



Earth isn't the only ocean world in our solar system. Oceans could exist in diverse forms on moons and dwarf planets, offering clues in the quest to discover life beyond our home planet.



SIZE COMPARISON

1 AU

DISTANCE FROM SUN

ACTIVE

Dynamic ocean, known to support life

OCEAN WORLD STATUS

TERRESTRIAL PLANET

EARTH

Our home planet, Earth, is the only body known to have life. Called the "ocean planet," Earth's surface-land-to-water ratio is 29% land to 71% water.



SIZE COMPARISON

5.2 AU

DISTANCE FROM SUN

#### ACTIVE?

Possibly a dynamic ocean, could support life

MOON OF JUPITER

## EUROPA

Scientists strongly suspect that a subsurface salty ocean lies beneath Europa's icy crust. Tidal heating from its parent planet, Jupiter, maintains this ocean's liquid state and could also create partially melted pockets, or lakes, throughout the moon's outer shell.





MOON OF JUPITER

# GANYMEDE

Ganymede is the largest moon in our solar system, and the only moon with its own magnetic field. Recent studies indicate a large, underground saltwater ocean is present at the Jovian moon. Ganymede could in fact have several layers of ice and water sandwiched between its crust and core.





SIZE COMPARISON

5.2 AU

DISTANCE FROM SUN

### LOCKED

Trapped ocean,
unlikely to support life



SIZE COMPARISON

5.2 AU

DISTANCE FROM SUN

### LOCKED

Trapped ocean, unlikely to support life OCEAN WORLD STATUS MOON OF JUPITER

## CALLISTO

Callisto's cratered surface lies at the top of an ice layer, which is estimated to be about 60 miles (100 km) thick. An ocean, which is thought to be at least 6 miles (10 km) deep, could be directly beneath the ice.





MOON OF SATURN

# ENCELADUS

Scientists predict that a regional reservoir about 6 miles (10 km) deep lies under a shell of ice 20 to 25 miles (30 to 40 km) thick at Enceladus' south pole. This underground ocean is thought to feed the moon's impressive jets, which spray from deep fissures (called "figer stripes") in the moon's surface.



SIZE COMPARISON

9.5 AU

DISTANCE FROM SUN

### ACTIVE

Dynamic ocean, could support life

OCEAN WORLD STATUS



SIZE COMPARISON

9.5 AU

DISTANCE FROM SUN

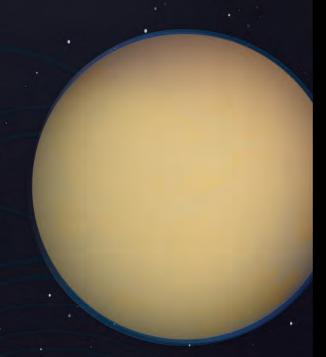
## LOCKED?

May have a trapped ocean, unlikely to support life if ocean is trapped

OCEAN WORLD STATUS

# MOON OF SATURN

Titan is believed to have a salty subsurface ocean -- as salty as the Dead Sea on Earth -- beginning about 30 miles (50 km) below its ice shell. It is also possible that Titan's ocean is thin and sandwiched between layers of ice, or is thick and extends all the way down to the moon's rocky interior.





SIZE COMPARISON

30.1 AU

DISTANCE FROM SUN

## POSSIBLE

Evidence of an ocean, biological potential unknown OCEAN WORLD STATUS

# MOON OF NEPTUNE

# TRITON

Active geysers on Triton spew nitrogen gas, making this moon one of the known active worlds in the outer solar system. Volcanic features and fractures mark its cold, icy surface, likely results of past tidal heating. A subsurface ocean at Triton is considered possible, but is unconfirmed.





## DWARF PLANET

# PLUTO

A world of many unknowns, Pluto could have rings and perhaps a subsurface ocean.

Data from NASA's New Horizons mission will provide new insights about this unexplored world.



