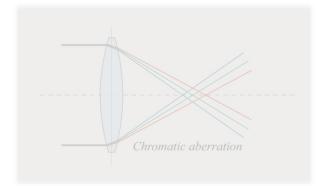


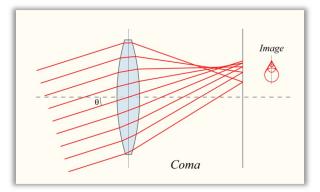


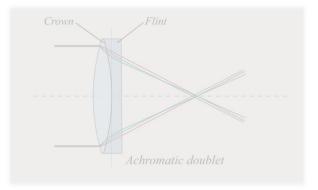


#### Lenses - Aberrations





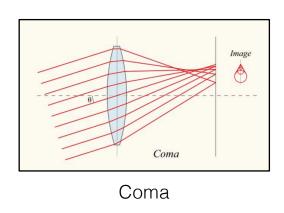


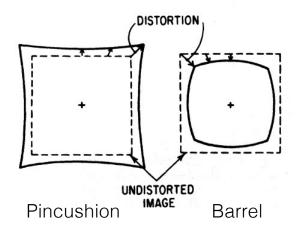


#### Oblique aberrations

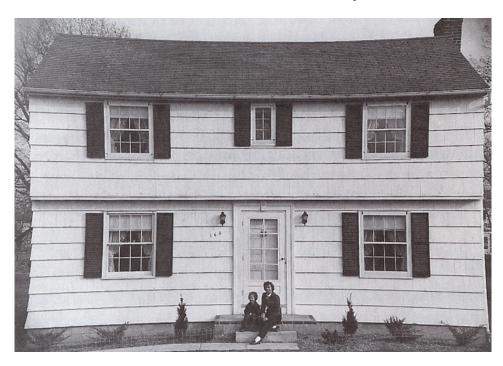
These appear only as we move further from the center of the field of view.

- Contrast with spherical and chromatic, which appear everywhere.
- Many other examples (astigmatism, field curvature, etc.).

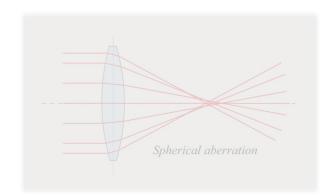


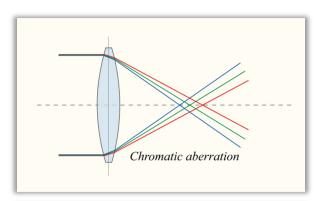


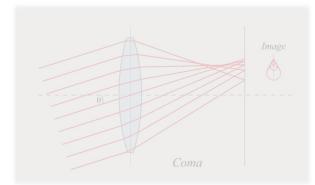
#### Distortion example

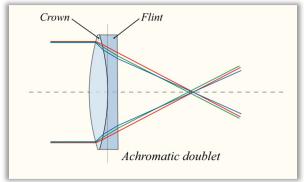


#### Lenses - Aberrations







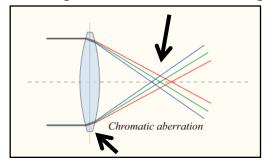


#### **Aberrations**

Deviations from ideal thin lens behavior (e.g., imperfect focus).

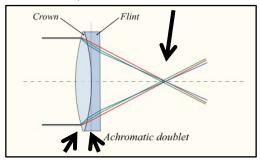
• Example: chromatic aberration.

focal length shifts with wavelength



glass has dispersion (refractive index changes with wavelength)

one lens cancels out dispersion of other



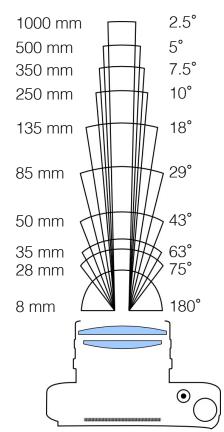
glasses of different refractive index

Using a doublet (two-element compound lens), we can reduce chromatic aberration.

