optical constructions 2

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optical constructions 1

laws of geometrical optics single interface planparallel plate prism and wedge

(thin) lens: magnifying glass microscope telescope lensmeter

matrix calculus

optical constructions 2

thick lens

antireflex coating camera obscura conversion lenses extension tubes

achromatised doublet

aberrations

anti-reflection coating



$$n_2 = \sqrt{n_1}$$

transmission over glass-air interface





spectrofotometric measurements







wavelength λ [nm]

	plan	-0,5 D	-1,0 D
T _{max} [%]	91,4	90,8	90,5
T _{pvg} [%]	90,7	90,3	89,8





nlan	

	plan	-0,5 D	-1,0 D
T _{max} [%]	99,0	98,8	98,8
T _{avg} [%]	97,8	97,5	87,5





	plan	-0,5 D	-1,0 D
T _{max} [%]	98,0	98,5	98
<i>T_{avg}</i> [%]	96,0	97,0	96,5



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	plan	-0,5 D	-1,0 D
T _{max} [%]	98,2	99,9	96,0
T _{avg} [%]	96,3	96,5	93,0

achromatised doublet





(here focus of a thin lens) depend on the wavelength

the light with different wavelength is focused into different points and with different magnification. The image gets blurred the image gets blurred with characteristic colouring of its contour:



this **chromatic aberration** is independent of refractive aberrations (it is present even for aspheric surfaces)

chromatic aberrations of a lens

Introduce the Abbe's number

$$v = \frac{n_d - 1}{n_f - n_f}$$

refractive indices:

n_D n_C

 n_F for blue line of hydrogen, $\lambda = 486, 1 \text{ nm}$

for yellow line of natrium, for red line of hydrogen, λ = 589,3 nmλ = 656,3 nm



high refractive index (thin lenses) and high Abbe's number (small chromatic aberration)

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R = R'

lens bulgyness

for thin submersed lens

$$\varphi' = \frac{n_L}{R} + \frac{n_L}{R'}$$
 ... lens maker equation
 $n\varphi' = (n_L - n) \left(\frac{1}{R} - \frac{1}{R'}\right) := (n_L - n)\rho$

we introduce the **bulgyness** *p*

(so that, in this case, the geometrical and material parameters get separated)

n - n

menisces

n

$$\rightarrow \rho = 0, f \rightarrow \infty$$

for $r1 \neq r2$ power of meniscus depends on its orientation

positive lenses





egative lenses
$$\rho = -\left(\frac{1}{|R|}\right)$$

$$= -\left(\frac{1}{|R|} + \frac{1}{|R'|}\right) < 0$$

 $\rho = \frac{1}{|R|} + \frac{1}{|R'|} > 0$

whether lens is positive or negative depends also on surrounding media

r

consider two adjacent lenses: $\phi' = \phi'_1 + \phi'_2$

using the bulginess,
$$\phi' = (n_1 - 1)\rho_1 + (n_2 - 1)\rho_2$$
 for some λ
 $\widetilde{\phi}' = (\widetilde{n}_1 - 1)\rho_1 + (\widetilde{n}_2 - 1)\rho_2$ for some $\widetilde{\lambda}$

if we want to suppress the chromatic aberration for these two wavelengths, it must hold $\phi' = \tilde{\phi}'$, ie.

$$\frac{\rho_1}{\rho_2} = -\frac{n_2' - n_2}{n_1' - n_1} = U$$

from Cauchy's formula this fraction of refractive indices is positive,

 $n_1 - n_1$ so that we need a **combination of a positive and a negative lens**

adjacent lenses share the inner radius of curvature, r_{z1} , the outer ones are free (r_{z1} , r_{z2})

the condition for the surfaces gets
$$\frac{1}{r_z} = \frac{\frac{1}{r_{z1}} + U \frac{1}{r_{z2}}}{1 + U}$$

by fulfilling it, we correct the chromatic aberration of position

usually, we choose $\lambda = \lambda_C$, $\lambda = \lambda_F$

we still have one degree of freedom, we use it to correct the chromatic aberration of magnification

achromatic doublet



a doublet gets achromatized when both chromatic error of position and magnigication get corrected



chromatic aberration of magnification is corrected when

$$\frac{f_{D1}}{f_{D2}} = -\frac{V_2}{V_1}$$

this brings a need to use one flint and one crown lens



alltogether, for prescribed focal power and chosen glass we have a system of three equations for three unknown radii, which gets solved uniquely:

$$\frac{1}{f_D} = (n_{D1} - 1)\rho_1 + (n_{D2} - 1)\rho_2 \qquad \frac{(n_{D2} - 1)\rho_2}{(n_{D1} - 1)\rho_1} = -\frac{\nu_2}{\nu_1} \qquad \qquad \frac{\rho_1}{\rho_2} = U$$

construction of optical systems

usually, the system are composed from groups of achromatized doublets:



eyepieces





Barlowov lens

pinhole camera

 $d = \sqrt{2f\lambda}$



pinhole camera





for optimal performance

$d = \sqrt{2f\lambda}$

pinhole camera

high aperture values:

- long expositions
 vignetting
 high field depth



www.paladix.cz

Pinhole camera picture: Telc main square, negative 6x9, pinhole 0.25mm, focal distance 20mm ... c=80

pinhole camera



Seidel aberrations



within paraxial space, the points are imaged into (perfect) points

in addition, points from a same plane image into same planes again:

we talk about focal planes.



in real systems, the perfect imaging is distorted by **aberrations** (longitudinal, transversal, etc.)

the images get blurred.