SIZE COMPARISON

39.5 AU

DISTANCE FROM SUN

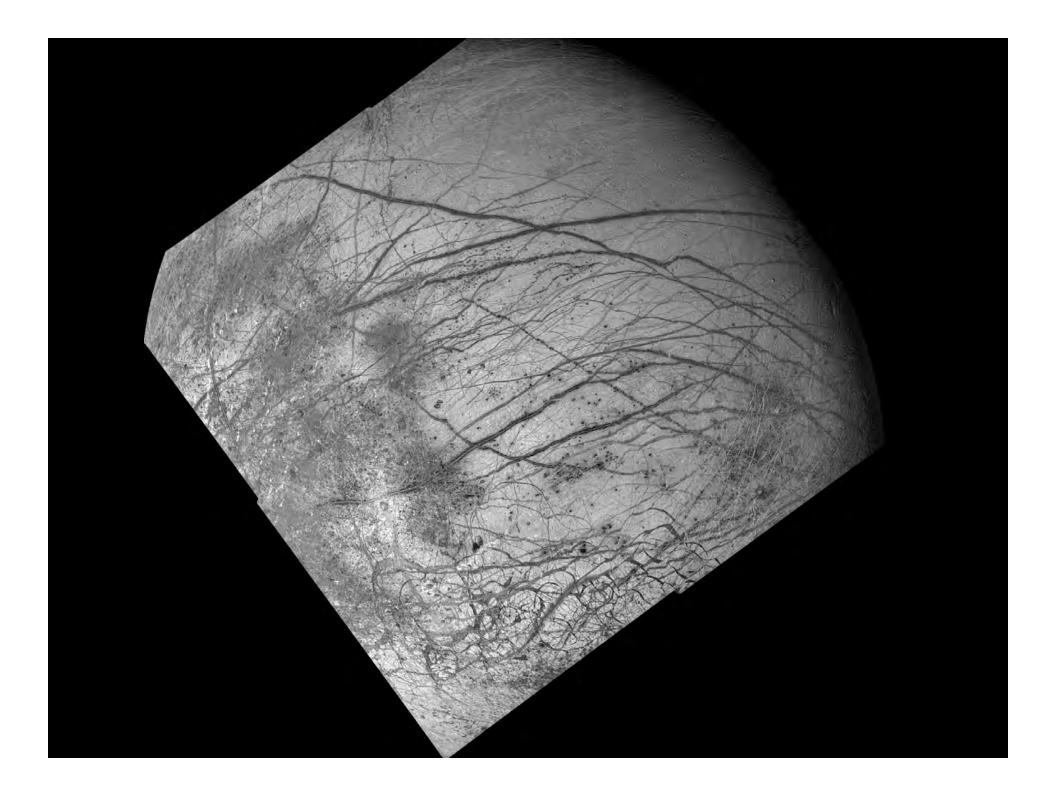
## POSSIBLE

Evidence of an ocean, biological potential unknown

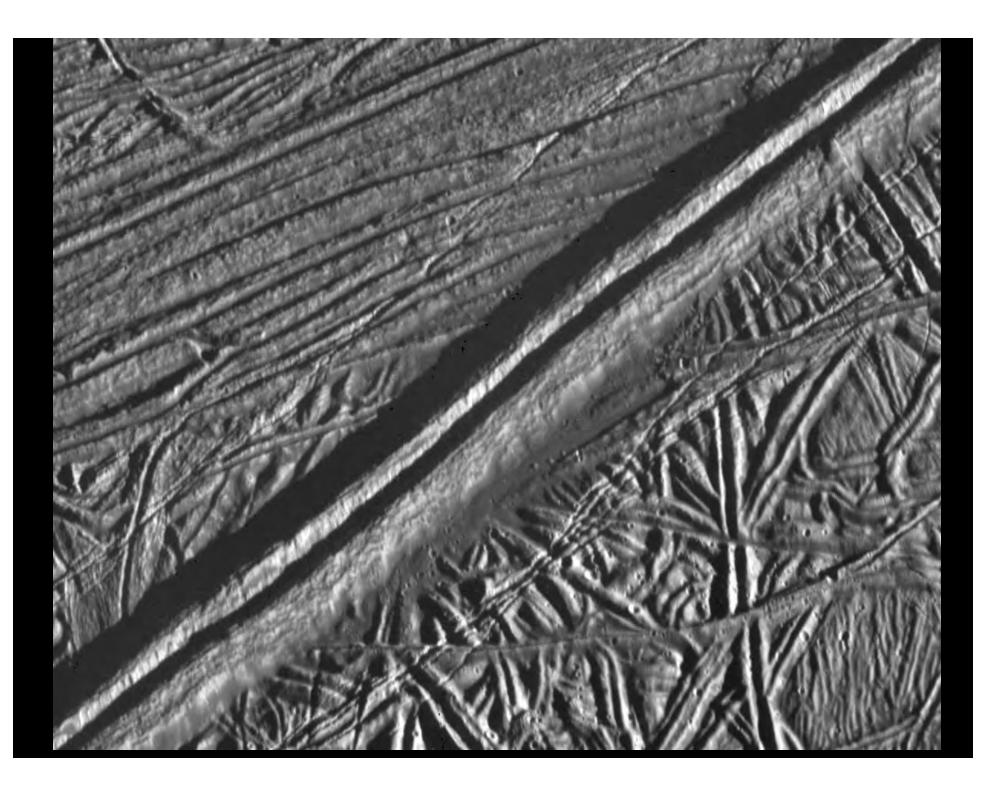
OCEAN WORLD STATUS

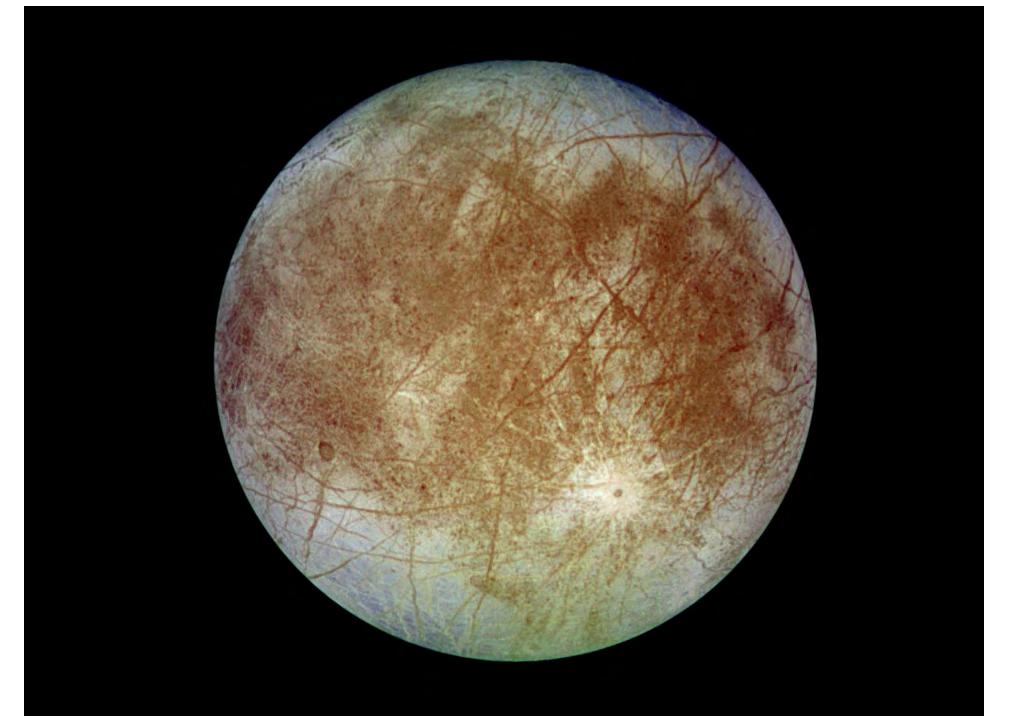


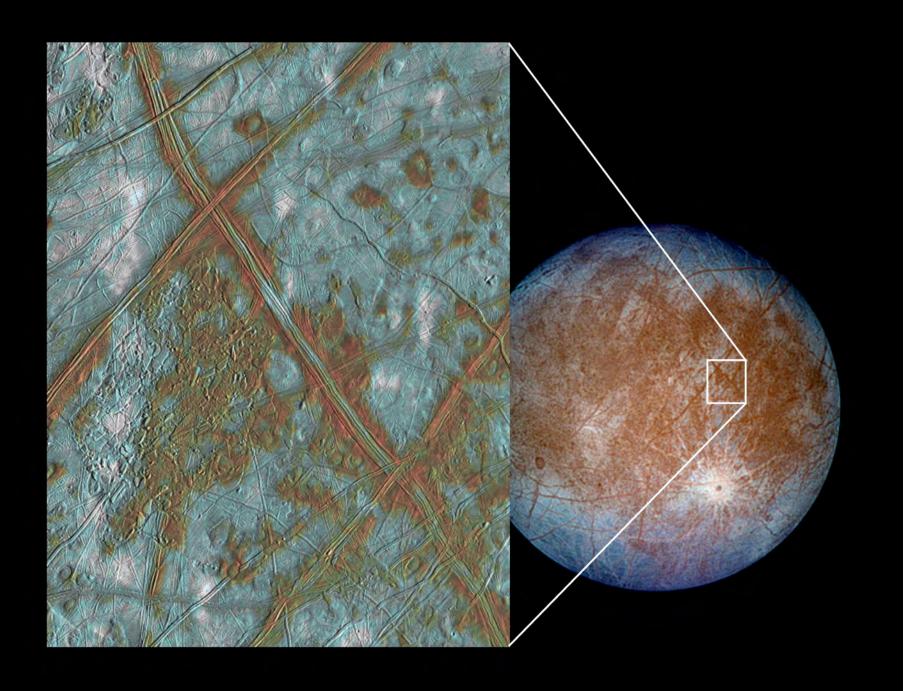


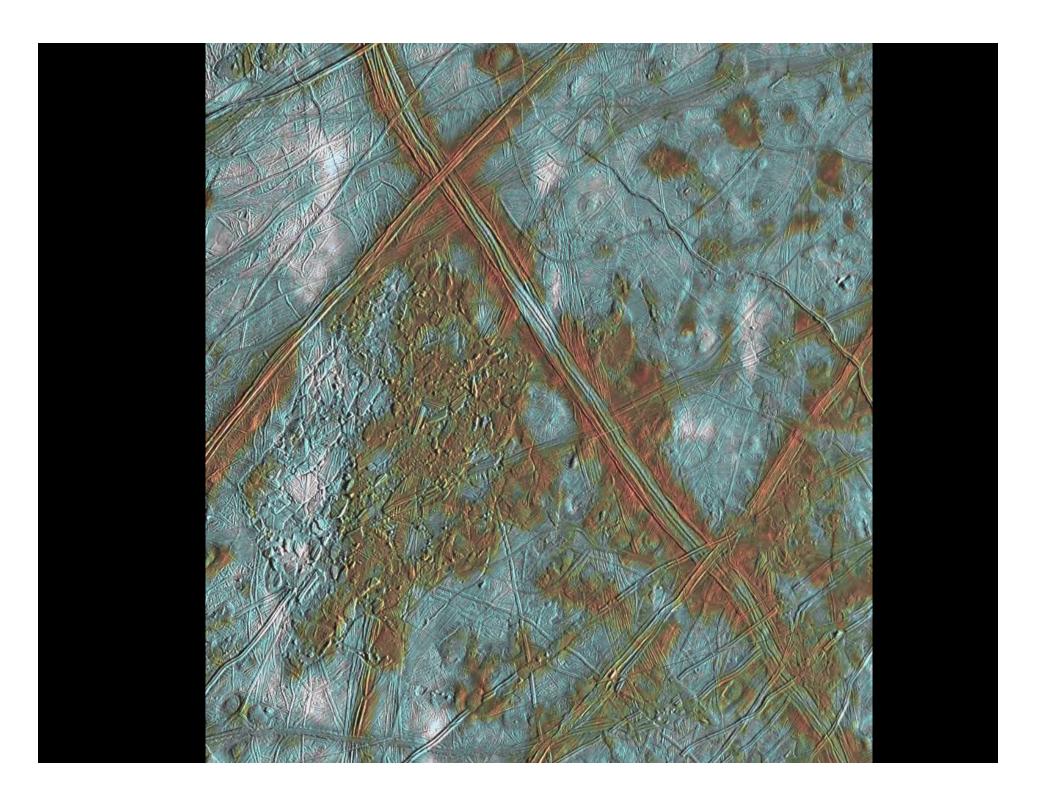


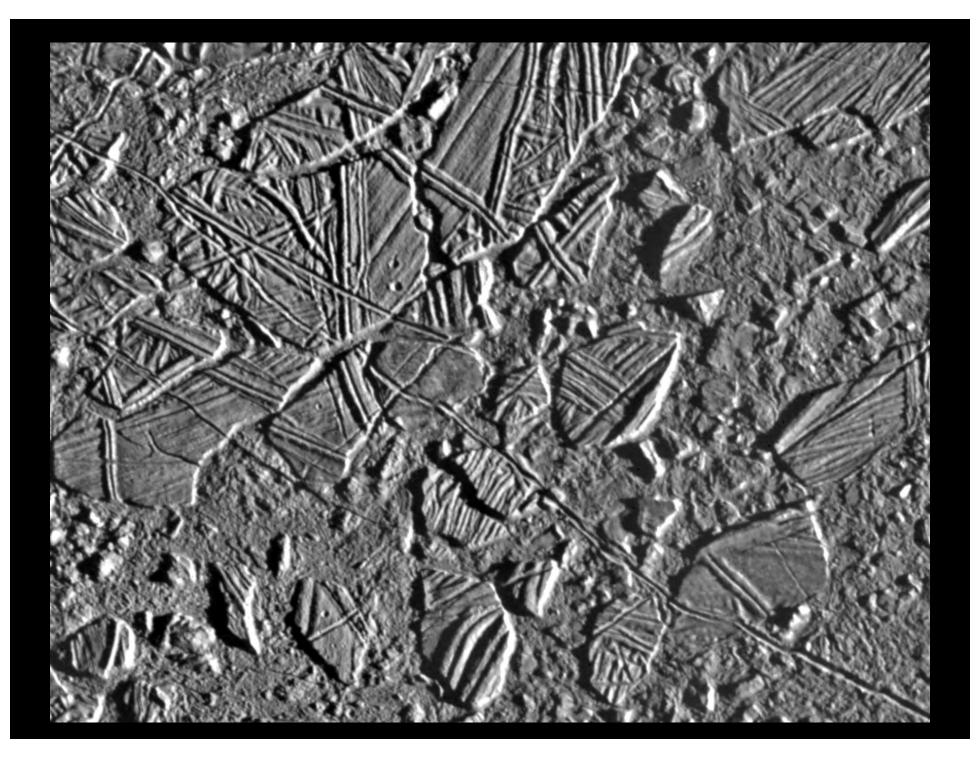


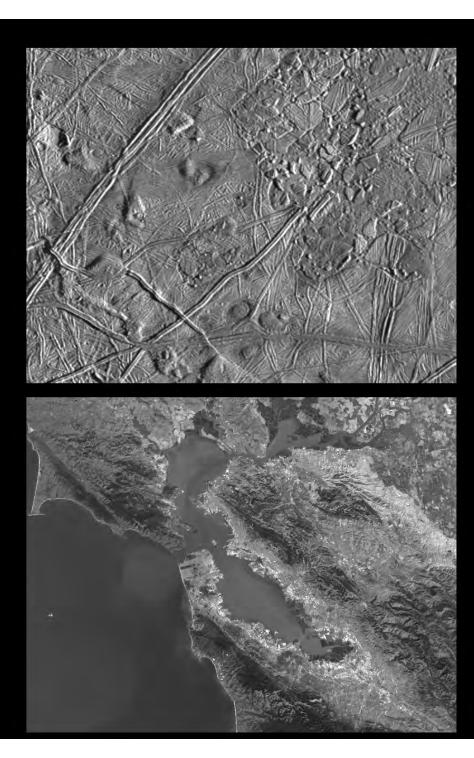




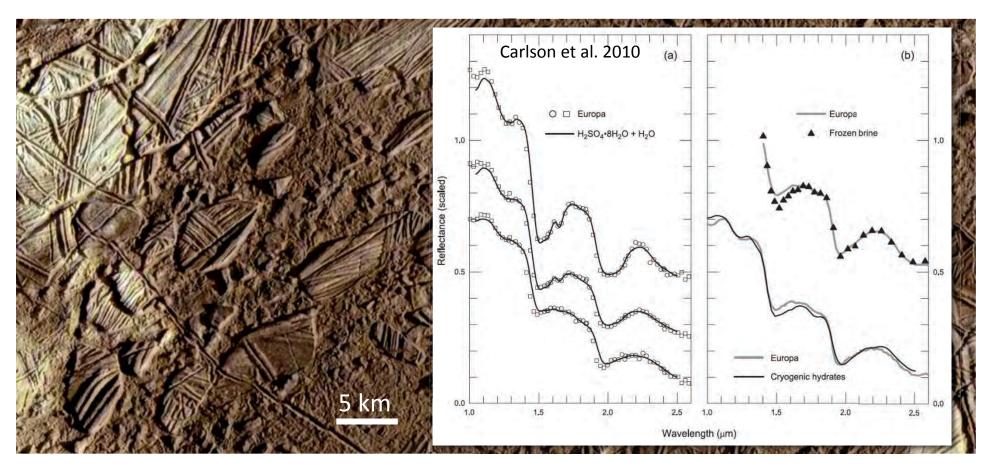




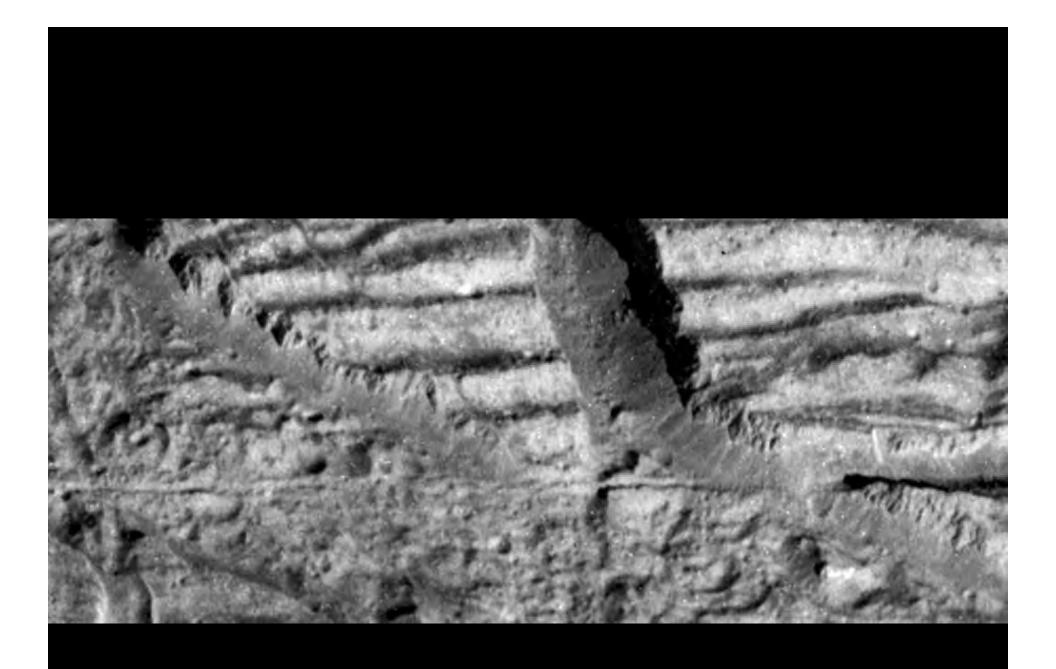


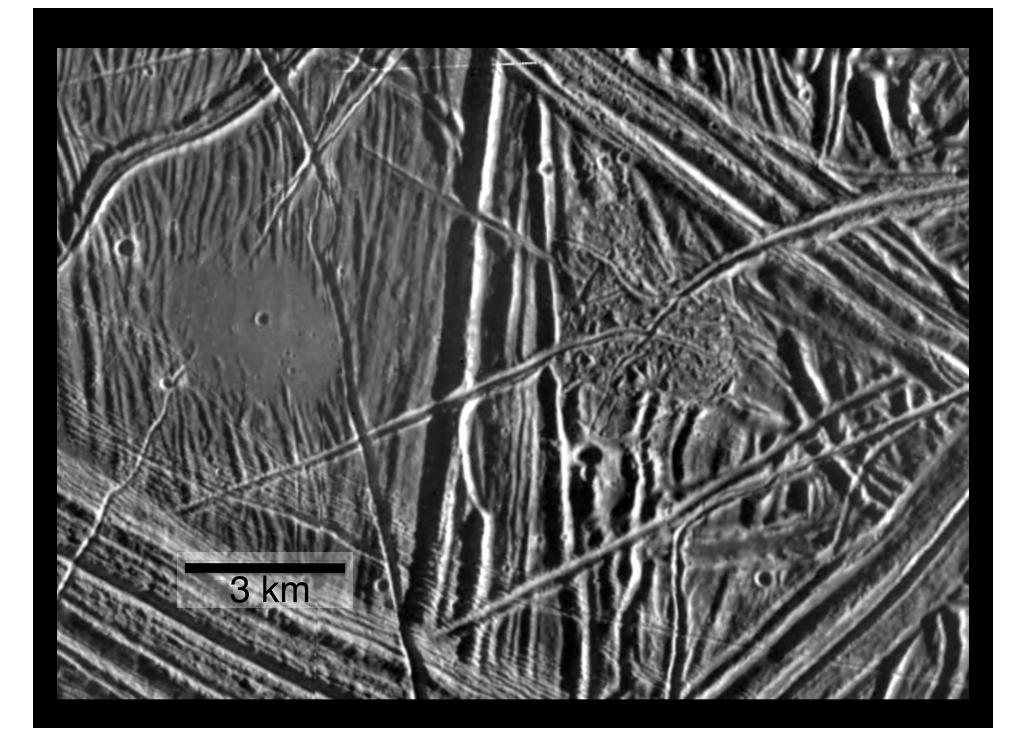


# Composition of Europa's non-ice

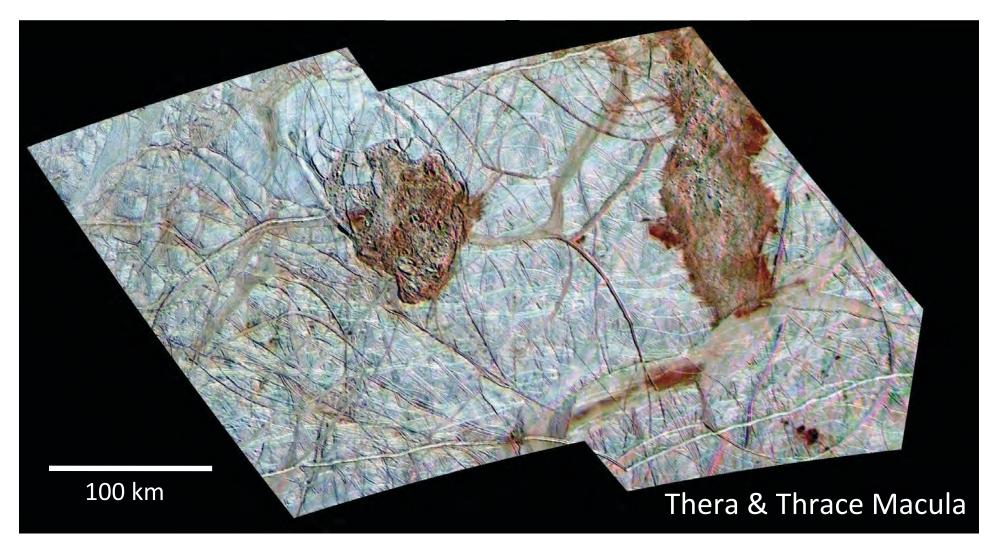


- Leading candidates are hydrated sulfates
- ♦ Sulfuric acid hydrate almost surely present, Mg-sulfate hydrate?
- ♦ Na, K seen in sputtered atmosphere