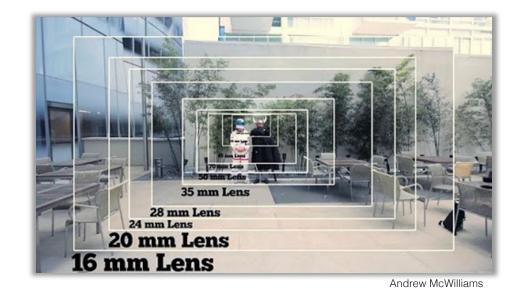
#### Chromatic aberration examples

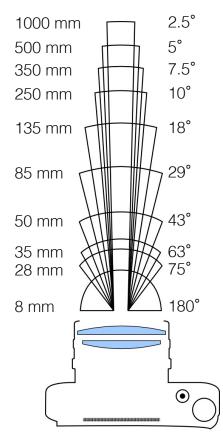






#### Field of View





#### Field of View

Hubble – what's the focal length?

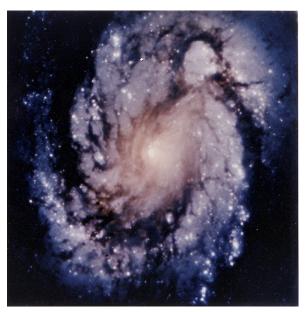


#### A costly aberration

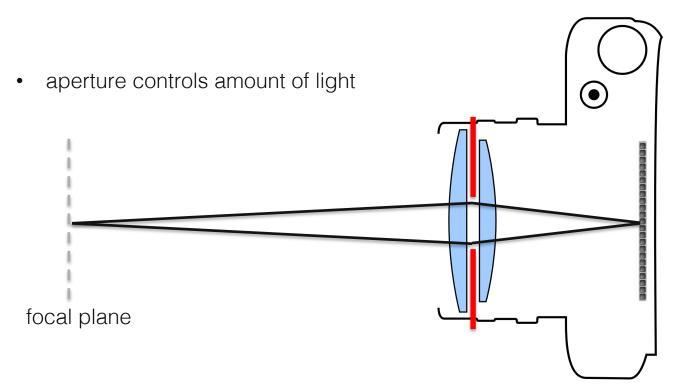
Hubble telescope originally suffered from severe spherical aberration.

• COSTAR mission inserted optics to correct the aberration.

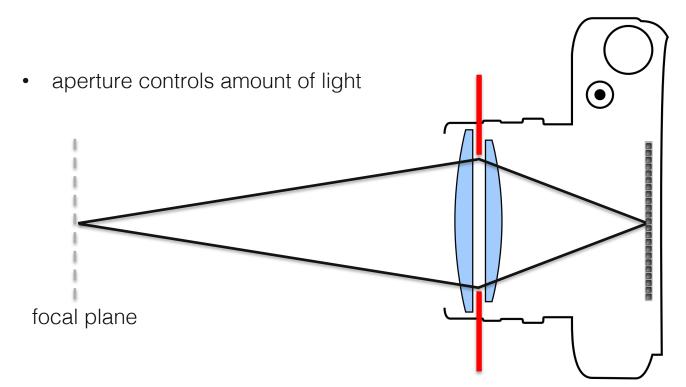




#### Aperture



#### Aperture



#### Aperture size

Most lenses have variable aperture size.

- F-number notation: "f/1.4" means f / = 1.4 (focal length / diameter).
- Usually aperture sizes available at steps of one-half or one-third stops.
- Older lenses have separate manual aperture ring.
- Modern lenses control the aperture through a dial on the camera body ("gelded" lenses).



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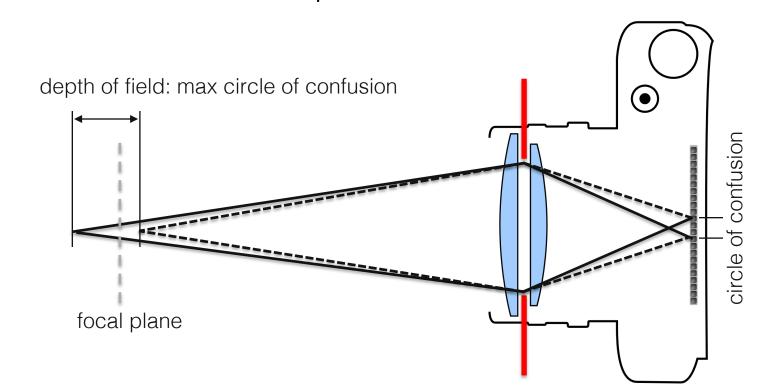


Reminder: A "stop" changes the amount of light by a factor of 2.