(we will not consider the chromatic aberrations here)

wave aberration



the **aberration coefficients** (of the expansion) can be minimised by the optical design we talk about producing the **circle of least confusion**

wave aberrations of axially symmetric systems

$$\vec{x}_{0} = (x_{0}, y_{0})$$
for description of optical systems we choose
position of (the point) source and
the coordinates of selected ray at the exit pupil
we can always put $y_{0} = 0$
thanks to axial symmetry, rotation must not matter,
the only important terms are the following: $\vec{x}_{0} \cdot \vec{x}_{0}, \vec{x} \cdot \vec{x}_{0}, \vec{x} \cdot \vec{x}$
the wavefront difference gets $H(x_{0}, y_{0}, x, y) \rightarrow H(x_{0}^{2}, x_{0}x, x^{2} + y^{2})$ ($y_{0} = 0$)
we use polar coordinates $x = \rho \cos \vartheta \ y = \rho \sin \vartheta$ (ρ is the aperture)
 $H(x_{0}^{2}, x_{0}\rho \cos \vartheta, \rho^{2}) = \sum_{k,l,m} W_{klm}x_{0}^{k}\rho^{l} \cos^{m} \vartheta =$
 $= W_{000} + W_{200}x_{0}^{2} + W_{111}x_{0}\rho \cos \theta + W_{002}\rho^{2} +$
 $+ W_{400}x_{0}^{4} + W_{040}\rho^{4} + W_{131}x_{0}\rho^{3} \cos \theta + W_{222}x_{0}^{2}\rho^{2} \cos^{2} \theta +$
 $+ W_{220}x_{0}^{2}\rho^{2} + W_{311}x_{0}^{3}\rho \cos \theta + \dots$

note the special case of axial source, $x_0 = 0$: only W_{o40} remains from green lines

Seidel aberrations 1856 (axially symmetric systems)

lowestorder terms can be organized into five (Seidel) aberrations:

$$H = \frac{1}{8}S_{I}\rho^{4} + \frac{1}{2}S_{II}x_{0}\rho^{3}\cos\vartheta + \frac{1}{2}S_{III}x_{0}^{2}\rho^{2}\cos^{2}\vartheta + \frac{1}{4}(S_{III} + S_{IV})x_{0}^{2}\rho^{2} + \frac{1}{2}S_{V}x_{0}^{3}\rho\cos\vartheta$$

spherical aberration S_{II} , coma S_{III} , astigmatism S_{III} , curvature S_{IV} , distortion S_{V}

the piston (W $_{\rm 200}$), tilt (W $_{\rm 111}$) and defocus (W $_{\rm 002}$) do not belong here (they can be removed by moving the system)

higher order aberrations bring more subtle behavior (secondary coma etc..)

good thing about Seidel coefficients:

the overall aberration can be found as a sum of aberrations from individual surfaces

 $S_{I} = S_{I}^{1} + S_{I}^{2} + S_{I}^{3} + S_{I}^{4} + \cdots$



spherical aberration

$$H(y) = \frac{1}{8} S_I \rho^4$$

the blur is symmetrically shaped (circles)

for aperture ρ , the blur extents to $\frac{1}{2} f S_I \rho^3$

(f is the focal distance)

- it is very effective to fight spherical aberration using small apertures ho
 ightarrow 0
- spherical aberration is present even for axial sources
- systems with spherical aberration repaired are called stigmatic at the axis





 $H(y) = \frac{1}{2} S_{II} x_0 \rho^3 \cos \vartheta$

coma

coma consists from circles

of radii

and offset

$$\frac{1}{2}fS_{II}x_{0}\rho^{2}$$
$$fS_{II}x_{0}\rho^{2}$$

- coma is observed only for nonaxial sources, small apertures help
- the circles fill up an angle of 60°, length : width of the whole blurr is 3:2
- systems without both spherical aberration and coma are called aplanatic





astigmatism

$$H(y) = \frac{1}{4} S_{III} x_0^2 \left(3x^2 + y^2 \right)$$

the aberration has a form of ellipses with axes ratio 1:3

- the sagital a meridional rays are each focused perfectly, but in different planes
- big aperture is not that harmful, but astigmatism rises steeply with nonaxial rays (ie. with field of view)