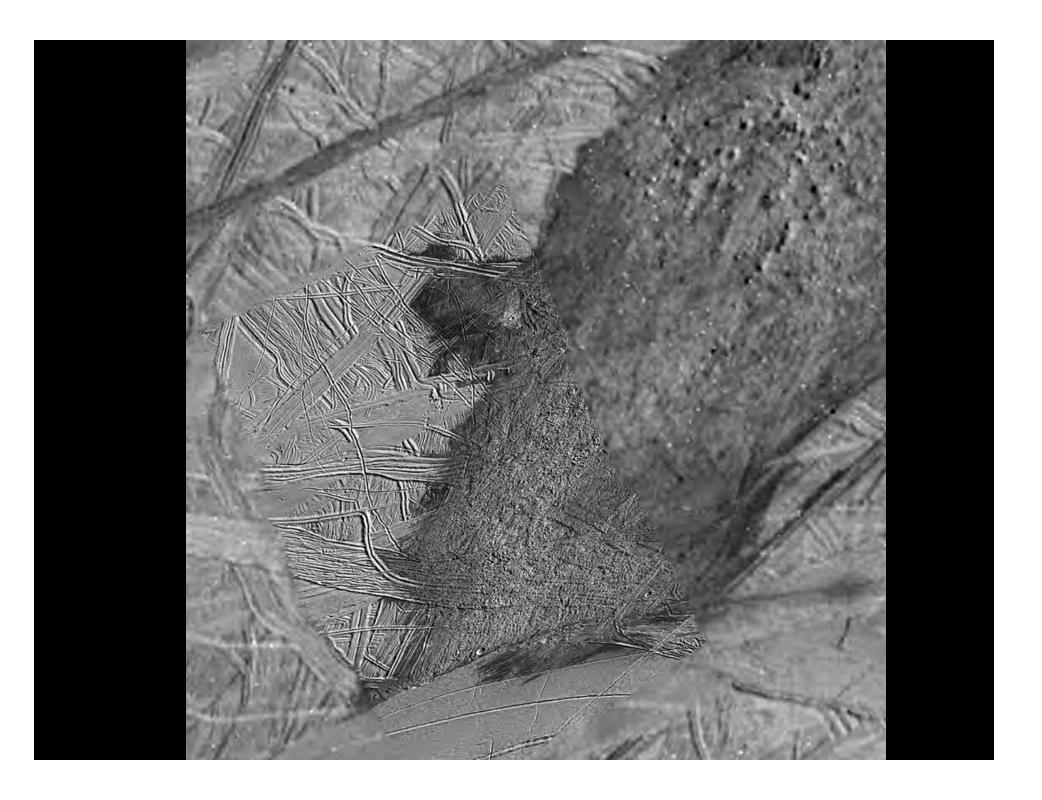


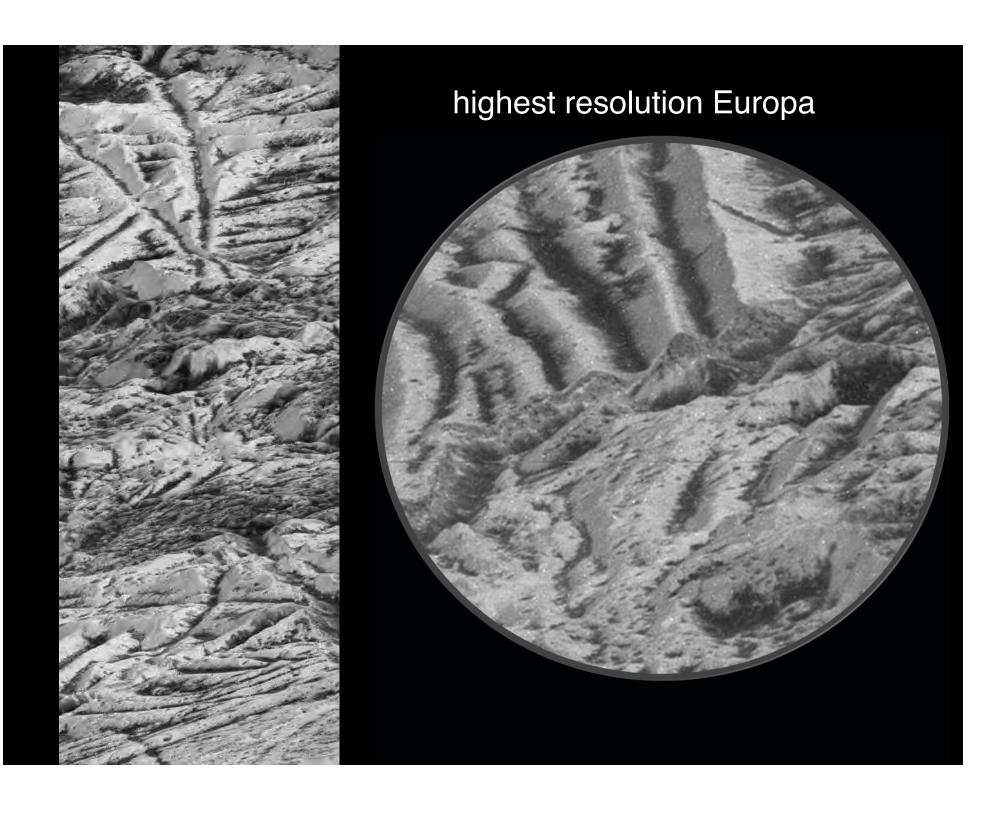


# Is **Europa** geologically active?

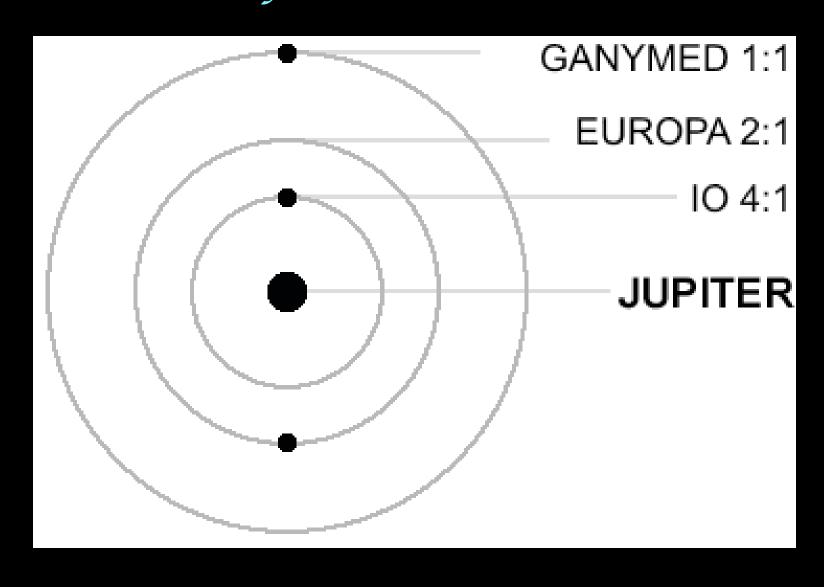


- ♦ Average surface age from impact crater density 65 ± 25 Ma
- ♦ Thera Macula has been proposed as presently active (Schmidt et al., 2011)

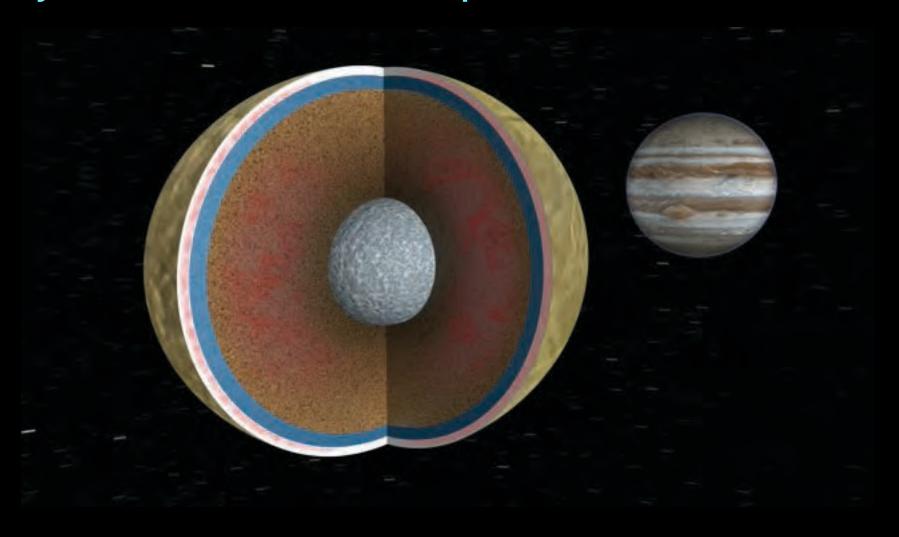




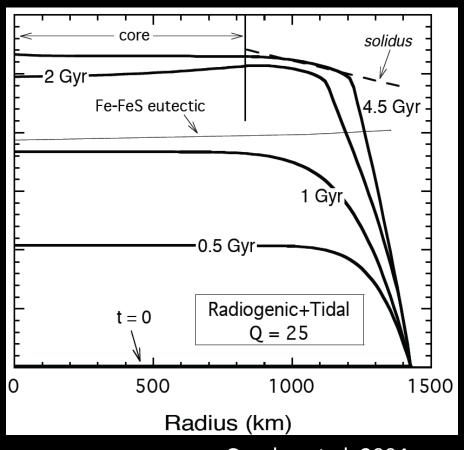
# the Laplace Resonance

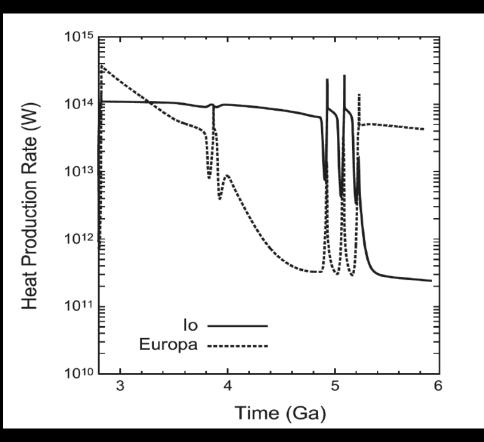


# jovian tides on Europa



# ENERGY: RADIOGENIC & TIDAL





Greeley et al. 2004

Moore & Hussmann 2009



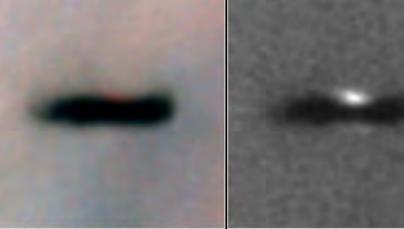
distance = 500 pc panels 3000 AU across



Protoplanetary Disks Orion Nebula

HST · WFPC2

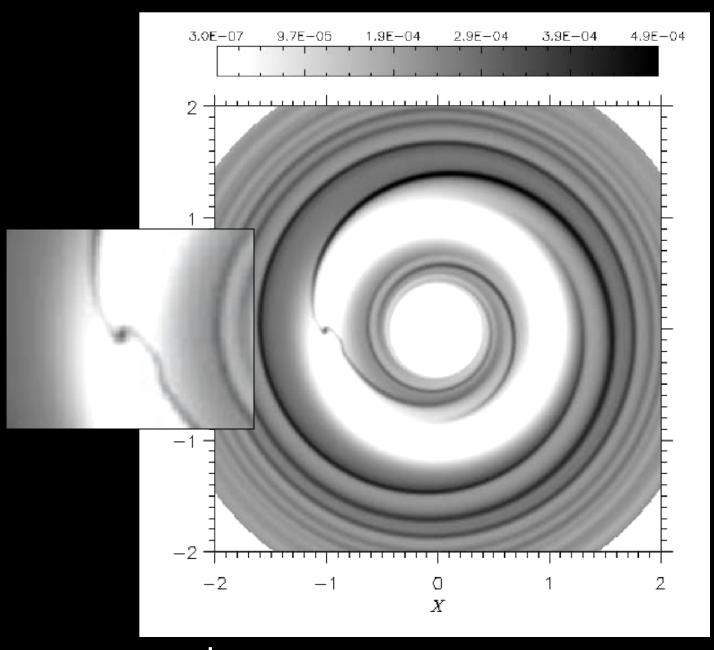
PRC95-45b · ST Scl OPO · November 20, 1995 M. J. McCaughrean (MPIA), C. R. O'Dell (Rice University), NASA



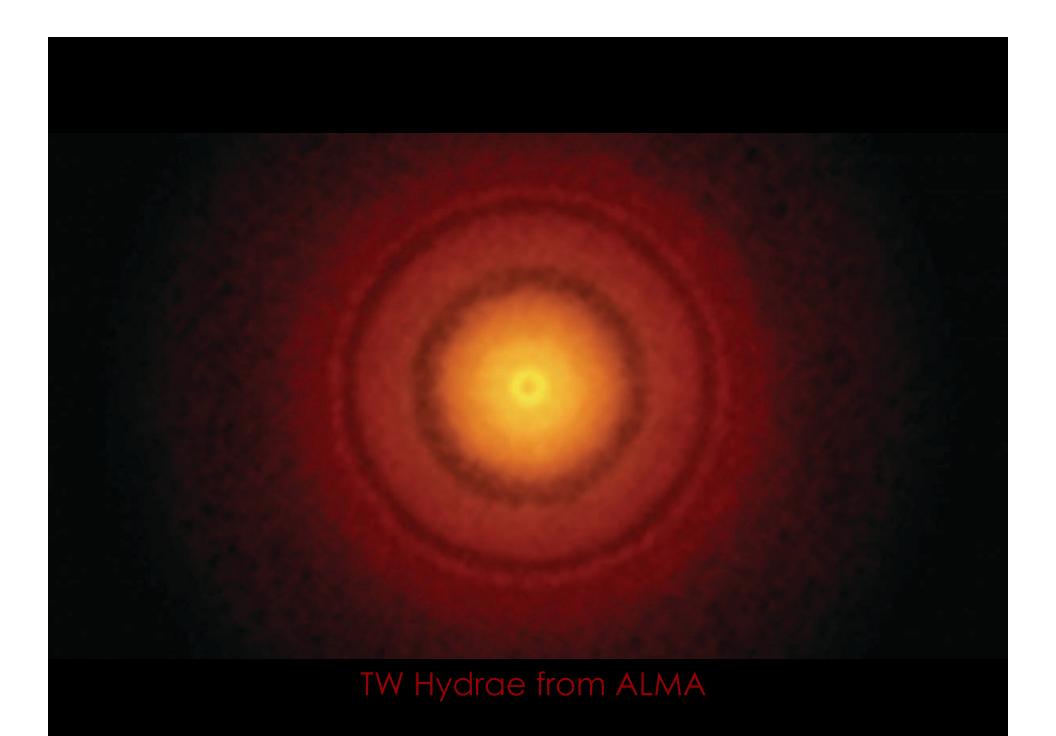
Edge-On Protoplanetary Disk Orion Nebula

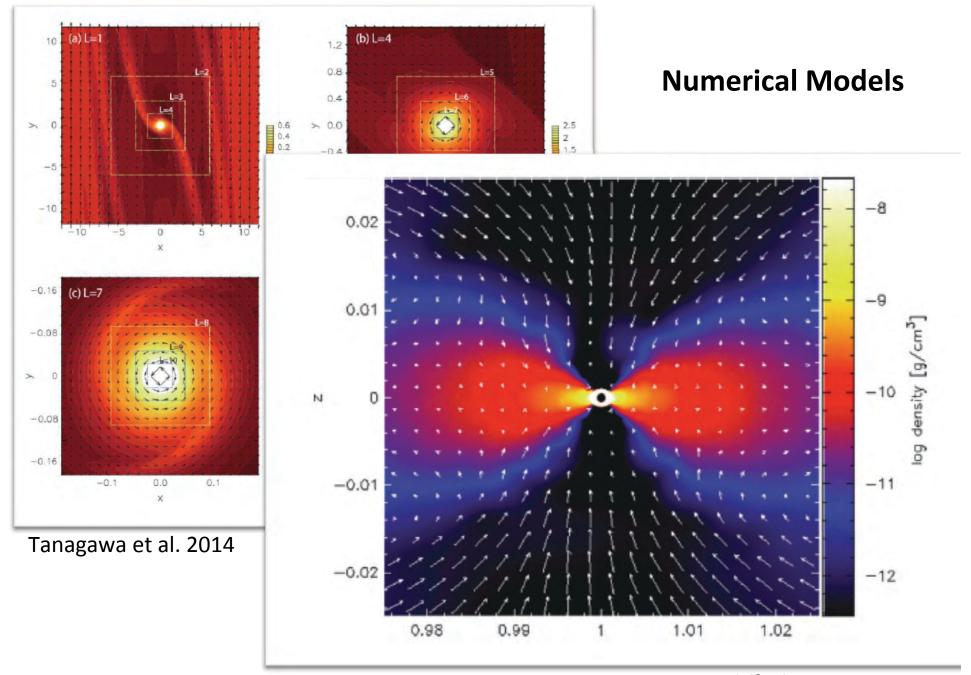
HST · WFPC2

PRC95-45c · ST ScI OPO · November 20, 1995 M. J. McCaughrean (MPIA), C. R. O'Dell (Rice University), NASA



D' Angelo et al. (2003) Astrophys. J. 599, 548-576.





Ayliffe & Bate 2009

# RECENT CARCONACEOUS METEORITE FALLS: CANDIDATE PRIMITIVE EUROPA ROCK





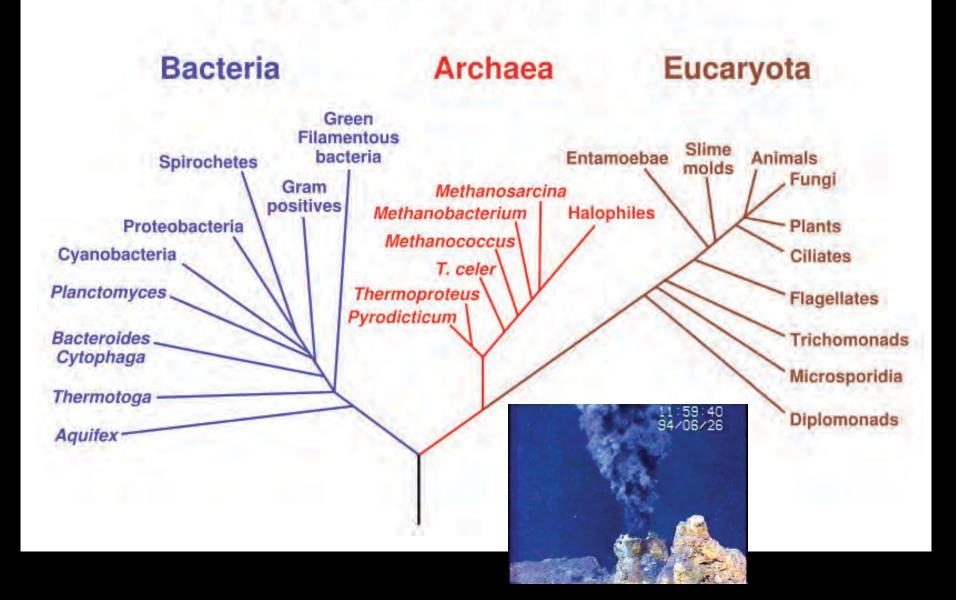


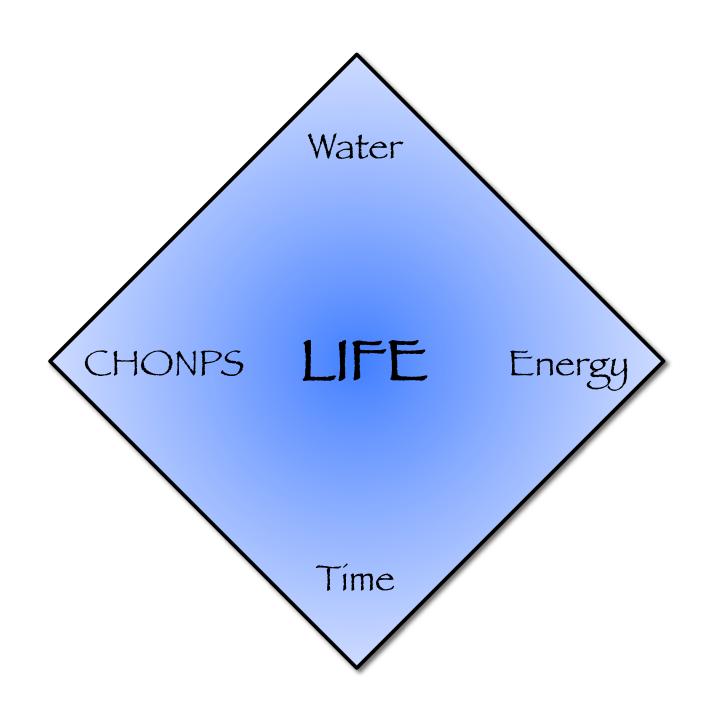
Hydrated & anhydrous silicates, sulfides, sulfates, oxides, carbonates, phosphides, phosphates, chondrules, CAIs, CHONPS—unequilibrated but ~ solar!