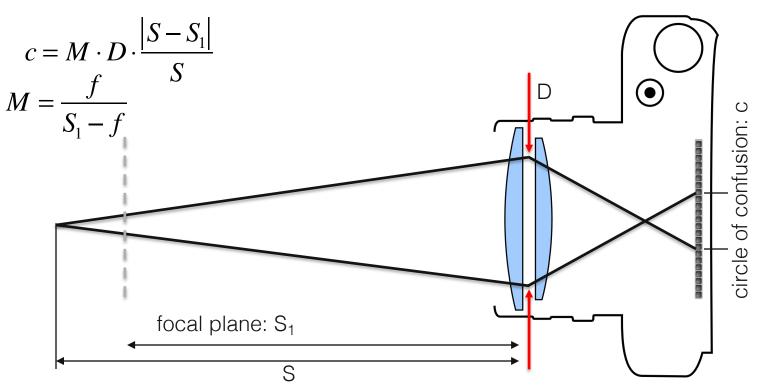
# Aperture out of focus blur circle of confusion focal plane

# Aperture out of focus blur circle of confusion focal plane

#### Depth of Field

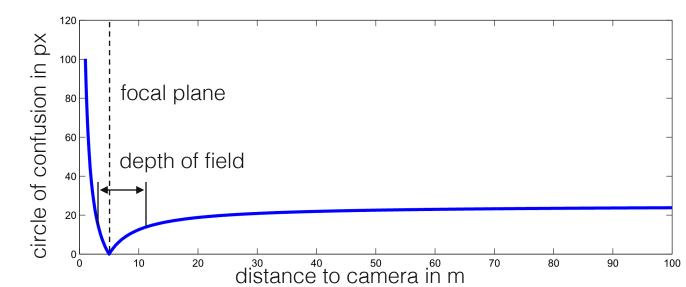


### Circle of Confusion



#### Circle of Confusion

$$c = M \cdot D \cdot \frac{\left| S - S_1 \right|}{S}$$
Canon 5D Mark III: f=50mm, f/2.8 (N=2.8), focused at 5m, pixel size=7.5um

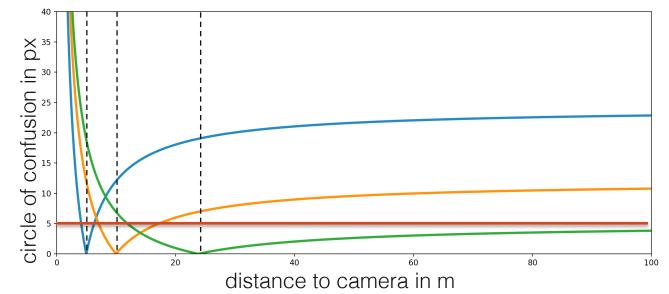


#### Hyperfocal Distance

$$H = \frac{f^2}{Nc}$$

<u>Canon 5D Mark III</u>: f=50mm, f/2.8 (N=2.8),

focused at 5m, pixel size=7.5um

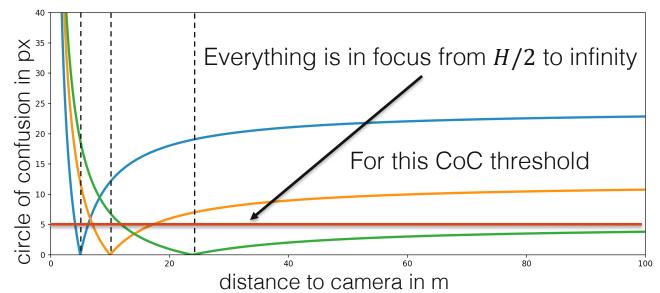


#### Hyperfocal Distance

$$I = \frac{f^2}{Nc}$$

<u>Canon 5D Mark III</u>: f=50mm, f/2.8 (N=2.8),

focused at 5m, pixel size=7.5um

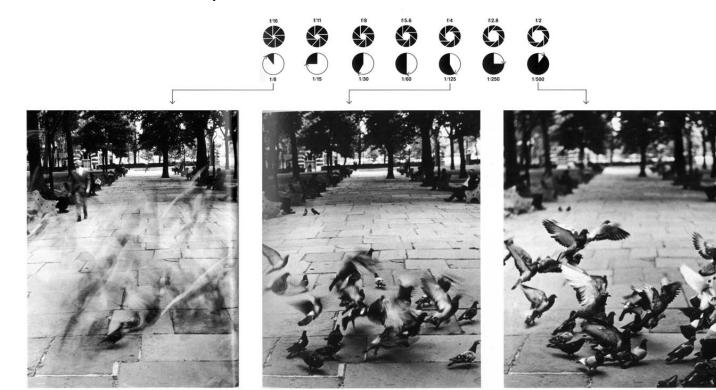


#### Depth of Field



aperture....f 1.8 shutter.....1/500 ISO......100 distance...~3ft aperture....f 4 shutter.....1/125 ISO......100 distance...~3ft aperture....f 8 shutter.....1/40 ISO......125 distance...~3ft http://photographywisdom

#### Depth of Field & Motion Blur



#### Bokeh

#### artistic use



#### coded aperture





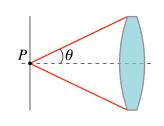


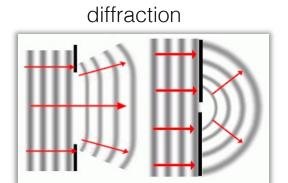


Levin et al., SIGGRAPH 2007

#### **Diffraction Limit**

• Ernst Abbe 1873:  $d = \frac{\lambda}{2n\sin\theta}$  spot radius (image space)





Airy pattern



## Diffraction Limit $\lambda \qquad \lambda \qquad \lambda$ $\lambda \qquad \lambda \qquad \lambda$

• Ernst Abbe 1873: 
$$d = \frac{\lambda}{2n\sin\theta} = \frac{\lambda}{2NA} \approx \lambda N$$
numerical aperture

- small f-number (large NA) = high resolution but shallow depth of field
  - inherent tradeoff between "3D" information and 2D resolution
  - space-bandwidth product (uncertainty principle)

microscope objectives today: NA 1.4-1.6  $\rightarrow$  d= $\lambda/2.8$ 

#### Fastest lens ever made?

Zeiss 50 mm f / 0.7 Planar lens





- Originally developed for NASA's Apollo missions.
- Stanley Kubrick somehow got to use the lens to shoot Barry Lyndon under only candlelight.

#### Fastest lens ever made?

Zeiss 50 mm f / 0.7 Planar lens



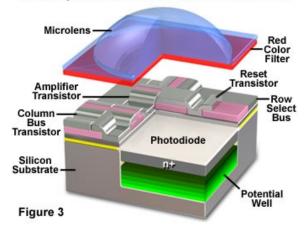


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- Stanley Kubrick somehow got to use the lens to shoot Barry Lyndon under only candlelight.

### Sensors

#### What's a Pixel?

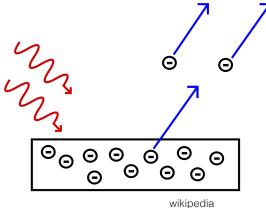
#### Anatomy of the Active Pixel Sensor Photodiode



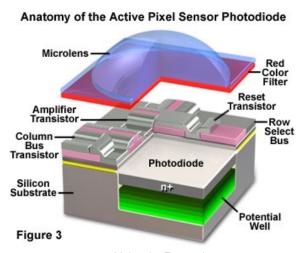
source: Molecular Expressions

photon to electron converter

→ photoelectric effect!

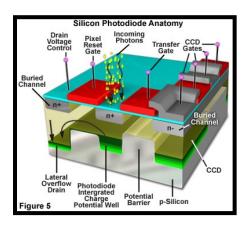


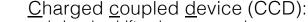
#### What's a Pixel?