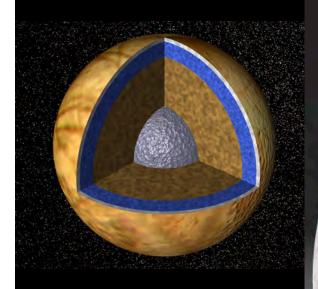


Tagish Lake chondrite fall,

# Phylogenetic Tree of Life







# Scientists say a moon of Jupiter could be the best chance for finding life outside Earth.

### Cracks in the ice

Strong tides knead the ice shell. A surface probe could detect life where water has spilled out.

### On thin ice?

The ice shell thickness is hotly debated. A thick shell could seal the ocean off from light A thin shell could allow life to reach the surface with daily tides.

#### Salt-water ocean

A strange magnetic signal was evidence for Europa's warm, salt-water ocean, bigger than all of the Earth's.

### **Hydrothermal vents**

Life on Earth may have begun near hot, ocean-floor vents. Some microbes can survive, without light and oxygen, on only the heat and chemicals near the vents.

CROSS-SECTION OF EUROPA'S SURFACE

## SOURCE: NASA | Post-Dispatch

### **Barren surface**

Charged particles from Jupiter would kill any surface life.
But these particles also make oxygen that could enter the ocean.

miles

Note: Drawing

100 miles

## SIZE



Earth and moon Europa

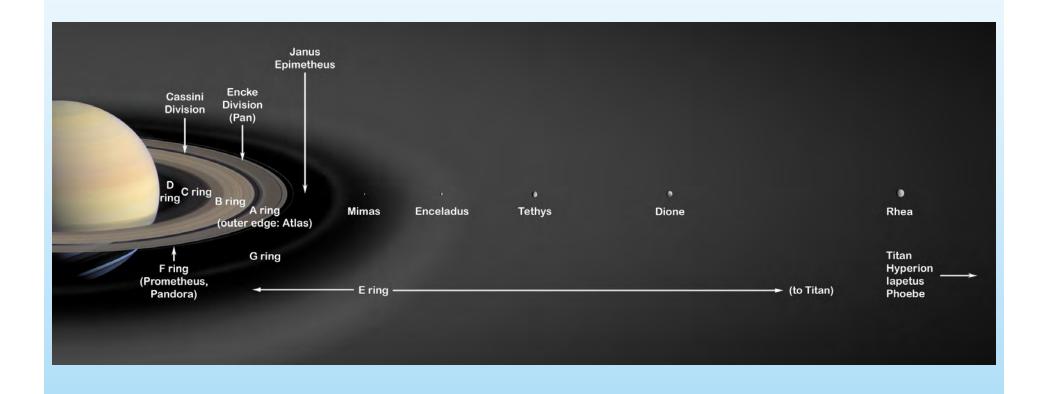


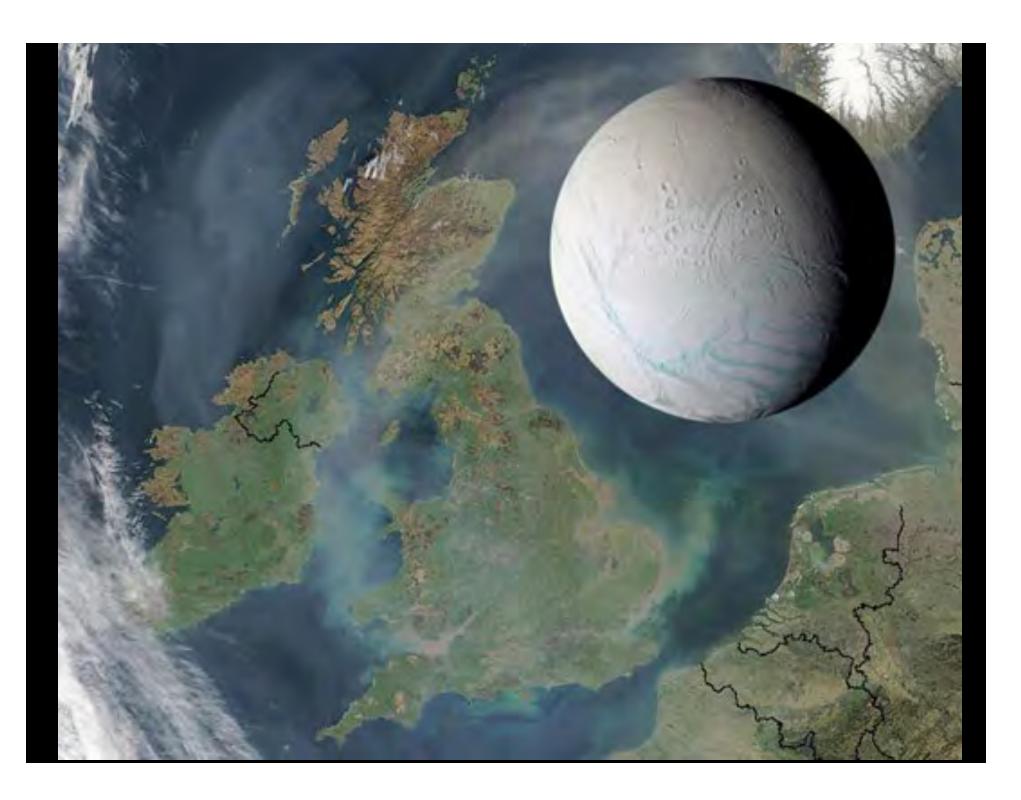
### The surface

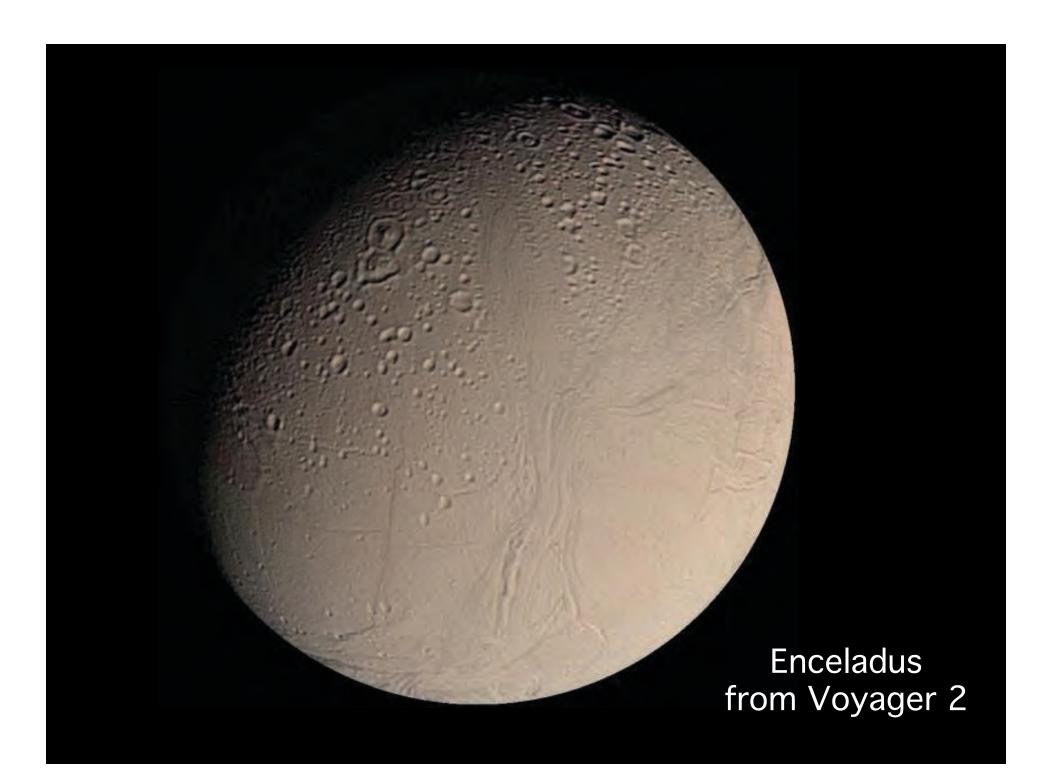
is fractured and mostly free of impact craters evidence for a young, stressed land that, like a hockey rink, may be resurfaced.

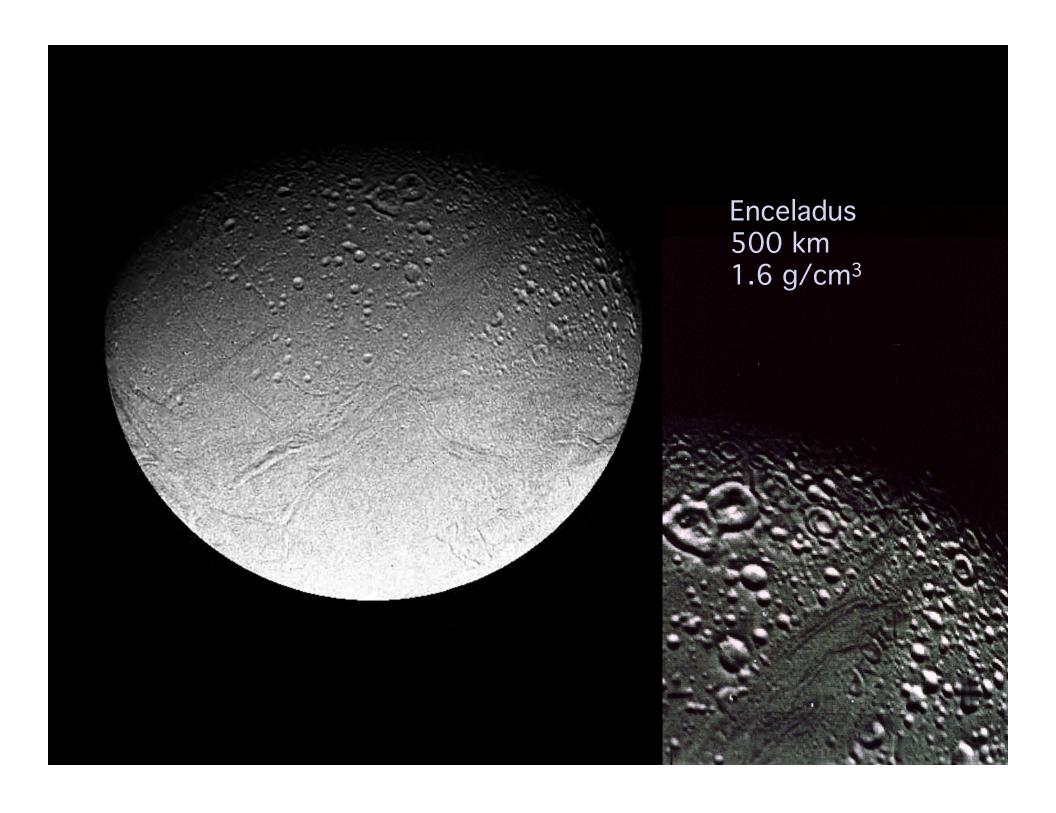
Eric Hand (2007) St. Louis Post-Dispatch

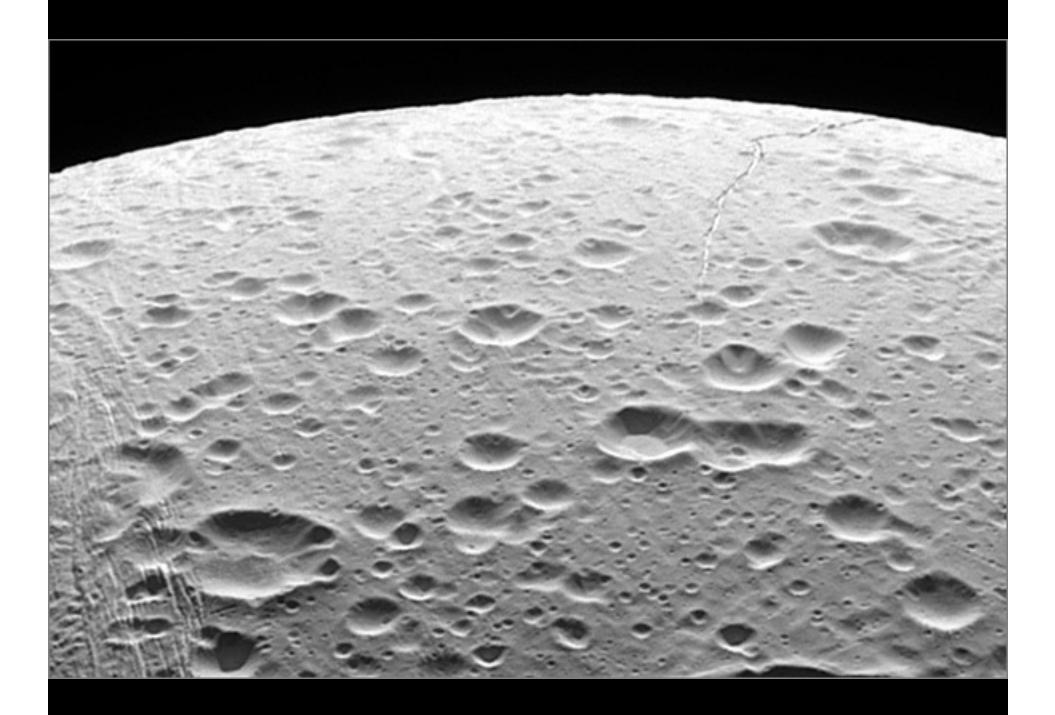
# On to Enceladus

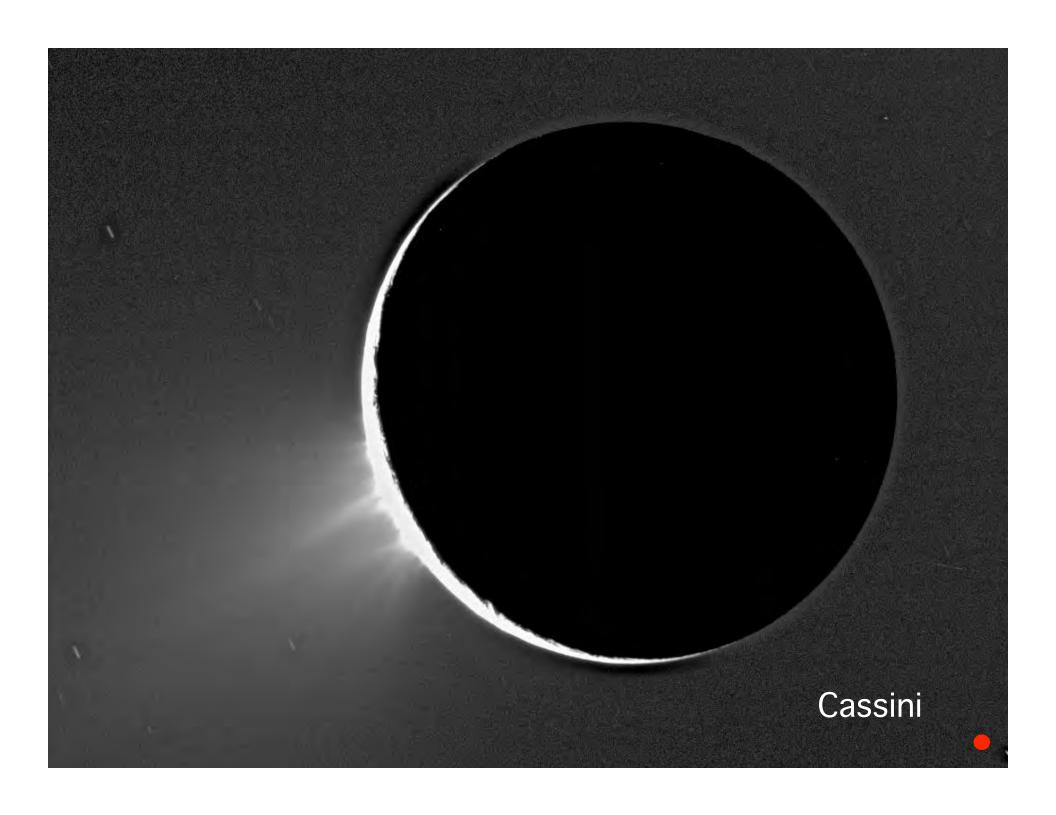




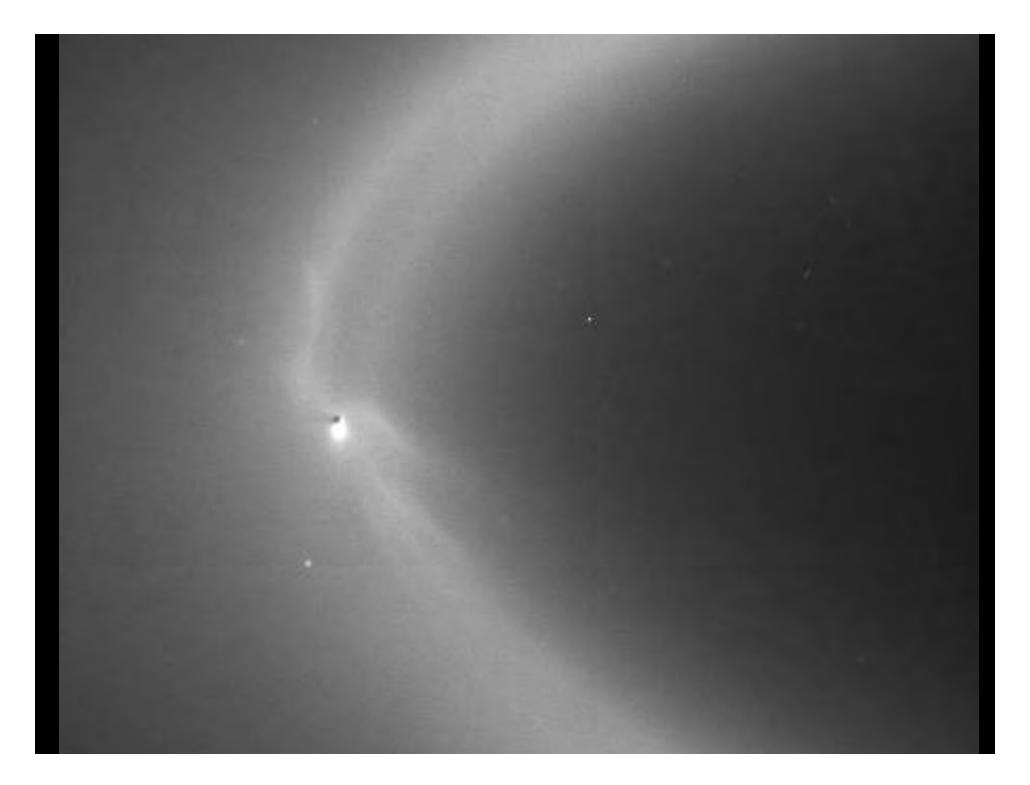


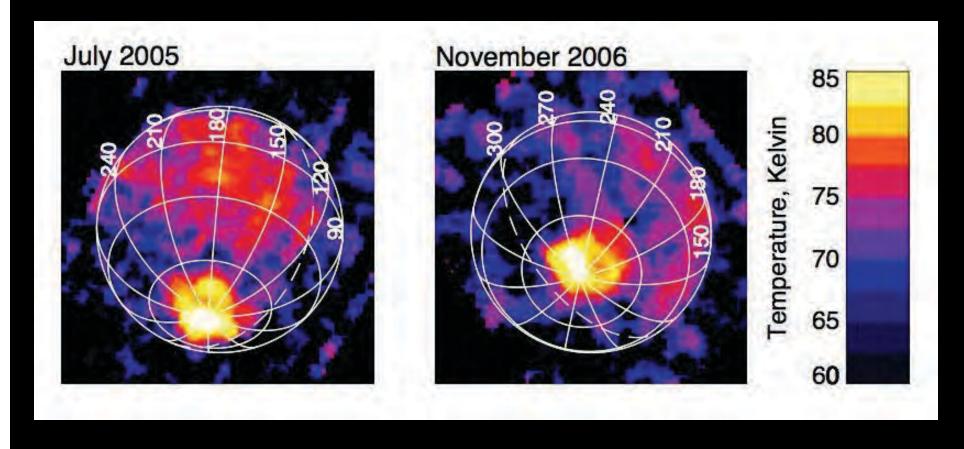




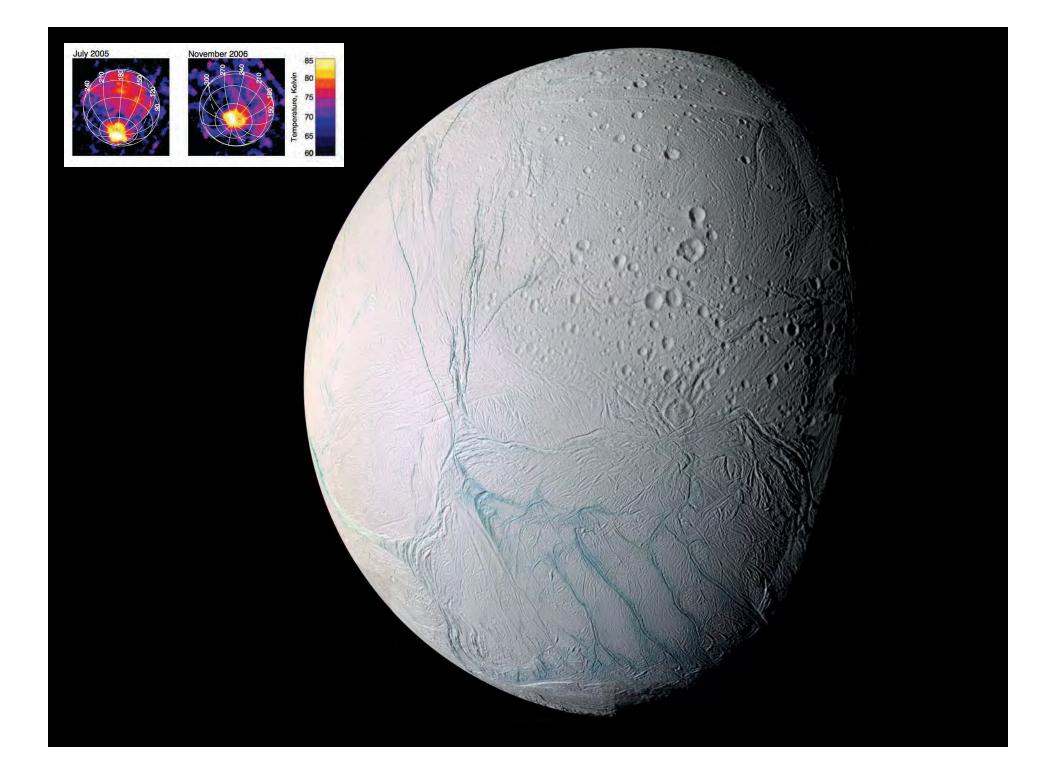


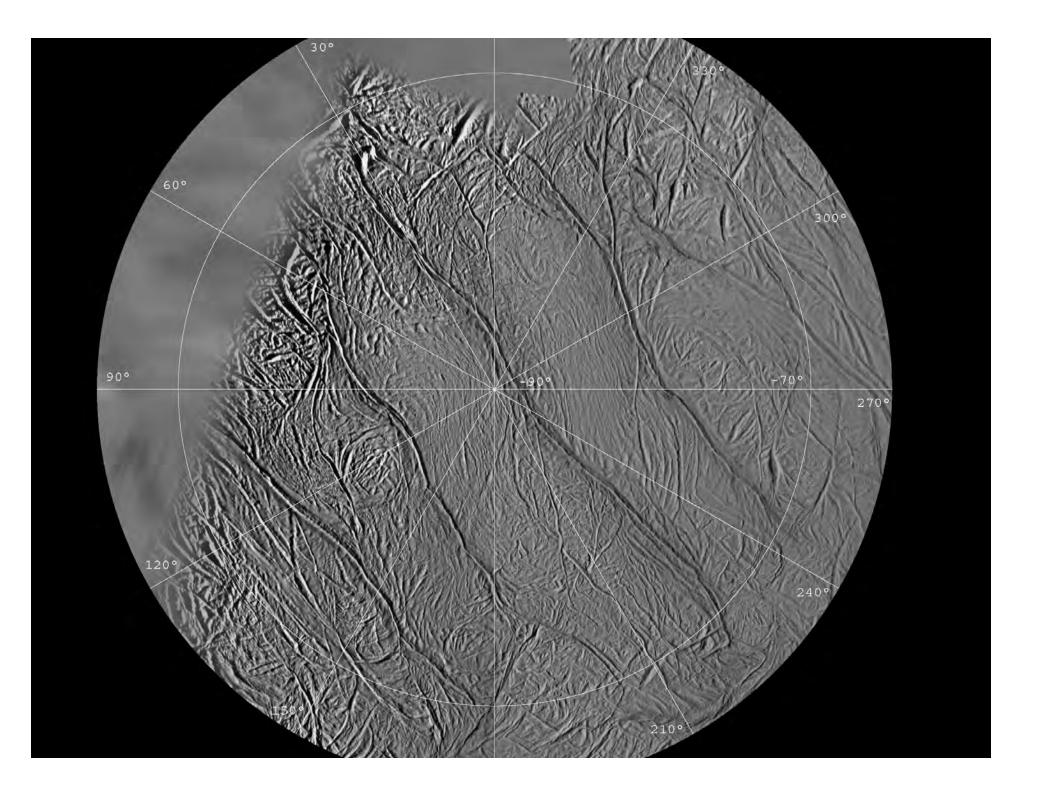


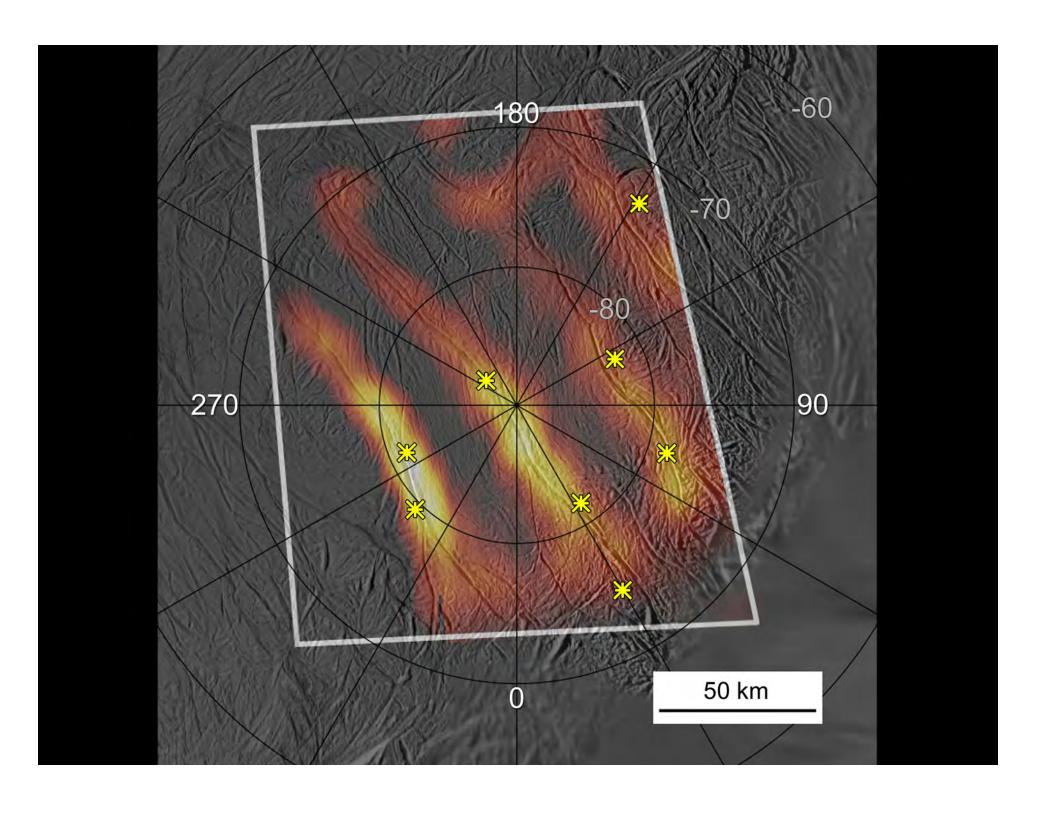


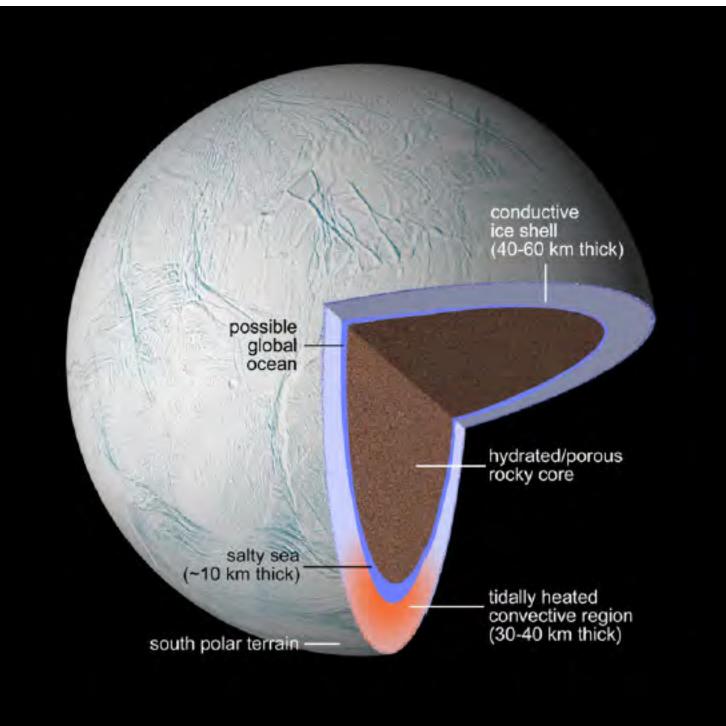


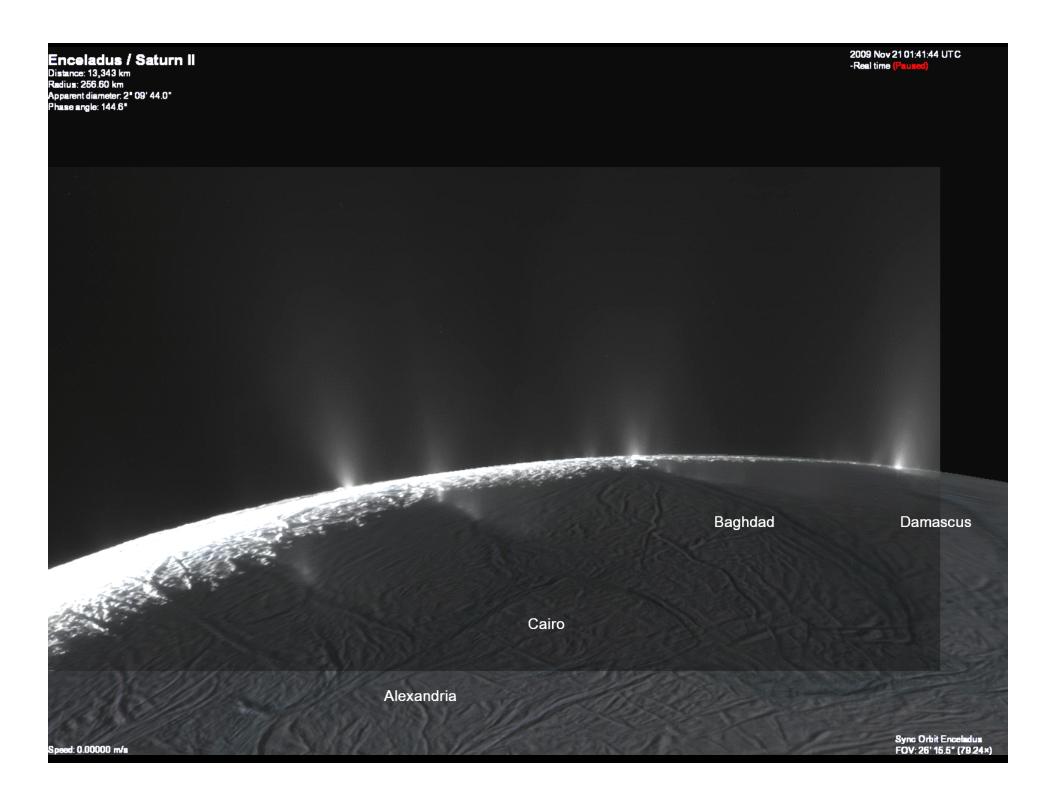
Cassini CIRS

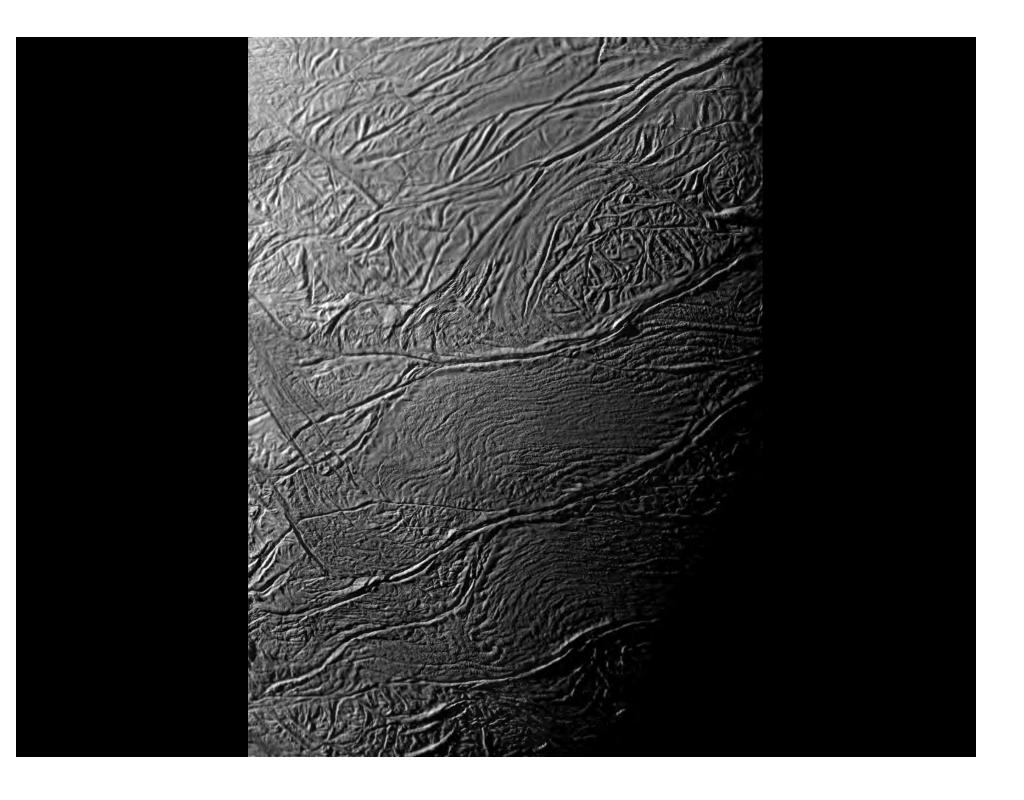


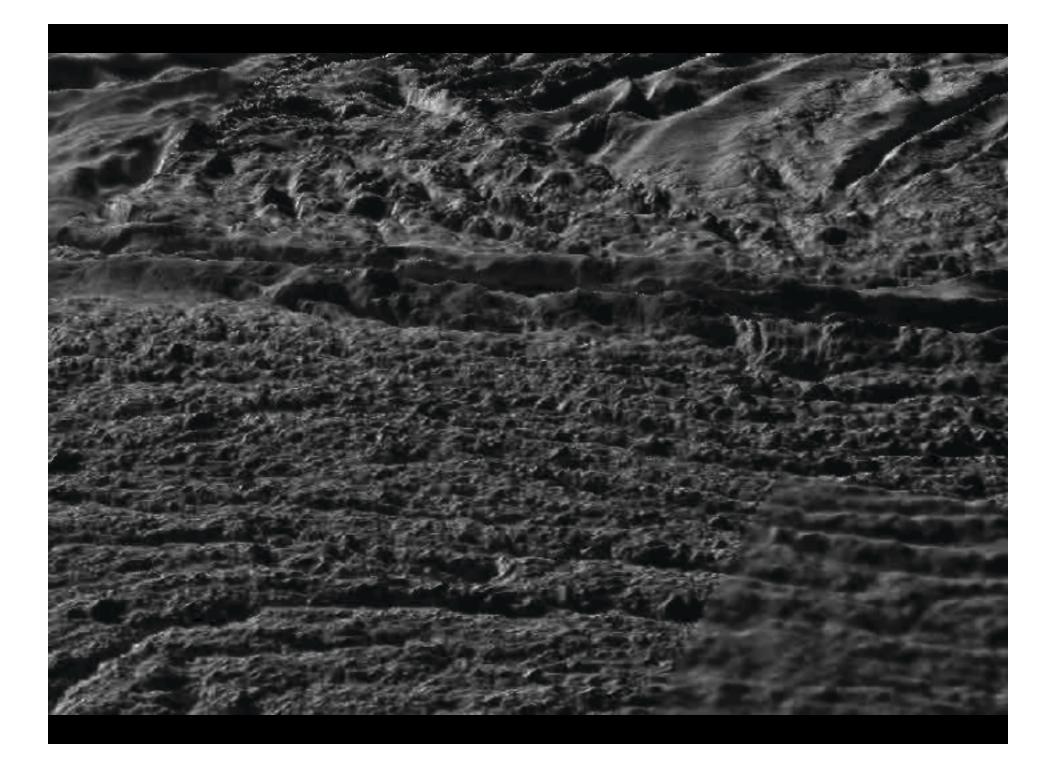


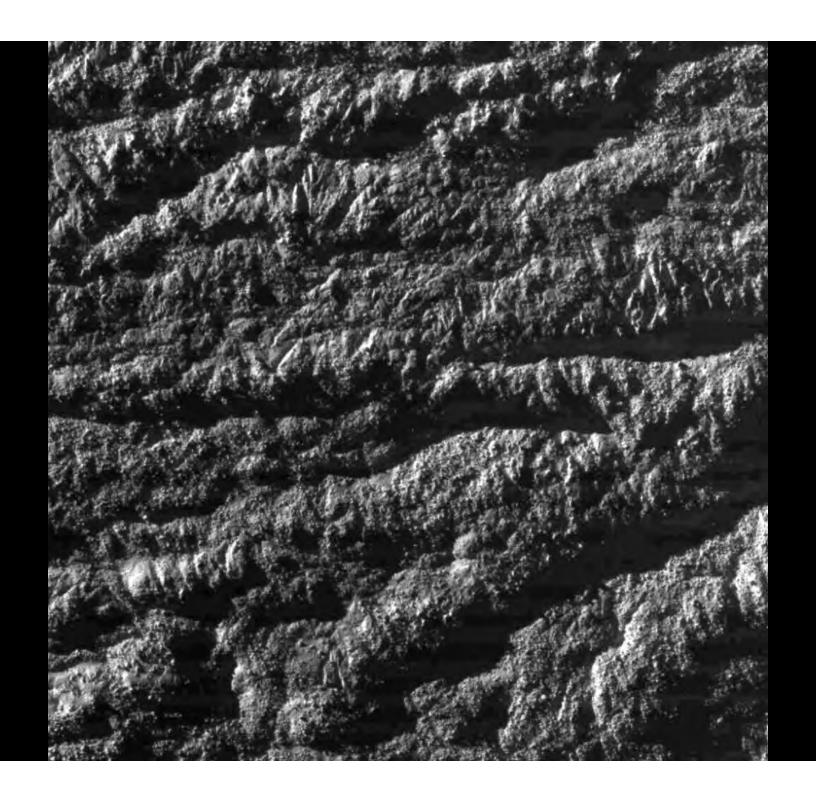


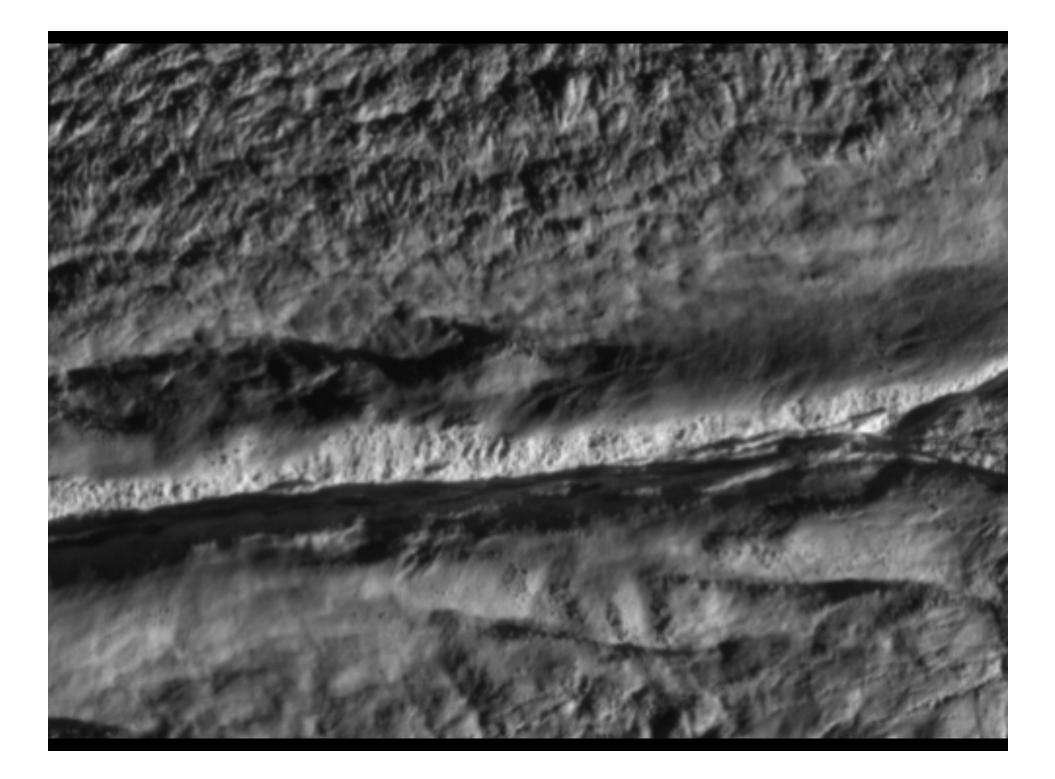


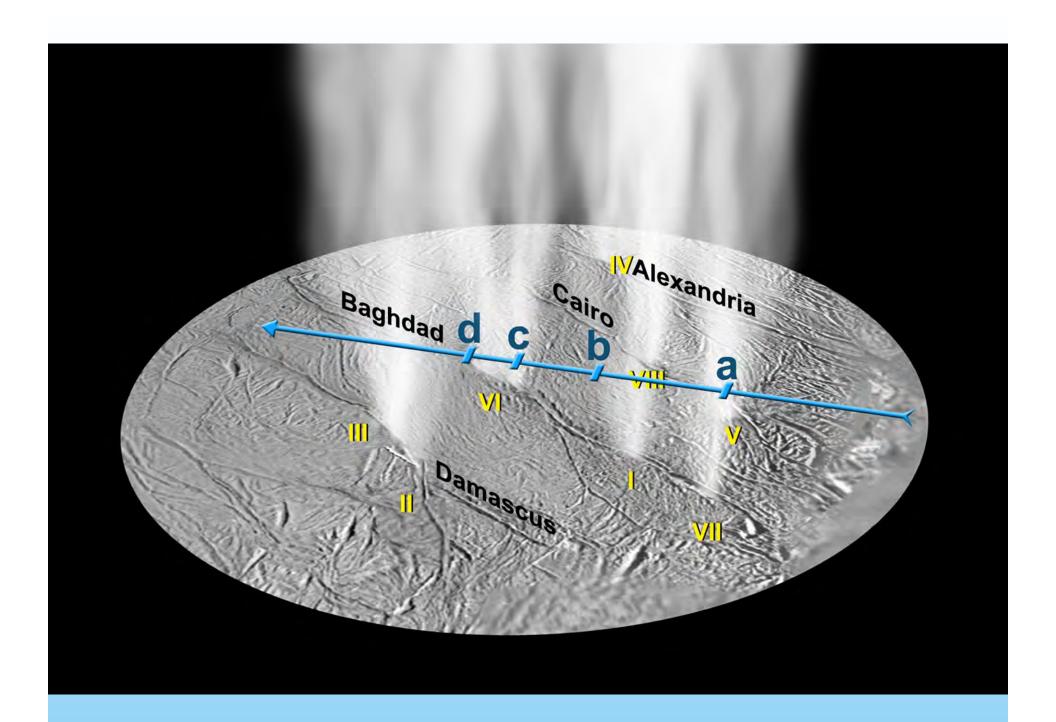


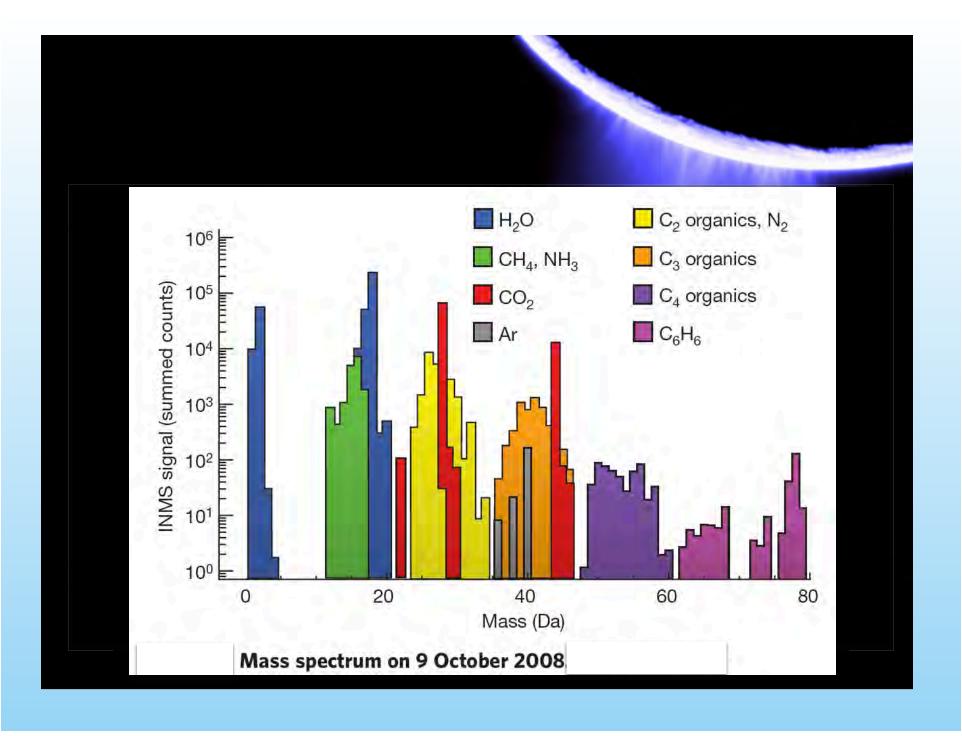