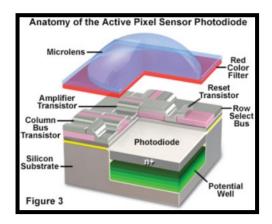
source: Molecular Expressions

- microlens: focus light on photodiode
- color filter: select color channel
- quantum efficiency: ~50%
- fill factor: fraction of surface area used for light gathering





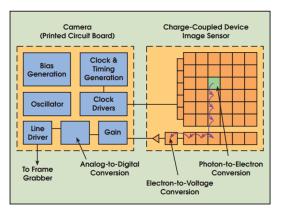
- row brigade shifts charges row-by-row
- amplifiers convert charges to voltages rowby-row



#### Complementary metal oxide semiconductor (CMOS):

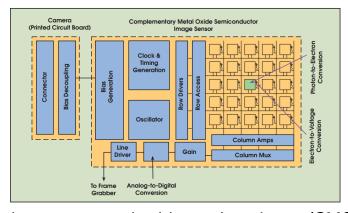
- per-pixel amplifiers convert charges to voltages
- multiplexer reads voltages row-by-row

Can you think of advantages and disadvantages of each type?





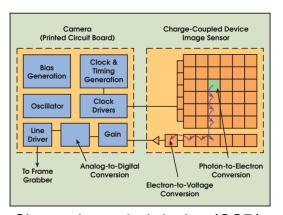
- row brigade shifts charges row-by-row •
- amplifiers convert charges to voltages row-by-row



#### Complementary metal oxide semiconductor (CMOS):

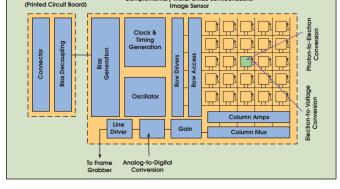
- per-pixel amplifiers convert charges to voltages
  - multiplexer reads voltages row-by-row

Can you think of advantages and disadvantages of each type?





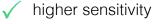
- row brigade shifts charges row-by-row •
- amplifiers convert charges to voltages row-by-row

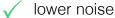


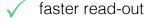
Complementary Metal Oxide Semiconductor

#### Complementary metal oxide semiconductor (CMOS):

- per-pixel amplifiers convert charges to voltages
- multiplexer reads voltages row-by-row









#### What's a Pixel?



### What's a Pixel?



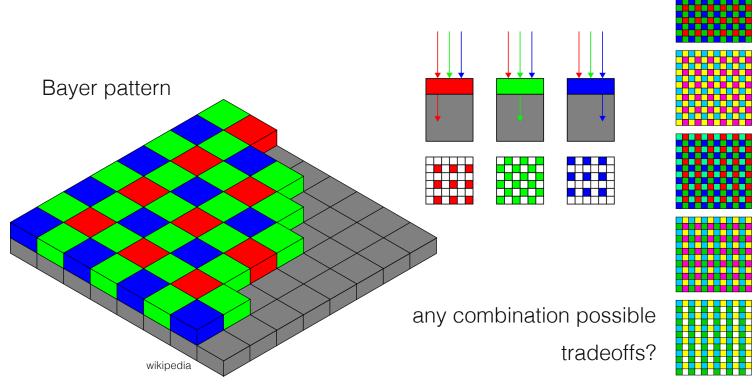
## What's a Pixel?



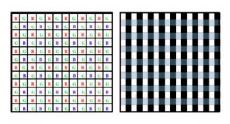
#### What's a Pixel?



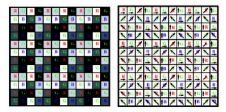
# Most Common: Color Filter Arrays

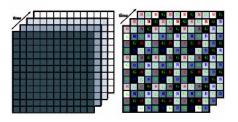


#### **Assorted Pixels**



- Narasimhan & Nayar @ Columbia
- multiplex anything: polarization, color, time, ND, ...



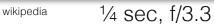




# Exposure (shutter speed)

exposure = time (e.g. 1/250, 1/60, 1, 15, bulb)







2 sec, f/6.3

# ISO ("film speed")

sensor sensitivity

analog gain applied before ADC!

















bobatkins.com

## Dynamic Range

ratio between largest and smallest possible value

bit depth also important! common bit depths: 12-14 bits RAW / 8

bits JPEG

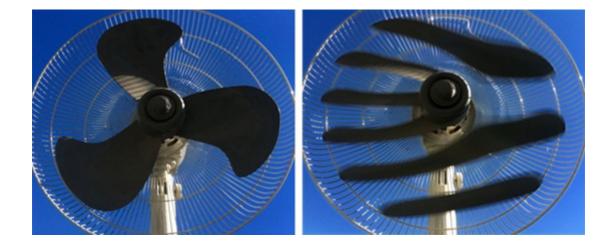
high dynamic range —







# Global Shutter vs. Rolling Shutter



All sensor pixels exposed at same time

Row-by-row readout of image

- shorter exposure times per pixel
- motion artifacts

What are these dark bands?