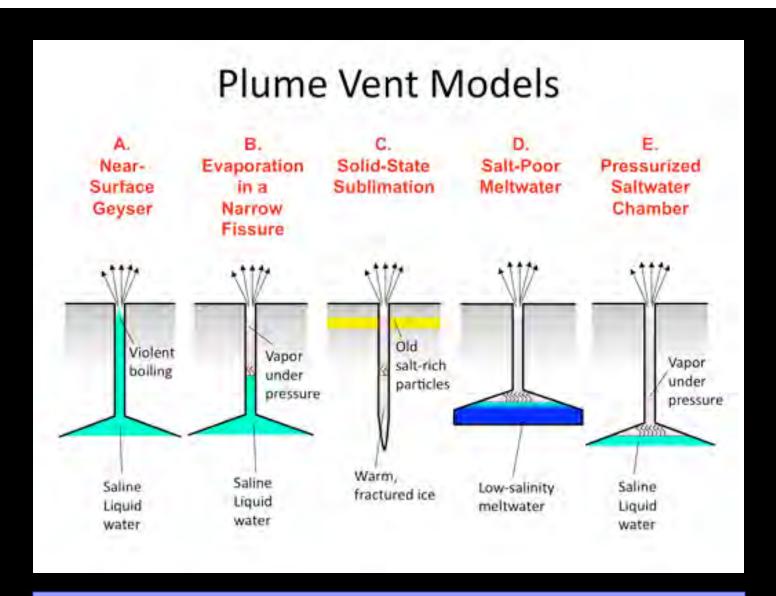






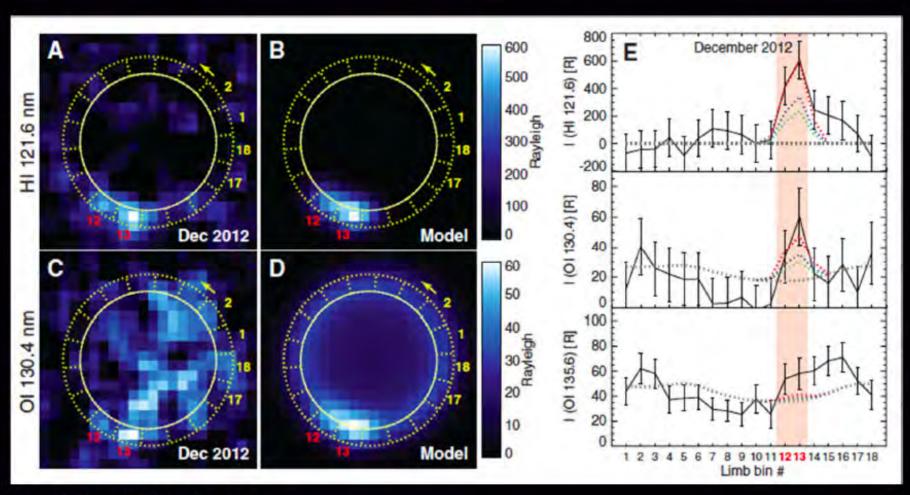




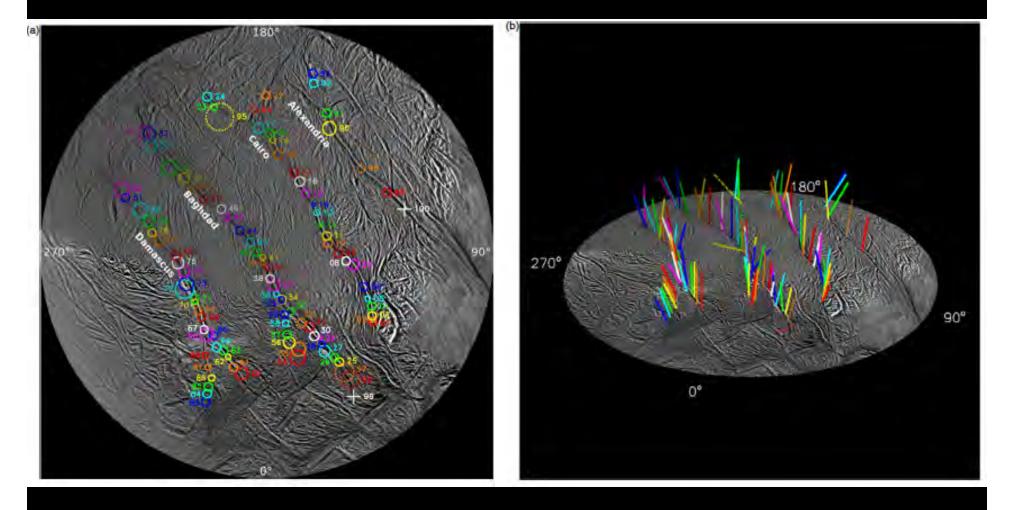


NaCl and bicarbonate in some E ring particles
There must be, or have been, water-rock interaction

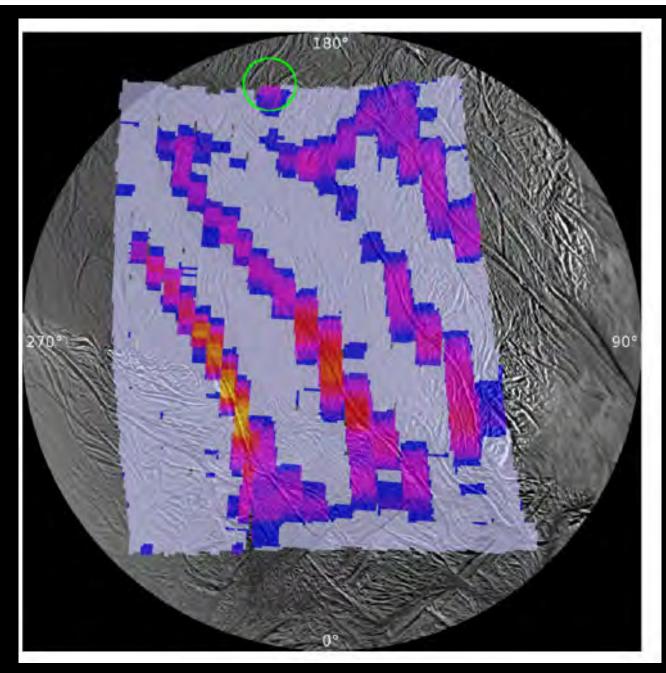
# **EUROPA PLUME?**



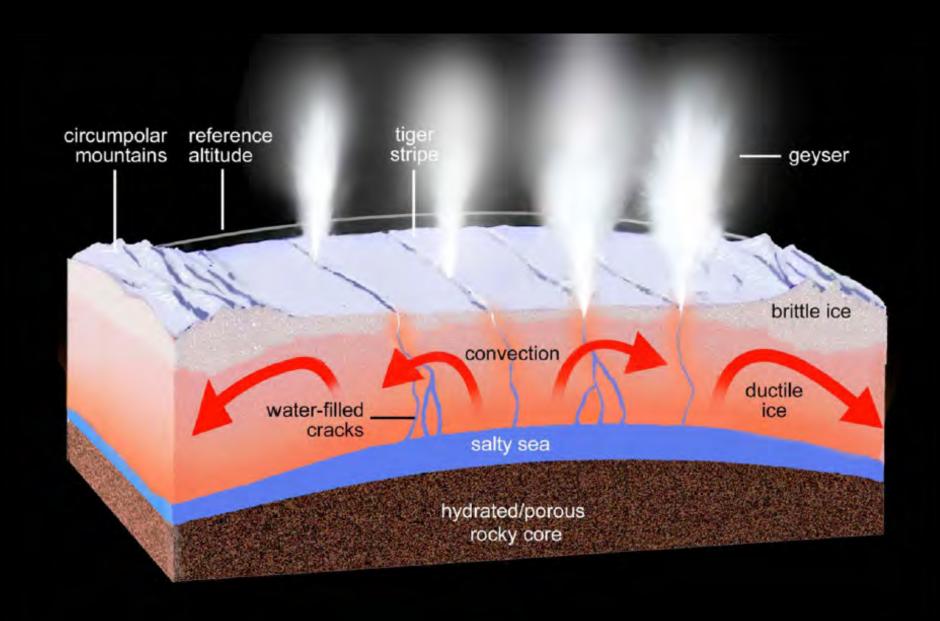
Roth et al 2014, Science



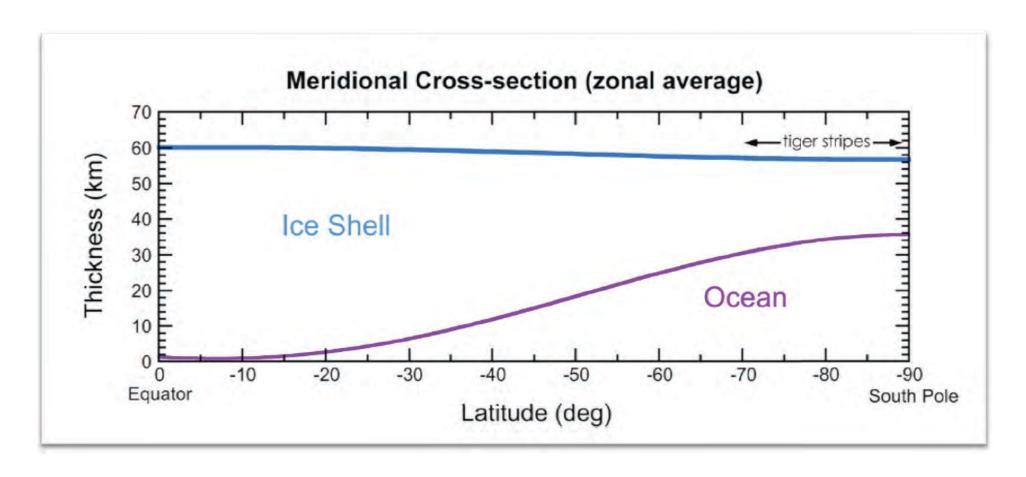
Porco et al. 2014



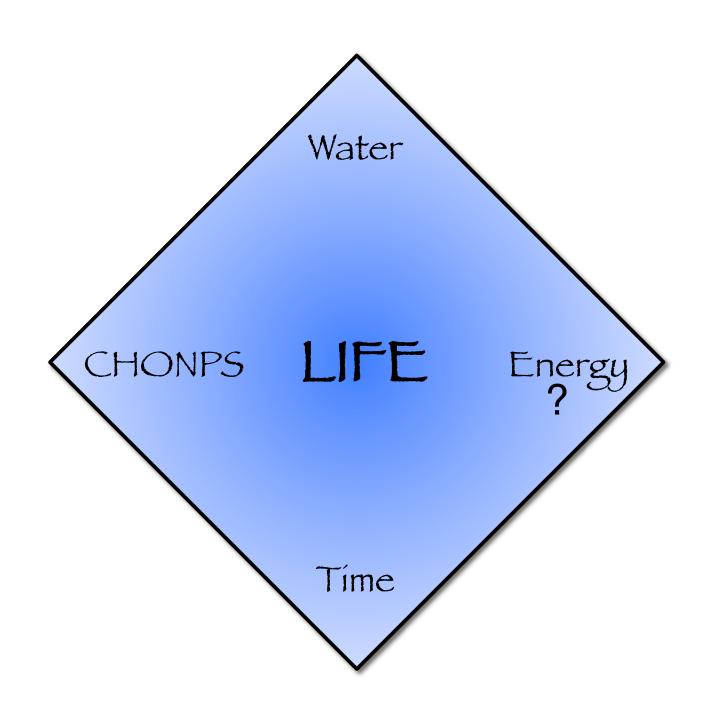
Howett et al. 2011

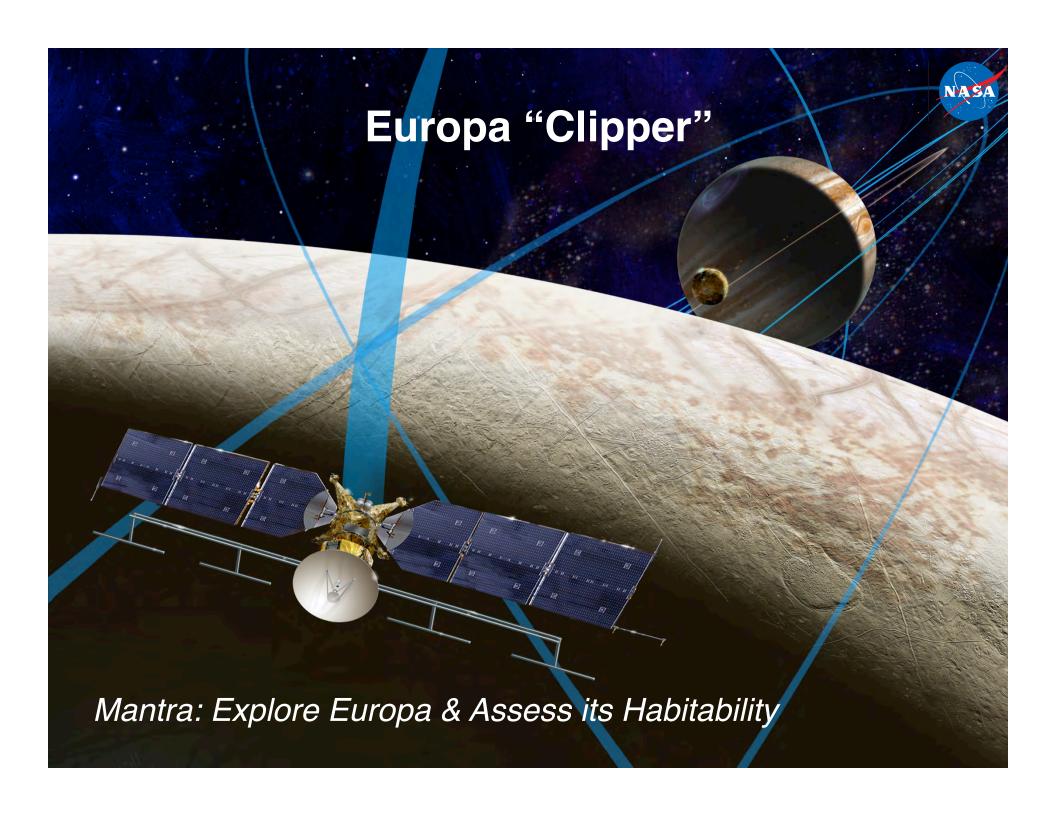


### **Enceladus Gravity and Topography Model** (McKinnon 2014)



A global ocean....now confirmed by libration measurements (Thomas et al. 2015)





# NASA-Selected Europa Instruments

Radiation Science Working Group WG Lead: Chris Paranicas JHU-APL

#### MASPEX

Mass Spectrometer Pl: J. Hunter Waite SwRl, San Antonio

#### SUDA

Dust Analyzer
Pl: Sascha Kempf
Univ. Colorado, Boulder

#### ICEMAG

Magnetometer
Pl: Carol Raymond
JPL-Caltech

#### PIMS

Faraday Cups Pl: Joe Westlake JHU-APL

# Europa-UVS UV Spectrograph Pl: Kurt Retherford SwRl, San Antonio

#### EIS

Narrow-Angle Camera + Wide-Angle Camera Pl: Zibi Turtle JHU-APL

#### MISE

IR Spectrometer
Pl: Diana Blaney
JPL-Caltech

#### **E-THEMIS**

Thermal Imager
Pl: Phil Christensen
Arizona State Univ.

#### **REASON**

Ice-Penetrating Radar PI: Don Blankenship Univ. Texas Inst. Geophys.

Gravity Science Working Group WG Lead: Sean Solomon Lamont-Doherty

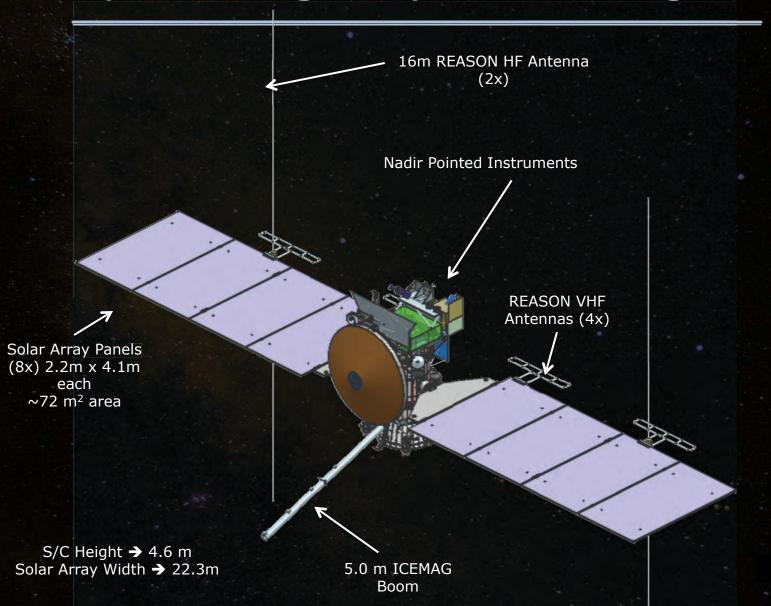


Remote Sensing

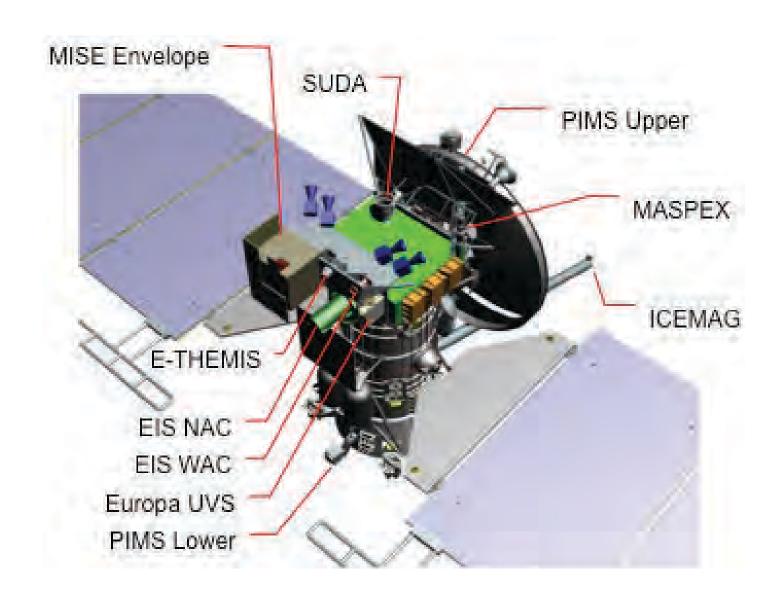


In Situ

# Proposed Flight System Configuration



# **Payload Accommodation**

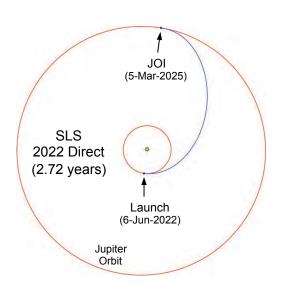


# **Current Jupiter Delivery Strategy**

#### **Baseline**

Launch Vehicle: SLS Block-1 Transfer: Earth-Jupiter Direct Time-of-flight: 2.5-2.7 yrs.





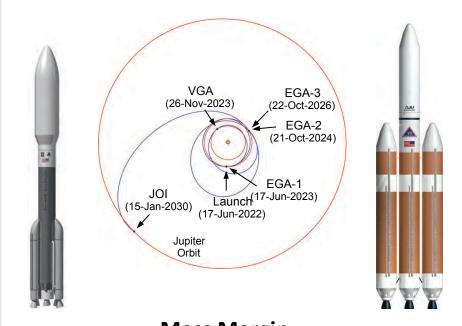
#### **Mass Margin**

35% - 2022 Launch 33% - 2023 Launch

#### **Backup**

Launch Vehicle: Atlas V 551 or Delta IV Heavy

Transfer: EVEEGA
Time-of-flight: 7.4 yrs.



	iviass iviargin	
<b>Atlas V 551</b>		Delta IV Heavy**
29%	2022 Launch	65%
<b>30</b> %	2023 Launch	66%

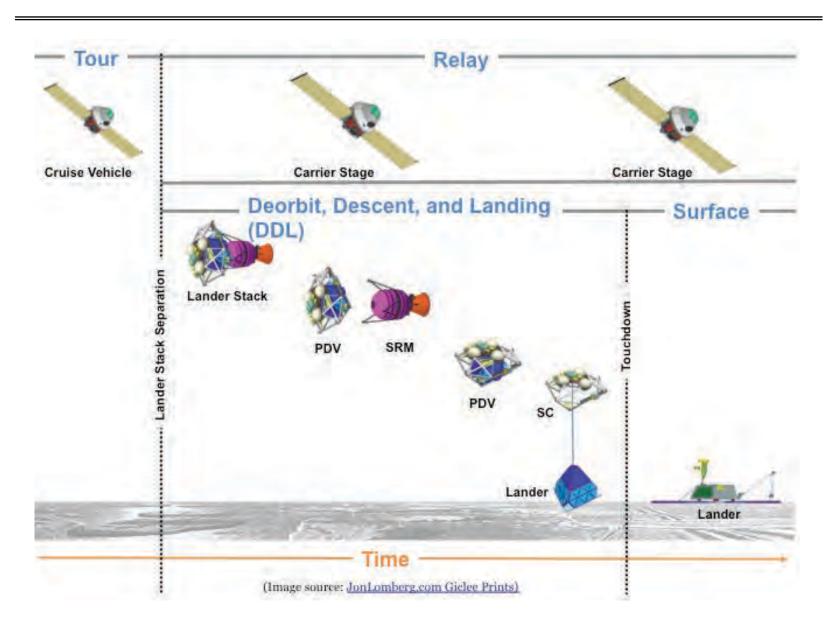
\*\* IF fully utilize L.V. capability

# **Lander Concept**

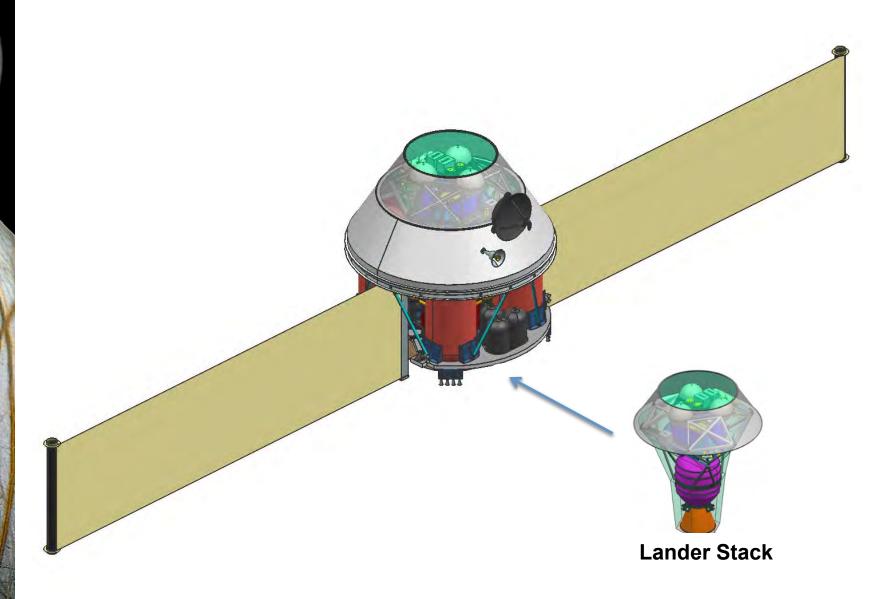
### **Concept Highlights**

- Spacecraft physically decoupled from Clipper
- Enter Jovian system and 'park' in a radiation safe orbit awaiting reconnaissance from Clipper to decide where to target landing
- Spacecraft components:
  - Carrier/Orbit Stage
    - Delivers system to Jovian system and eventually targets lander stack (everything bellow)
    - Provides relay capability (Clipper can be backup) to earth
  - De-orbit Module
    - Decelerates lander to capture a Europa descent trajectory
  - Descent Module
    - Slows down lander, terminal descent to Europa
  - Lander
    - Science!!!

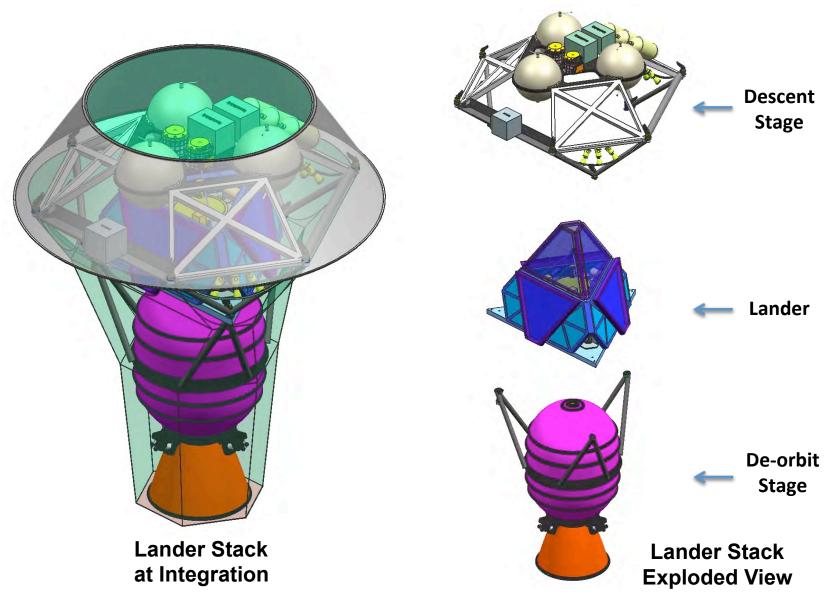
### **Top-Level Mission Event Sequence**



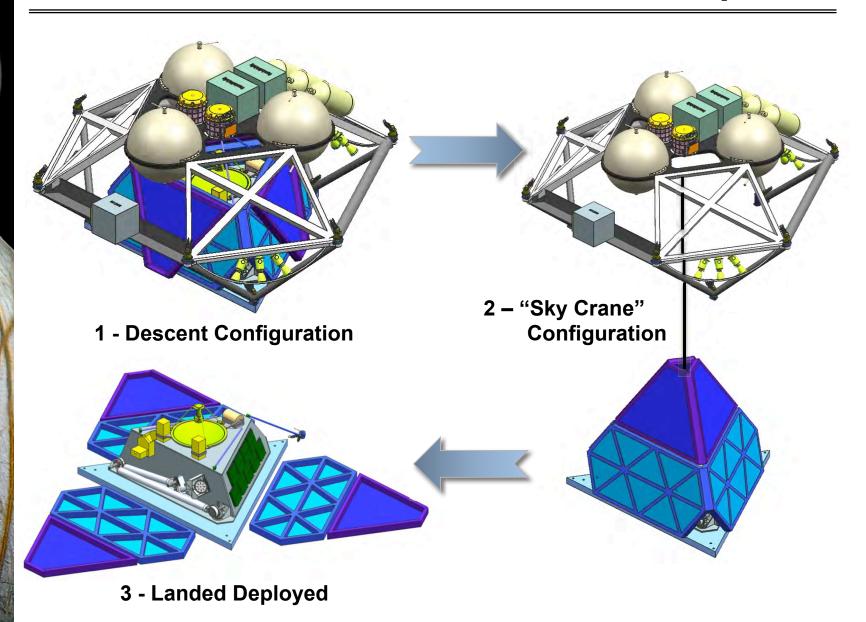
# **Carrier / Orbit Stage Concept**



# **Lander Full System Concept**



### **Lander Descent and Surface Concepts**



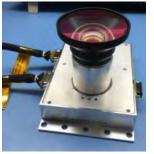
# Model Payload (Total Mass: 25 kg)

- Centerpiece Instruments for Astrobiology
  - GCMS: VCAM GC + Ion Trap MS, 8.3 kg CBE
  - Raman: SHERLOC 5.4 kg CBE
- Auxiliary Instruments
  - Context LanderCams (x2),0.5 kg each CBE
  - Microscopic SampleCam,0.5 kg CBE
- Baseline Instrument (not included in Threshold)
  - 3-axis Geophone, 0.8 kg







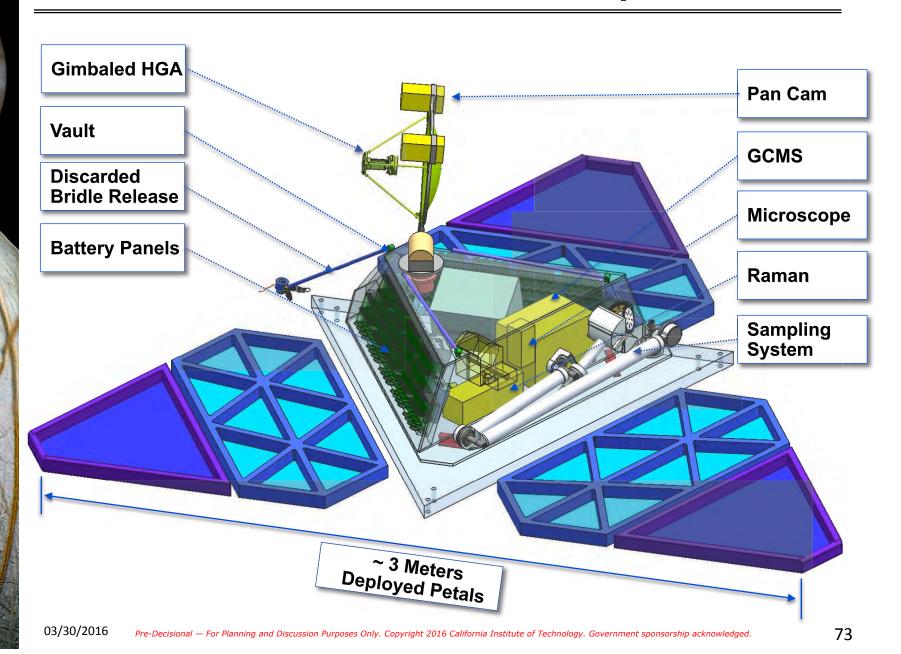








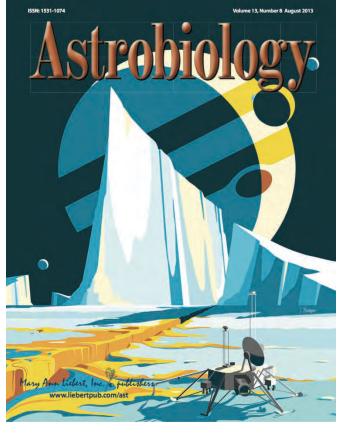
## **Lander Surface Concept**





 2012 Science Definition Team consensus on key questions best answered by a

lander:



Pappalardo et al. 2013

### **NASA Ocean Worlds Program**

NASA is supporting the House recommendation (Commerce, Justice, Science, and Related Agencies Appropriations Bill, 2016) to create an Ocean Worlds program:

"Ocean Worlds Exploration Program.—The recommendation provides \$226,000,000 for Outer Planets, of which not less than \$140,000,000 is for the Jupiter Europa Clipper, or comparable mission, to support the process of finalizing the mission design concept that meets the scientific objectives described in the most recent Planetary Science decadal survey. To support sustained momentum in this program, NASA shall ensure that future funding requests are consistent with achieving a launch no later than 2022, with the goal of launching on a Space Launch System platform as discussed elsewhere in this report.

Many of NASA's most exciting discoveries in recent years have been made during the robotic exploration of the outer planets. The Cassini mission has discovered vast oceans of liquid hydrocarbons on Saturn's moon Titan and a submerged salt water sea on Saturn's moon Enceladus. The Committee directs NASA to create an Ocean World Exploration Program whose primary goal is to discover extant life on another world using a mix of Discovery, New Frontiers and flagship class missions consistent with the recommendations of current and future Planetary Decadal surveys. "



# Backup