#### 60 Hz AC power results in 120 Hz flicker!



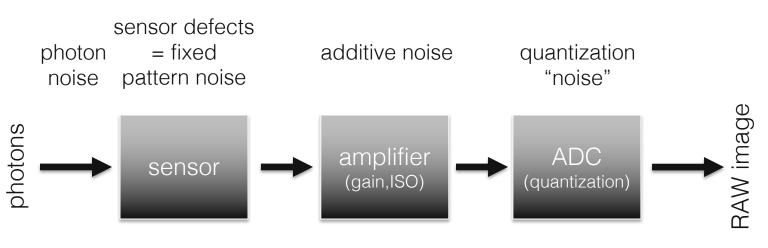
YouTube: user cameratest



26 frames over 10 ms [Sheinin et al. '17]

26 frames over 10 ms [Sheinin et al. '17]

## Photons to RAW Image



#### Sensor Noise

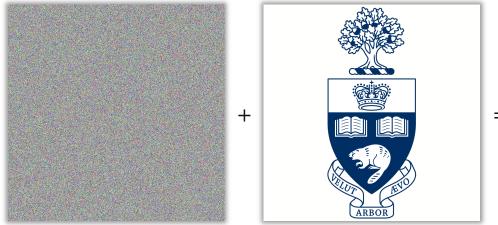
noise is (usually) bad!

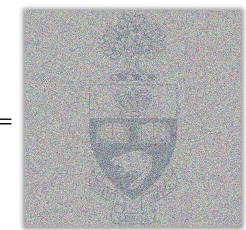
• many sources of noise: heat, electronics, amplifier gain, photon to electron conversion, pixel defects, read, ...

- different noise follows different statistical distributions, two crucial ones:
  - Gaussian
  - Poisson

#### Gaussian Noise

- thermal, read, amplifier
- additive, signal-independent!





#### Photon or Shot Noise

- signal dependent
- Poisson distribution:

$$f(k;\lambda) = \frac{\lambda^k e^{-\lambda}}{k!}$$
$$\sigma = \sqrt{\lambda}$$

N photons:  $\sigma = \sqrt{N}$ 2N photons:  $\sigma = \sqrt{2}\sqrt{N}$ 

nonlinear!



# Signal-to-Noise Ratio (SNR)

$$SNR = \frac{\text{mean pixel value}}{\text{standard deviation of pixel value}} = \frac{\mu}{\sigma} \leftarrow \frac{\text{signal}}{\text{noise}}$$

$$=\frac{PQ_{e}t}{\sqrt{PQ_{e}t+Dt+N_{r}^{2}}}$$

P = incident photon flux (photons/pixel/sec)

 $Q_{e}$  = quantum efficiency

t = eposure time (sec)

D = dark current (electroncs/pixel/sec), including hot pixels

 $N_r$  = read noise (rms electrons/pixel), including fixed pattern noise

#### Scientific Sensors

- e.g., Andor iXon Ultra 897: cooled to -100° C
- scientific CMOS & CCD
- reduce pretty much all noise, except for photon noise



## Digital Photography

- optics
- aperture
- · depth of field
- field of view
- exposure
- noise
- color filter arrays
- image processing pipeline



## Digital Photography – Additional Resources

What we left out: metering, autofocus, autoexposure, anti-aliasing filter,
 IR filter (and probably much more)

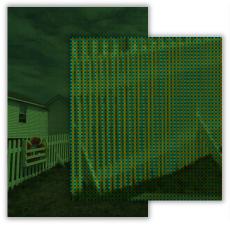
 Stanford CS 178 – Digital Photography: slides, applets, and other material online

CMU Computational Photography 15-862

looking for a camera? check dpreview.com

### Next: The Image Processing Pipeline

- RAW images
- demosaicking
- denoising
- deblurring
- white balancing
- gamma correction
- compression





## References and Further Reading

- London, Upton, Stone, "Photography", Pearson, 11th edition, 2013
- Stanford CS 178, "Digital Photography", Course Notes
- CMU Computational Photography course
- wikipedia